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48	Abstract	<p>We reconstruct past accretion rates of a salt marsh on the island of Sylt, Germany, using measurements of the radioisotopes ^{210}Pb and ^{137}Cs, as well as historical aerial photographs. Results from three cores indicate accretion rates v arying between 1 and 16 mm year⁻¹. Comparisons with tide gauge data show that high accretion rates during the 1980s and 1990s coincide with periods of increased storm activity . We identify a critical inundation height of 18 cm below which the strength of a storm seems to positively influence salt marsh accretion rates and above which the frequency of storms becomes the major factor. In addition to sea level rise, we conclude that in low marsh zones subject to higher inundation levels, mean storm strength is the major factor affecting marsh accretion, whereas in high marsh zones with lower inundation levels, it is storm frequency that impacts marsh accretion.</p>	
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Salt Marsh Accretion and Storm Tide Variation: an Example from a Barrier Island in the North Sea

Mark Schuerch · J. Rapaglia · V. Liebetrau ·
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Abstract We reconstruct past accretion rates of a salt marsh on the island of Sylt, Germany, using measurements of the radioisotopes ^{210}Pb and ^{137}Cs , as well as historical aerial photographs. Results from three cores indicate accretion rates varying between 1 and 16 mm year $^{-1}$. Comparisons with tide gauge data show that high accretion rates during the 1980s and 1990s coincide with periods of increased storm activity. We identify a critical inundation height of 18 cm below which the strength of a storm seems to positively influence salt marsh accretion rates and above which the frequency of storms becomes the major factor. In addition to sea level rise, we conclude that in low marsh zones subject to higher inundation levels, mean storm strength is the major factor affecting marsh accretion, whereas in high marsh zones with lower inundation levels, it is storm frequency that impacts marsh accretion.

Keywords Salt marsh · Accretion rate · Geochronology · Storm activity · Barrier Island · Sylt

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Introduction

Coastal salt marshes in the Wadden Sea (southeastern North Sea) are abundant leeward of the East Frisian and North Frisian barrier islands as well as in front of the dike system on the Dutch–German–Danish North Sea coast. Salt marshes serve as coastal protection structures by reducing the impact of waves on the upper shoreline (e.g., Möller 2006) and as important habitats for specialized plants and animals including migratory and breeding waterfowl and breeding birds (Reise et al. 2010). The existence of salt marshes is critically dependent on how fast sediment accretes relative to sea level rise (Orson et al. 1985; Redfield 1972).

Salt marshes are complex coastal features, governed by various physical and biological processes creating a system that exists in dynamic equilibrium with relative sea level rise (Allen 2000). Salt marsh accretion, defined as the vertical growth of the marsh, occurs when organic and/or inorganic sediments are deposited onto the marsh during inundation (allochthonous growth), as well as when salt marsh plants grow and decompose (autochthonous growth; Dijkema 1987; Kolker et al. 2009). Projected acceleration in SLR, however, may outpace accretion rates in the future, if an insufficient amount of material is deposited on the marsh surface (Kirwan and Temmerman 2009; Orson et al. 1985). Several investigators studied marsh response to mean sea level rise (French 1993; Kirwan and Temmerman 2009; e.g., Orson et al. 1985; Reed 1995); others investigated the influence of tidal range on salt marsh accretion (e.g., Chmura et al. 2001; Harrison and Bloom 1977; Kirwan and Guntenspergen 2010), but relatively little literature is available discussing the impact of non-tidal

61 short-term sea level variations on salt marsh accretion
 62 (Bartholdy et al. 2004; French 2006; Kolker et al. 2009;
 63 Temmerman et al. 2003b). While it is a widespread
 64 assumption that macrotidal environments are more resilient
 65 against SLR than microtidal environments (Kirwan and
 66 Guntenspergen 2010), Allen (2000) and Kolker et al.
 67 (2009) conclude that marsh accretion is dominated by
 68 wind-induced (short-term) sea level variations in microtidal
 69 and by SLR-induced long-term sea level variations in
 70 macrotidal environments. It remains unclear how wind-
 71 induced sea level variations enhance the resilience of the
 72 salt marsh and, in particular, whether many minor storms
 73 are more effective in triggering accretion than a few major
 74 storms. This is noteworthy as certain studies suggest that
 75 the storm activity in the German Bight has increased in
 76 recent years and will most likely increase in the near future
 77 leading to more frequent storm surges and/or higher storm
 78 surge water levels (Beniston et al. 2007; Fischer-Bruns
 79 et al. 2005; Rockel and Woth 2007; von Storch and Weisse
 80 2008; Woth et al. 2006).

81 The goal of this work is to reconstruct the development
 82 of a back-barrier salt marsh on the island of Sylt, Germany,
 83 by using both ²¹⁰Pb and ¹³⁷Cs as geochronometers. ²¹⁰Pb
 84 and ¹³⁷Cs are often used for studying soil and sediment
 85 processes (Armentano and Woodwell 1975; Delaune et al.
 86 1978; Goodbred and Kuehl 1998; He and Walling 1996b;
 87 Kolker et al. 2009). These data will be combined with
 88 information gathered from aerial photographs. We report
 89 how the marsh has been growing during the last decades
 90 and compare accretion rates with historical tide gauge data
 91 in order to infer how storm activity affects marsh accretion
 92 and its resilience against accelerated SLR.

93 **Study Site**

94 The island of Sylt is located in the German Wadden Sea
 95 (54°47'18" N, 008°17'30" E), in the southeastern North Sea
 96 (Fig. 1). Its core is made of Pleistocene protruding outcrops.
 97 About 7,000 years B.P., during a period of SLR deceleration
 98 (Milne et al. 2005), local sea level reached its current
 99 position and sandy spits were formed north and south of the
 100 original Pleistocene outcrop (Ahrendt and Thiede 2001). Sylt
 101 is an elongated barrier island with a total area of 99 km²
 102 extending 40 km from north to south and ranging from 1 to
 103 13 km in width (Kelletat 1992). The western coastline is
 104 characterized by extensive beaches and dunes, and due to a
 105 relatively steep offshore elevation gradient, it is exposed to
 106 highly energetic wave activity from the North Sea (Ahrendt
 107 and Köster 1998). The sheltered eastern coastline is fringed
 108 by extensive tidal flats. Salt marshes are found in the
 109 transition zone from dunes to tidal flats and have a total area
 110 of approximately 5 km² (The Trilateral Monitoring and
 111 Assessment Program—TMAP 2006).

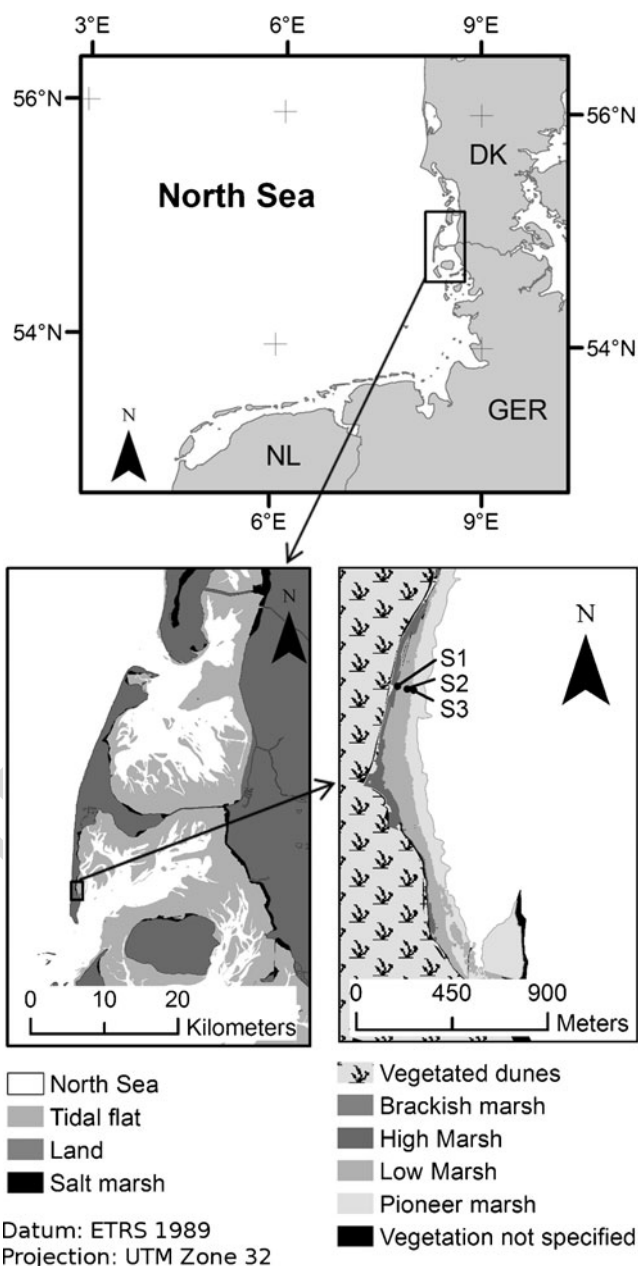


Fig. 1 The study area is located in the southeastern North Sea (cell 1), in the southern part of the island of Sylt (cell 2). The cores were taken on a transect in three dominant vegetation zones, indicated by points (cell 3)

112 The mean tidal range, as of 2010, at the nearby tide
 113 gauge *Hörnnum Hafen* is 2.06 m (WSA—Wasser- und
 114 Schifffahrtsamt Tönning 2007), varying from 1.8 m at neap
 115 tide to 2.3 m at spring tide (BSH—Bundesamt für
 116 Seeschifffahrt und Hydrographie 2008). Tidal range has
 117 constantly increased in the Wadden Sea, including the
 118 *Hörnnumtief*, since 1955 (Jensen and Mudersbach 2004).

119 The immediate study area is located in the southern third
 120 of Sylt between the villages of Rantum and Hörnum and is
 121 characterized by the typical morphology of a barrier island.

122 A transect from the open North Sea to the tidal flats is
 123 marked by a sequence of beaches, dunes, and salt marshes
 124 (Hildebrandt et al. 1993). These dunes developed a dense
 125 grass and scrub vegetation and became stable, as a
 126 consequence of intense planting activities since 1864
 127 (ALW—Amt für Land- und Wasserwirtschaft Husum
 128 1997). The area of the investigated marsh is about
 129 0.3 km², with a length of about 2 km and a width ranging
 130 from 80 to 280 m. The salt marsh itself has a typical
 131 zonation ranging from pioneer marsh vegetation at the
 132 seaward edge to high marsh vegetation close to the foot of
 133 the dunes (Fig. 1). The pioneer marsh is dominated by the
 134 introduced cordgrass *Spartina anglica*, the low marsh zone
 135 is dominated by *Atriplex portulacoides*, and the high marsh
 136 zone is dominated by *Juncus gerardi* and *Elymus athericus*.
 137 The vegetation is not grazed by domestic livestock.

138 The elevation of the salt marsh ranges from about 0.7 to
 139 1.5 m above NN (NN: German reference datum; Fig. 2),
 140 while the mean high water (MHW) level is at about 1 m
 141 (NN; WSA—Wasser- und Schifffahrtsamt Tönning 2007).
 142 The topography of the marsh is rather homogeneous, with a
 143 small marsh cliff (10–20 cm height) at the transition from
 144 pioneer to low marsh (Fig. 2). It is therefore assumed
 145 possible to reconstruct the historic evolution of this salt
 146 marsh by analyzing one transect only.

147 **Methods and Analysis**

148 **Sample Collection and Preparation**

149 Three marsh cores were collected on 22 January 2009 using
 150 PVC tubes with an inner diameter of 11.8 cm. Cores were
 151 collected on a shore-normal transect, within the three dominant
 152 vegetation zones (Figs. 1 and 2) and had a length of about 50–

70 cm. S3, the most seaward core, was collected about 153
 30 m from the seaward edge of the pioneer marsh, 154
 while S2 and S1 were located about 60 m and 110 m 155
 inland from the seaward edge, respectively (Fig. 2). 156
 The large diameter of the tube was chosen in order to reduce 157
 possible compaction of the soil during collection, as well 158
 as to amass a large volume of material for accurate 159
 analysis. Compaction during core collection was measured 160
 in situ as the difference between the measures from the top 161
 of the PVC tube to the salt marsh surface outside and 162
 inside the core. This difference was found to vary between 163
 2% and 4% of the core length (in all three cores) and was 164
 therefore assumed to be negligible. 165

166 The transect where the cores were taken was leveled 166
 by hand using a dumpy level on a tripod. The height of 167
 the (seaward) starting point was determined by compar- 168
 ison of measured flooding and HW times between the 169
 starting point and the tide gauge in *Hörnnum Hafen*. 170
 Elevations of the three core locations, as of today, are 171
 0.94 m (S3), 1.34 m (S2), and 1.44 m (S1) above NN 172
 (Fig. 2). S3 is the only core location that is regularly 173
 inundated, while S1 and S2 are only inundated during 174
 storm tides. 175

176 After extraction, the cores were sliced into layers with a 176
 thickness of 1–5 cm. The thickness of the layers progres- 177
 sively increased toward the bottom of the core. The 178
 sediment was freeze-dried and ground manually in order 179
 to disintegrate the conglomerates of sediment. The ground 180
 material was filled into Petri dishes with a diameter of 181
 52 mm and embedded into epoxy resin, which prevented 182
²²²Rn from degassing. The dry weight of the aliquots 183
 embedded for radiometric analysis ranged from 17.25 to 184
 43.44 g. Samples were allowed to sit for at least 3 weeks in 185
 order to reach equilibrium between ²²⁶Ra and ²¹⁴Bi before 186
 analysis. 187

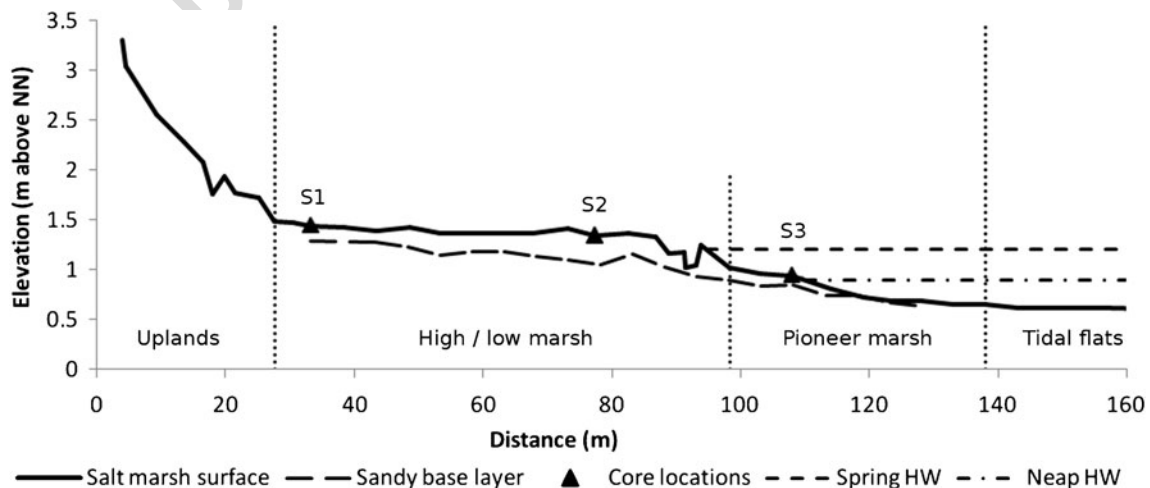


Fig. 2 Topographic profile and marsh zonation of the investigated salt marsh. Core locations and the sandy base layer are indicated by *triangles* and a *dashed line*, respectively

188 Sedimentology: Grain Size, Organic Carbon Content,
189 and Bulk Density

190 Salt marsh sediment was analyzed for grain size,
191 organic carbon content, and bulk density in order to
192 derive the depths of the base layer and for calculation
193 of radioisotope inventories. Grain size measurements
194 were conducted using laser diffractometry (Malvern
195 Instruments Ltd 2010). Although this method is supposed
196 to measure grain sizes ranging from 0.02 to 2,000 μm
197 (Malvern Instruments Ltd 2010), gradually increasing
198 uncertainties in the coarser spectrum of grain sizes were
199 observed during pre-measurements. Therefore, the sedi-
200 ment was sieved at 1,000 μm before the measurements.
201 The preparation of the samples included the destruction of
202 organic carbon, using H_2O_2 , and iron if necessary, using
203 sodium bicarbonate, sodium citrate, and sodium dithionite.
204 The measurements were conducted at obscuration rates
205 between 5% and 30%.

206 The organic carbon content was measured using the
207 element analyzer Euro EA, which performs C/N analysis.
208 The organic carbon is oxidized in the system and escapes as
209 CO_2 into a gas chromatographer where total C (and N)
210 content is determined.

211 The bulk density (ρ) was indirectly calculated from the
212 organic carbon and the water content (O_C and W) in the
213 samples as not enough sample material was available in the
214 upper layers for direct measurement. Assuming a constant
215 density for water ($\rho_W=1.02 \text{ g cm}^{-3}$) and for mineral and
216 organic matter ($\rho_M=2.6 \text{ g cm}^{-3}$, $\rho_O=1.2 \text{ g cm}^{-3}$, respectively),
217 the bulk density was calculated using the following formula
218 (Kolker et al. 2009):

$$\rho = \frac{1 - W}{\frac{W}{\rho_W} + \frac{1 - W}{(1 - O_C \times \rho_M) + (O_C \times \rho_O)}} \quad (5)$$

219 Additional measurements on dry bulk density and
220 relative water content were conducted on cores
221 extracted on the 18 May 2011 in order to validate
222 these bulk density calculations. It was shown that the
223 results of the model fit the measured data ($R^2=0.98$,
224 $p<0.01$).
225

226 Assessment of Autocompaction

227 Bulk density calculations were used for the assessment of
228 autocompaction within the cores S1, S2, and S3. Bulk
229 densities within the uppermost 5 cm of the salt marsh cores
230 were investigated for autocompaction (Bartholdy et al.
231 2010), since variability of bulk density triggered by
232 changes of grain size and organic carbon content in the
233 lower parts of the core was too large to derive reliable
234 estimates for autocompaction. A logarithmic curve (Eq. 6)

was fitted through the calculated bulk density data (BDD) 235
within the uppermost 5 cm. 236

$$\text{BDD} = A \times \ln(z) + B \quad (6)$$

where A and B are constants to be optimized and z is the 238
depth of the respective salt marsh layer. 239

If the model was found to significantly represent the 240
bulk density data ($p<0.05$), autocompaction was included 241
into the calculation of accretion rates by decompacting the 242
core. Decompaction was performed by calculating the 243
thickness of a layer in depth z (T_z) to the time when it 244
was deposited using the inverse version of the equation 245
presented by Williams (2003): 246

$$T_z = \frac{\text{BDD}_z}{\text{BDD}_0} \times T_0 \quad (7)$$

where BDD_0 and T_0 are the bulk dry density and the 248
thickness of today's surface layer, respectively. 249

Radiometric Measurements 250

^{137}Cs ($t_{1/2}=30.2$ years) is a product of nuclear fission and is 251
highly particle reactive (He and Walling 1996b). In Europe, 252
it is marked by two main periods of atmospheric fallout 253
(1963 and 1986) producing two useful marker horizons 254
(Delaune et al. 1978; Kirchner and Ehlers 1998; Pedersen et 255
al. 2007). The ^{137}Cs peak around 1963 is produced by a 256
series of tests of nuclear bombs during the 1950s and 257
1960s, while the 1986 peak is a result of the nuclear 258
disaster in Chernobyl. Some authors have found an 259
additional ^{137}Cs signal in cores taken along the North and 260
the Baltic Sea, peaking between 1974 and 1977 (Andersen 261
et al. 2000; Kunzendorf et al. 1998). However, given the 262
low initial activity and the 30-year half-life of ^{137}Cs , these 263
activities are probably below detection limits today. ^{210}Pb 264
($t_{1/2}=22.3$ years), on the other hand, is constantly being 265
supplied from the atmosphere via the decay of its 266
grandparent nuclide ^{222}Rn (Koide et al. 1972; Walling 267
et al. 2003). This continuous deposition makes ^{210}Pb an 268
excellent geochronometer for deriving accretion rates in 269
coastal salt marshes (Appleby and Oldfield 1978, 1983; 270
Bartholdy et al. 2004; Bellucci et al. 2007; Kirchner and 271
Ehlers 1998). 272

A low-background coaxial Ge(Li)detector was employed 273
to measure total gamma-ray activity of ^{210}Pb , ^{226}Ra , and 274
 ^{137}Cs . It is a non-destructive counting method that allows 275
simultaneous measurement of all three radionuclides 276
(Nikulina 2008). Measurement of ^{226}Ra is necessary 277
because it is a proxy for supported ^{210}Pb . The analysis 278
was conducted by the "Laboratory for Radioisotopes" in 279
Goettingen, Germany. ^{210}Pb was measured via its gamma 280
peak at 46.6 keV, ^{137}Cs via its peak at 661.7 keV, and ^{226}Ra 281
via the peaks of its granddaughters ^{214}Pb and ^{214}Bi at 352 282

283	and 609.3/1,120.3 keV, respectively. The measurement time	Hydrological Data	330
284	for all samples was 250,000 s.		
285	Errors/Detection Limit		
286	Errors generally increase toward the low energy gamma-ray	Annual data about mean sea level (MSL) were calculated	331
287	spectrum, since the background radiation is elevated as a	from annual MHW and mean low water (MLW) levels	332
288	consequence of the Compton effect of higher energetic	(Hofstede 2010, personal communication). Storm frequen-	333
289	radionuclides (Compton 1923). Meanwhile, for higher	cy and storm intensity were aggregated from the high water	334
290	energy spectra, the background radiation decreases contin-	levels, collected for every tidal cycle between 1938 and	335
291	uously. For ²¹⁰ Pb that is measured at 46.6 keV (very low	2007 (Wahl 2010, personal communication; Wahl et al.	336
292	energy), this is of particular importance, since the back-	2011). Eustatic sea level rise was included into the analysis	337Q2
293	ground radiation within the sample may be elevated above	as an underlying trend that was removed from the original	338
294	the ²¹⁰ Pb activity of the soil sample, resulting in measure-	dataset before storm parameterization was conducted.	339
295	ments below detection limit.	Global SLR data, presented by Church and Church and	340
296	Another source of error for radiometric measurements is	White (2011), were used for that purpose due to a lack of	341
297	the background radiation of the environment, which is	reliable estimates of local eustatic SLR.	342
298	quantified by measuring a parallel blank sample and	Storm activity was parameterized by storm frequency,	343
299	minimized by increasing the sample volume as well as the	defined as the number of tides above a certain storm level,	344
300	counting time. The combined “internal” and environmental	and by storm intensity, defined as the mean height of these	345
301	background radiation at 46.6 keV is therefore higher than	storm tides. For storm parameterization, we used these	346
302	the activity measured for ²²⁶ Ra in higher energy spectra,	definitions in order to ensure that storm intensity is	347
303	and detection limits can be included into the analysis as	statistically independent from storm frequency. Mean sea	348
304	maximum values for ²¹⁰ Pb activities.	level in turn was calculated as a function of MHW and	349
305	Dating Model	MLW (Wahl et al. 2010).	350
306	The constant rate of supply model (CRS) is employed for	Mean values of storm intensity and frequency data were	351
307	the dating of sediment characterized by variations of initial	aggregated for all years in which an estimate for accretion	352
308	concentrations (Appleby and Oldfield 1978; Appleby and	rate was available, using a moving average filter. Taking	353
309	Oldfield 1983). It is therefore useful in study sites with	into account an expected error and temporal resolution of	354
310	irregular inundation, where accretion rates strongly vary	the dating of the sediment layers, the window size was	355
311	with time. This model allows for the calculation of the age	chosen to be 5 years.	356
312	of the sediment by using radioisotope inventories (Appleby	Analysis of Storm-Related Salt Marsh Accretion	357
313	and Oldfield 1978; Kolker et al. 2009).		
314	Due to the relative elevation of the core locations within	One of the main objectives of this study is to investigate the	358
315	the tidal frame (Fig. 2), irregular inundation occurs during	influence of storm frequency and storm intensity on the	359
316	storm surges, negating the assumption that accretion rates	observed accretion rates. Multiple linear regression analysis	360
317	are constant over time. Therefore, the CRS model was used	was employed to assess how much of the variation within	361
318	for this study.	the accretion rate time series can be explained by one of	362
319	Aerial Photographs	these parameters. This analysis was carried out by including	363
320	Aerial photographs of the investigated area, provided by the	all high water levels that flooded the marsh from 1938 to	364
321	Landesbetrieb für Küstenschutz, Nationalpark und Meer-	2007. By gradually increasing the storm level and therefore	365
322	eschutz Schleswig-Holstein (LKN-SH), were utilized in	excluding the low inundation events, we investigated how	366
323	order to analyze the historic extension of the salt marsh	different high water levels were influencing accretion rates	367
324	during the last century and to verify the radiometric	in the past. The standardized coefficients (β) of the multiple	368
325	measurements. Photographs were available from 1937,	linear regressions, as well as the p values, were compared	369
326	1958, 1988, 1999, 2003, and 2007. After georeferencing	for the different storm levels.	370
327	the pictures and overlaying the core locations, visual	Results	371
328	interpretation was performed in order to classify the core		
329	location into the categories “marsh” and “no marsh.”	Grain Size Analysis and Organic Carbon Content, and Bulk	372
		Density	373
		In each of the cores, a fine-grained upper layer is	374
		underlain by layers comprised of mainly coarse-grained	375

376 sediments. The transition zone in core S1 is very
 377 distinct, and fine-grained sediments are not present
 378 below a depth of about 10 cm (Fig. 3). In contrast, the
 379 transition zones in cores S2 and S3 are more gradual.
 380 Fine-grained sediment is, in fact, present throughout the
 381 entire core S2 (Fig. 3).

382 Bulk densities reproduce this pattern of fine-grained
 383 sediments on top of a coarse-grained base layer by
 384 increasing at the interfaces of salt marsh sediments and
 385 the sandy base layer (Fig. 3). In core S1, a clear increase of
 386 bulk density is observed at a depth of about 10 cm,
 387 reaching 1,574 kg m⁻³. The high bulk density of
 388 1,469 kg m⁻³ at the depth of 4.5 cm is considered as an
 389 outlier and was therefore excluded from the analysis.
 390 Instead, the average from the neighboring upper and lower

layers was employed. The increase in S2 is more gradual
 and constant, with a maximum bulk density of 939 kg dm⁻³
 at 27 cm. In core S3, a considerable increase in bulk density
 is observed at a depth of about 5 cm, with a maximum
 value of 1591 kg m⁻³. Within the silty salt marsh
 sediments, bulk densities in cores S2 and S3 do not
 significantly increase, mostly varying between 500 and
 600 kg m⁻³. In core S1, a significant increase from 326 to
 656 kg m⁻³ within the uppermost 6 cm is observed, and
 bulk density then decreases to 450–500 kg m⁻³ in the lower
 silty layers.

Salt marsh sediments are characterized by fine-grained
 material with high organic content, which allows for visual
 and experimental determination of the thickness of the
 marsh layer deposited on top of the sandy base. Herein, the
 base of the marsh can clearly be identified in cores S1 and
 S3 at 8.5 and 4.5 cm, respectively, whereas it is not as clear
 in core S2 (Fig. 3). The fine-grained fraction rapidly
 decreases at 21 cm but increases at 23 and 27 cm again.
 Although high fractions of sand at depths of 21 and 25 cm
 are found, marsh development seems to have started at a
 depth of about 27 cm.

The increase of bulk density within the uppermost 5 cm
 of core S1 was found to fit the logarithmic model
 (equation 6; R²=0.97, p<0.02), which is considered as
 evidence for autocompaction. Model parameters A and B
 were estimated to be 0.0879 and 0.3791, respectively. The
 autocompaction rate during the growth of the salt marsh
 therefore averages 1 mm year⁻¹.

Aerial Photographs

The first photograph available, taken in 1937 (Fig. 4a),
 provides strong evidence for the presence of marsh at
 locations S1 and S2; however, no marsh is yet present at
 location S3. A sand bar seems to have formed between the
 seaward edge of the marsh and the adjacent tidal flat. At
 the landward edge of the marsh sand appears to be
 transported onto the marsh surface. However, in the
 following photograph from 1958 (Fig. 4b), sand transport
 onto the marsh at the landward edge seems to be less
 visible than in 1937; meanwhile, the sand bar at the
 seaward edge has developed further offshore. Recent
 photographs from 1988 and 1995 (Fig. 4c, d) do not show
 considerable changes of the marsh extension toward the
 land, but do indicate a seaward expansion. In 1988
 (Fig. 4c), core location S3 is found to be in a channel
 between the developing pioneer marsh and the mature
 main part of the marsh. The photo from 1995 (Fig. 4d)
 shows that the channel appears to be closed and covered
 by pioneer marsh vegetation. Marsh development at core
 location S3 is, therefore, estimated to have started between
 1988 and 1995.

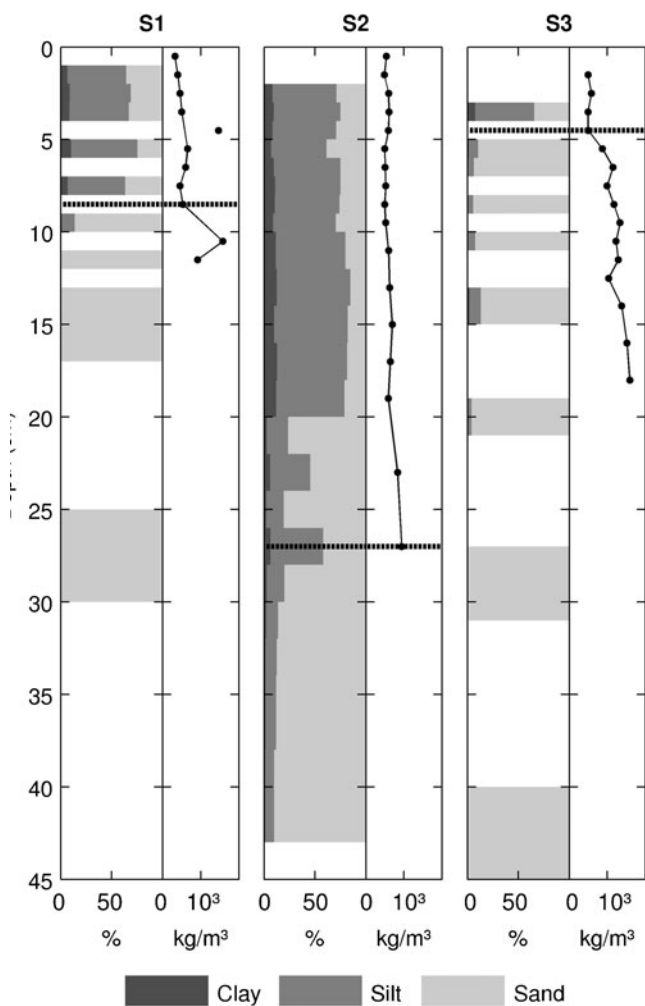


Fig. 3 Grain size composition (in percent) and bulk density calculations (in kilograms per cubic meter), as described in Eq. 5, for all three cores and all depths (in centimeters) that were measured. Grain size composition is depicted in the bar graph and bulk density by the line graph. The depth of the sandy base layer is indicated by the dashes line

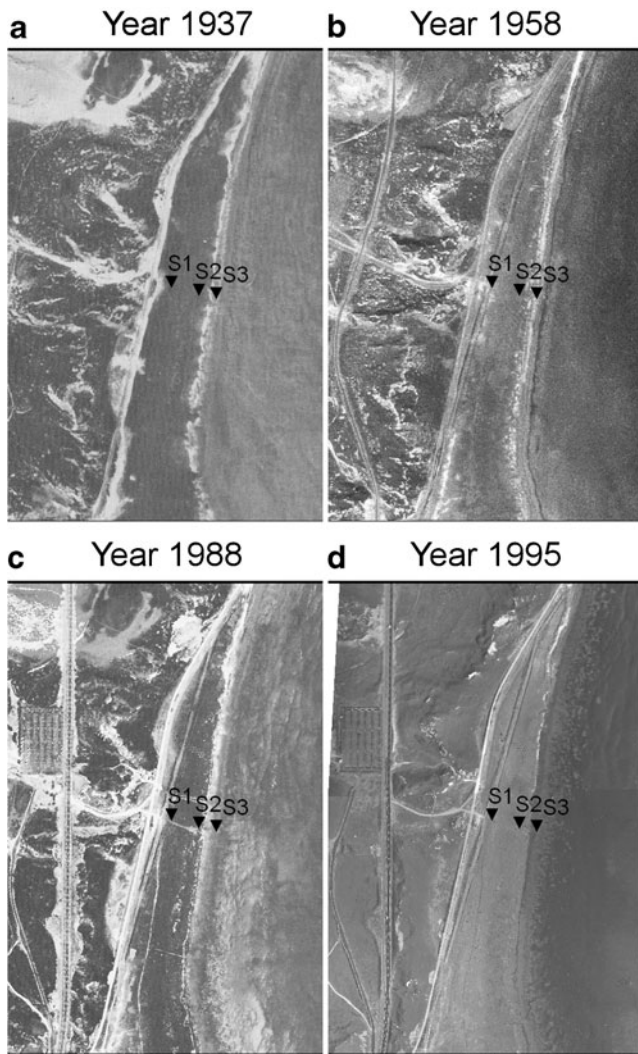


Fig 4 Aerial photographs of the investigated salt marsh from 1937 (a), 1958 (b), 1988 (c), and 1995 (d). Dark colors indicate vegetated areas. Bright colors indicate sand. Source: LKN-SH

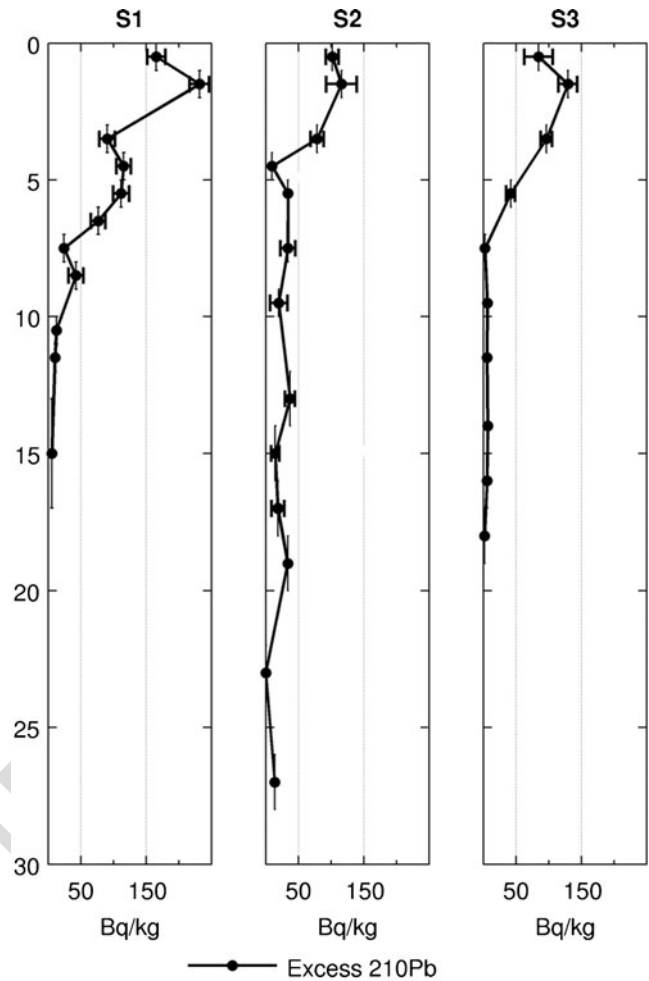


Fig. 5 Excess ^{210}Pb (in Becquerel per kilogram) in all three cores and all depths (in centimeters) that were measured. Horizontal error bars indicate propagated errors for excess ^{210}Pb . Measurements without horizontal error bars were measured below detection limit. Vertical error bars represent the thickness of the measured layer

442 Radioisotope Data

443 ^{210}Pb Data

444 All three depth profiles of excess ^{210}Pb ($^{210}\text{Pb}_{\text{XS}}$) generally
 445 decrease as expected (Fig. 5). Large variations are found in
 446 the surface concentrations, with the highest activity
 447 (232 Bq kg^{-1} in 1.5 cm) in core S1. Distinct peaks are
 448 found at a depth of 1.5 cm in all three cores.

449 In core S1, a slow decrease is observed, due to relatively
 450 high activities ($116, 112 \text{ Bq kg}^{-1}$) at depths of 4.5 and
 451 5.5 cm. Activities below 10 cm are all less than the
 452 detection limit and therefore less than 12.6 Bq kg^{-1} (Fig. 5).
 453 The initial decrease in S2 is much greater and influenced by
 454 low activity at 4.5 cm (below detection limit of
 455 9.5 Bq kg^{-1}). Activities below 4.5 cm decrease slower,
 456 eventually resulting in activities below the detection limit

for all depths lower than 17 cm (Fig. 5). In core S3, $^{210}\text{Pb}_{\text{XS}}$ 457
 also quickly decreases within the upper 7 cm from 458
 129 Bq kg^{-1} at 1.5 cm to less than 2.6 Bq kg^{-1} at 7.5 cm. 459
 $^{210}\text{Pb}_{\text{XS}}$ activities at lower depths are all below the 460
 detection limit (Fig. 5). 461

As ^{210}Pb activities in the sandy substrate of the cores S1 462
 and S3 are all below the detection limit, normalization for 463
 grain size and organic carbon content of these activities is 464
 impossible. In core S2, where the fraction of fine-grained, 465
 highly organic sediment is rather constant within salt marsh 466
 layers, normalization is not necessary. 467

468 ^{137}Cs Data

^{137}Cs peaks are found in cores S1 and S2. In core S3, no 469
 ^{137}Cs peak was found. S1 is characterized by a major peak 470
 at a depth of 5.5 cm (267 Bq kg^{-1}) and a minor peak at 471
 8.5 cm (38 Bq kg^{-1} ; Fig. 6). ^{137}Cs appears to decrease 472

473 toward zero at about 15 cm, although it should be
 474 considered that the sediment below 10 cm is very coarse
 475 and the effect of coarse-grained sediment might reduce the
 476 ^{137}Cs activity (He and Walling 1996a). In core S2, the
 477 situation is clearer: Two peaks are found in depths of
 478 5.5 cm (64 Bq kg^{-1}) and 17 cm (52 Bq kg^{-1}); we note that
 479 the upper peak's activity is significantly lower than the
 480 activity in S1 (Fig. 6). The ^{137}Cs activity in core S3 is
 481 missing a distinct peak and has a very low activity
 482 ($<1 \text{ Bq kg}^{-1}$) in the sandy base layer. Finer sediments in
 483 the upper layers are marked by a slight increase in activity
 484 up to 8 Bq kg^{-1} (Fig. 6).

485 *Age of Base Layer*

486 The age of the marsh and the mean accretion rates were
 487 determined by radioisotope dating (^{210}Pb) and compared
 488 with aerial photographs (Table 1). The base layer at core

Table 1 Ages of base layers and mean sedimentation rates: comparison of ^{210}Pb datings with information from aerial photographs

Core name	Depth of base layer	Age of base layer (^{210}Pb)	Age of base layer (aerial photographs)	Sedimentation rate (mm/year)
S1	8.5 cm	1925–1955	<1937	1–1.2
S2	27	1915	<1937	2.8
S3	4.5 cm	No measurement	1988–1995	2.5 (since 1996)

location S1 is found at a depth of about 8.5 cm (Fig. 2 and 3). Aerial photographs show that the marsh was present in 1937. However, analysis of ^{210}Pb data does not support this finding, suggesting the marsh to be younger. Several ^{210}Pb data in core S1 and S3 are below the detection limit due to a relatively high baseline activity in these samples. Since the detection limit is considered as a maximum activity and allows for activities within the range of zero to the detection limit, the original error range needs to be extended. Therefore, the year when the base layer was deposited (according to ^{210}Pb dating) is calculated to be between 1925 (± 5 years) and 1955 (± 5 years) (Table 1). Utilizing these dates and the aerial photographic evidence, we conclude that the marsh started to develop between 1925 (± 5 years) and 1937, resulting in a mean accretion rate between 1 and 1.2 mm year^{-1} . In core S2, the base layer is found at a depth of 27 cm (Figs. 2 and 3). ^{210}Pb dating suggests that this layer was deposited in 1915 (± 5 years). The aerial photograph of 1937 confirms the existence of the marsh at that time. Therefore, the mean accretion rate for core location S2 is estimated to be 2.8 mm year^{-1} (Table 1). S3 has a marsh layer that is only 4.5 cm thick (Figs. 2 and 3), but no ^{210}Pb measurement is available for the base layer at that depth. Meanwhile, the sample at 3.5 cm is calculated to be deposited in 1996, suggesting a mean accretion rate of 2.5 mm year^{-1} . Aerial photographs indicate the beginning of the pioneer (*Spartina*) marsh development between 1988 and 1995.

Accretion Rates

A reliable time series of accretion rates could only be drawn for the core S2 because the marsh layer in S3 is too thin to show significant variations, while ^{210}Pb dating for S1 resulted in small accretion rates associated with large uncertainties. However, for S1, a general trend toward an increase of accretion rates can be observed from the beginning of the 1960s, resulting in a recent accretion rate between 3 and 3.5 mm year^{-1} . Accretion rates for S2 indicate strong variations over the last 75 years, ranging from about 1 up to 16 mm year^{-1} (Fig. 7). No clear trend can be observed, although accretion rates in the 1980s and 1990s seem to be higher. Two distinct peaks are found in 1982 and 1992 (according to ^{210}Pb dating).

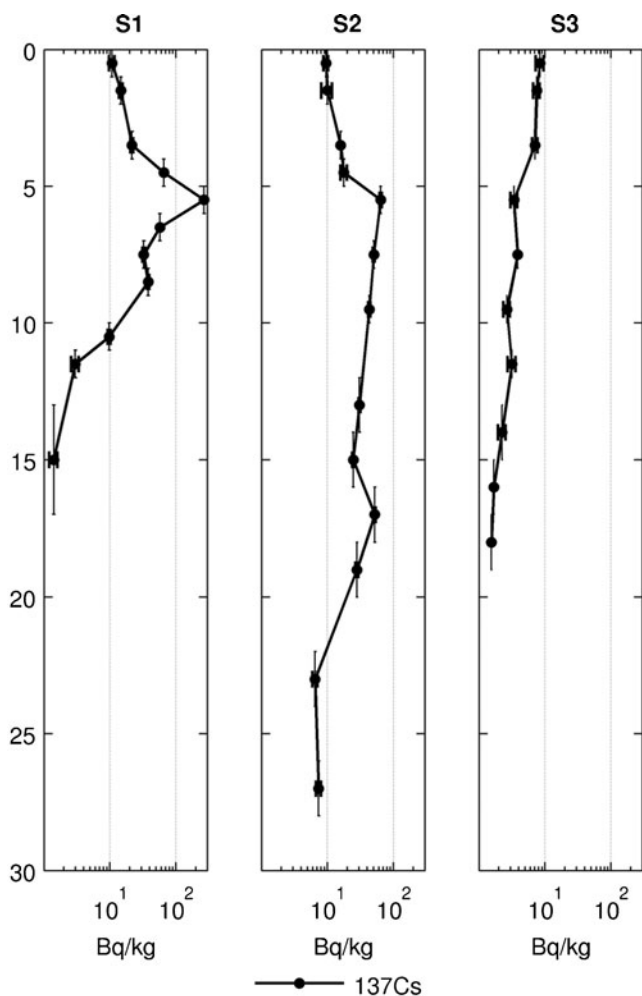


Fig. 6 ^{137}Cs (in Becquerel per kilogram) in all three cores and all depths (in centimeters) that were measured. Horizontal error bars indicate measurement errors of ^{137}Cs . Vertical error bars represent the thickness of the measured layer

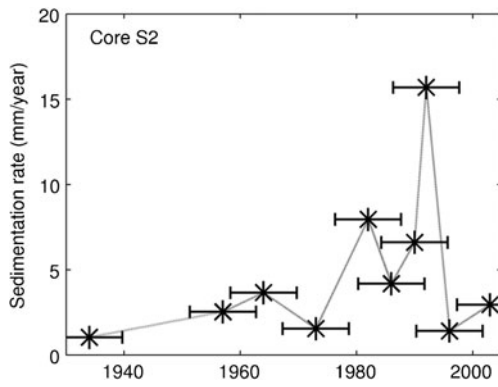


Fig. 7 Marsh sedimentation rates from 1938 to 2003 (millimeters per year). Horizontal error bars represent the averaged uncertainty of the ²¹⁰Pb dating

531 Hydrology and Meteorology

532 Tide gauge data from 1938 to 2007 were analyzed for
 533 historic trends (Fig. 8) and compared to ²¹⁰Pb-derived
 534 marsh elevations (Fig. 9). It is shown that the MSL at the
 535 tide gauge *Hoernum Hafen* within that time period rose by
 536 about 2.1 mm year⁻¹, slightly greater than the global rate of
 537 SLR of 1.8 mm year⁻¹ during the same period (Church and
 538 White 2011). The rise of MSL was accompanied by a mean
 539 increase of the MHW of about 4 mm year⁻¹, while the
 540 MLW did not change significantly. The increase of both
 541 MHW and MSL was observed to accelerate since about the
 542 early 1980s. This corresponds to data analyzed by Wahl et
 543 al. (2010), who further pointed out that SLR rates during
 544 the last 10–15 years within the German Bight were higher
 545 than the global average, as reported by Church et al. (2008).

The frequency of storm floods was analyzed based on their
 546 definition given by the German Maritime and Hydrographic
 547 Agency (BSH), where storm floods are water levels
 548 exceeding 1.5 m above MHW. Considering that the
 549 averaged yearly MHW from 1938 to 2007 is 0.9 m above
 550 NN, this corresponds to a storm level of 2.4 m NN. It was
 551 shown that storm frequency experienced a linear increase of
 552 0.06 events year⁻¹, while 1984 and 1990 were marked by
 553 the greatest number of storms with 10 and 11 events,
 554 respectively. During the same period, the intensity of these
 555 storms did not significantly change.
 556

As MHW increased rapidly within the last 50 years, salt
 557 marsh growth was not able to keep pace with this rise. An
 558 average of 2.6 and 1.3 mm year⁻¹ was lost relative to MHW
 559 at core location S1 and S2, respectively. Core location S3
 560 was observed to accrete at a rate of 2.5 mm year⁻¹ since
 561 1996. Meanwhile, during that period, MHW strongly
 562 increased with a rate of 11 mm year⁻¹ (Fig. 9). Marsh
 563 elevation of S2 relative to MSL, on the other hand,
 564 increased by 0.7 mm year⁻¹, while S1 decreased by
 565 0.6 mm year⁻¹ (Fig. 9).
 566

Influence of Storm Frequency and Intensity on Accretion Rates

The marsh at core location S2 floods when the water level
 567 reaches 134 cm (NN; Fig. 2). Multiple linear regression
 568 analysis shows that storm intensity ($\beta=0.88, p<0.01$)
 569 describes accretion rates better than storm frequency
 570 ($\beta=0.06, p>0.1$), when considering all inundation events
 571 (Fig. 10). However, with an increasing maximum storm
 572
 573
 574

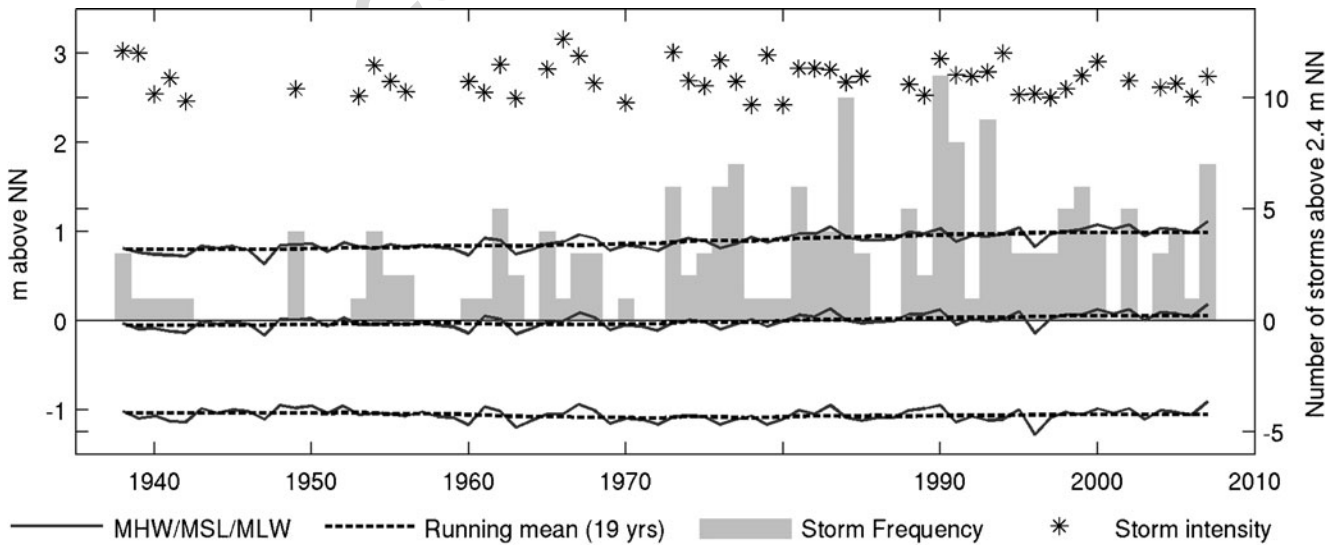


Fig. 8 Hydrological data from 1938 to 2007 at tide gauge *Hoernum Hafen* (Sylt). Storm intensity (stars), mean tidal high water (MHW), mean sea level (MSL), and mean tidal low water (MLW) are given in meter above “normal null” (NN). A 19-year running mean is given for

MHW, MSL, and MLW. Storm frequency (bars) is the number of storm events that exceeded the level of 2.4 m above NN. Source: Wahl (2010, personal communication) and Hartwig (2010)

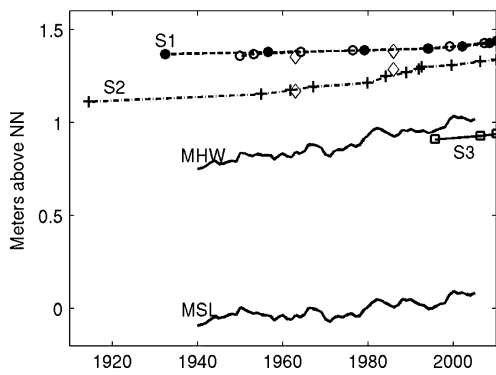


Fig. 9 Historic marsh elevations for all three cores derived from ^{210}Pb and ^{137}Cs data. Due to several data points below the detection limit in core S1, two graphs are displayed, comparing the fastest possible accretion rates (open circles) with the slowest possible accretion rates (filled circles). For validation of the data, the ^{137}Cs peaks are included (open diamonds). The error bars (as shown in Fig. 6) were omitted for clarity. The mean high water and the mean sea level (MHW and MSL; solid lines) are displayed as 5-year running averages

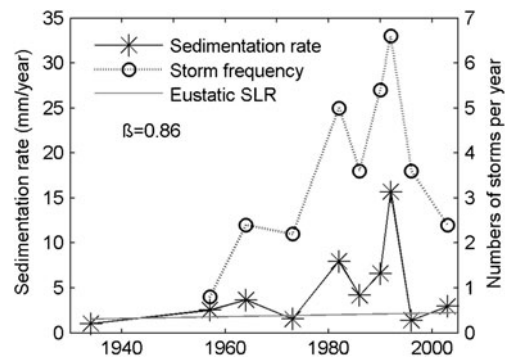


Fig. 11 Comparison of sedimentation rates (stars) at core location S2 with storm frequency (open circles), defined as the number of water levels, exceeding 2.4 m (NN)

575 level, there is a decreasing influence of the storm intensity,
 576 whereas the influence of storm frequency gradually
 577 increases. At a storm level of 260 cm, for example, storm
 578 frequency gives a β of 0.86 ($p < 0.01$), while storm intensity
 579 cannot explain accretion rates significantly ($\beta = 0.04$, $p > 0.1$;
 580 Fig. 11). It appears that the storm level of 152 cm (equals
 581 an inundation height of 18 cm at core location S2) is the
 582 turning point where the influence of storm frequency
 583 becomes statistically significant ($p < 0.1$); meanwhile, the
 584 influence of storm intensity progressively diminishes
 585 (Fig. 12).

Discussion

587 ^{210}Pb geochronologies were used at three salt marsh levels
 588 (a) to determine the age of the marsh and to reconstruct its
 589 evolution, (b) to assess temporal variations of accretion
 590 rates at all three core locations, and (c) to identify the
 591 hydrological parameters influencing marsh accretion.

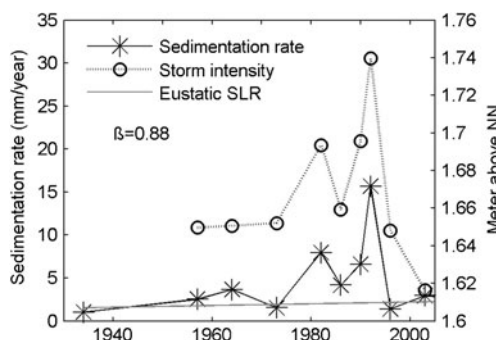


Fig. 10 Comparison of sedimentation rates (stars) at core location S2 with storm intensity (open circles), defined as the mean height for all high water levels exceeding 1.34 m (NN)

Support for Accretion Rate Calculation

592

In contrast to the continuous deposition of ^{210}Pb on the
 593 marsh surface, ^{137}Cs serves as a marker horizon that can be
 594 used to calculate mean accretion rates in between the years
 595 1963, 1986, and 2009, when the cores were extracted. It is,
 596 therefore, an additional independent geochronometer that
 597 allows for comparison with the ^{210}Pb dating method, but
 598 not for derivation of higher resolved temporal accretion
 599 rates. Both the 1963 and 1986 peaks were found in cores S1
 600 and S2. In core S1, ^{137}Cs overestimates accretion rates
 601 compared to the ^{210}Pb data and the aerial photograph of
 602 1937 (Fig. 9). It seems that ^{137}Cs has been transported
 603 further down in the soil column, although we are not certain
 604 of the mechanism. A very good agreement between the
 605 three methods (^{137}Cs , ^{210}Pb , and aerial photographs) is
 606 found in core S2 and in core S3. ^{137}Cs in core S2 shows
 607

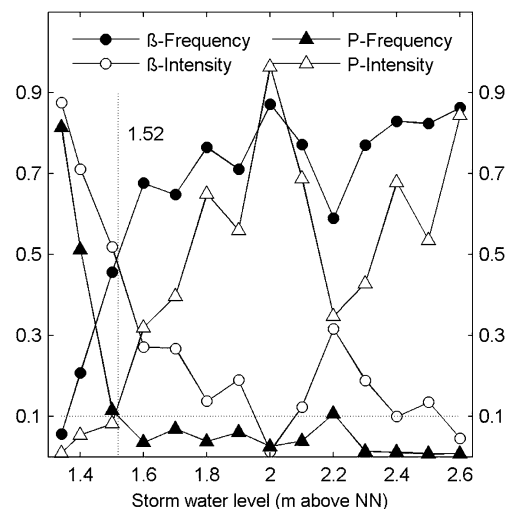


Fig. 12 Results of linear regression analysis: β and p values (circles and triangles, respectively) for different storm water levels (meters above NN) are shown. β and p values refer to results of multiple linear regression analyses of storm frequency (filled symbols) and storm intensity (open symbols) with sedimentation rates at core location S2

608 two distinct peaks at depths of 5.5 and 17 cm. ²¹⁰Pb dating
 609 is supported by these peaks; the layers at depths of 5.5 and
 610 17 cm were estimated to be deposited in 1992 and 1961,
 611 respectively (Fig. 9). The absence of a peak in core S3 also
 612 supports what we can observe in the aerial photographs, as
 613 well as the ²¹⁰Pb dating suggesting that the marsh
 614 development started after 1988.

615 Grain size data also support the ²¹⁰Pb dating in core S2
 616 to a certain degree. Two layers (5.5 and 9.5 cm) show an
 617 elevated fraction of coarse-grained sediments in comparison
 618 to the other layers (Fig. 3). ²¹⁰Pb dating calculates that
 619 these layers were deposited in 1992 (±5 years) and 1985
 620 (±5 years). In 1990, the highest number of storms occurred
 621 during the last 70 years (Fig. 8), two of which were among
 622 the five most severe storm surges (measured at *Hörnum*
 623 *Hafen*) in recorded history. Meanwhile, the year 1984 was
 624 marked by the second most frequent storm events during
 625 the last 70 years (Fig. 8). We expect that the higher energy
 626 of storms allows for the transport of coarser-grained
 627 materials onto the marsh by both waves and currents.

628 Mean Accretion Rates

629 The spatial pattern of mean accretion rates shows a
 630 decrease from the low marsh toward the inner marsh
 631 (Table 1). This is a typical spatial phenomenon observed on
 632 many marshes by various authors (Bartholdy et al. 2004;
 633 Cahoon and Reed 1995; French and Spencer 1993; Pethick
 634 1981; Temmerman et al. 2003a). Meanwhile, the mean
 635 accretion rate for the pioneer marsh zone is lower than the
 636 one in the low marsh zone. Resuspension of sediment
 637 during the early stage of the marsh development could be
 638 the reason for this pattern. The measured accretion rates
 639 generally compare well with mean accretion rates measured
 640 on the peninsula of Skallingen, where Bartholdy et al.
 641 (2004) found the accretion rate on the Skallingen marsh to
 642 vary between 2 mm year⁻¹ on the inner marsh and
 643 4 mm year⁻¹ on the outer marsh. Measurements of Kirchner
 644 and Ehlers (1998) in the eastern part of Sylt, however, have
 645 shown much higher accretion rates between 5.8 and
 646 15.2 mm year⁻¹, but grain sizes in this section of Sylt are
 647 much finer, indicating different hydrodynamic conditions
 648 favoring settling of sediment.

649 Temporal Variations of Accretion Rates

650 S2 shows strong temporal variations of accretion with a
 651 period of higher rates found during the 1980s and 1990s
 652 including two distinct peaks in 1982 and 1992 and a rapid
 653 decrease in accretion rates during the last 20 years. This
 654 pattern strongly resembles the periods of high storm
 655 activity (Figs. 10 and 11). Although temporal variations of
 656 accretion rates in S1 should only be considered as an

approximation, a trend toward increasing accretion rates 657
 current values between 3 and 3.5 mm year⁻¹ is observed. 658
 The few data points that exist for S3 show a constant 659
 accretion rate of 2.5 mm year⁻¹. 660

661 Accretion rates at core location S2 during calm periods
 662 seem to lie close to the rates of eustatic SLR (Figs. 10 and
 663 11), while, during stormy years, accretion rates are much
 664 higher than SLR. This behavior shows the importance of
 665 storms for the resilience of salt marshes toward eustatic
 666 SLR. Even though accretion rates during calm years are
 667 still sufficiently high, a considerable decrease in storminess
 668 would lower the marsh elevation and possibly affect the
 669 marsh zonation.

Historic Development of the Marsh 670

671 Using the data presented in the previous section, it is
 672 possible to reconstruct the historic evolution of the marsh:
 673 Marsh development at our study site seems to have started
 674 after about 1915. Prior to marsh development, we assume a
 675 bare sandy beach slope at that location. The age of the base
 676 layer at core location S2 indicates that marsh development
 677 started during a period of rapidly increasing MHW and
 678 MLW levels (Jensen and Mudersbach 2004; Wahl et al.
 679 2010). Increasing inundation frequencies may have trig-
 680 gered the development of pioneer marsh vegetation, such as
 681 *Salicornia* and *Suaeda*, probably sheltered by a seaward
 682 sandy ridge. At some point, the sandy ridge at the seaward
 683 edge moved offshore and a cliff arose due to increased
 684 hydrodynamics caused by more frequent inundation events.
 685 As a consequence, a channel developed in front of the marsh,
 686 which probably acted as a tidal channel, building up levees at
 687 the landward and seaward side (Pedersen and Bartholdy
 688 2007). According to the aerial photograph in 1988, the
 689 pioneer marsh started to develop on top of the sand bar and
 690 slowly spread toward the cliff at the marsh edge, stabilizing
 691 the latter. Decreased hydrodynamics in the vicinity of the
 692 sand bar, a constantly low MLW level, and a spread of the
 693 invasive species *S. anglica* after 1987 (Loebl et al. 2006)
 694 may have triggered that process. Today the cliff is located far
 695 inside the marsh, and the pioneer marsh seems to expand
 696 further onto the tidal flat and accrete faster than MSLR.

697 The inner part of the marsh accreted with the slowest
 698 overall accretion rate. While the base layer is approximately
 699 30 cm above the base layer in the central part of the marsh,
 700 marsh development started between 1925 and 1937. During
 701 the first few decades of growth, aerial photographs indicate
 702 that considerable amounts of sand may have been blown
 703 onto this part of the marsh, although accretion rates were
 704 too low to detect significant temporal variations. The
 705 topography of the marsh platform has been flattened during
 706 the last 75 years due to higher accretion rates in the middle
 707 of the marsh than on the landward side of the marsh.

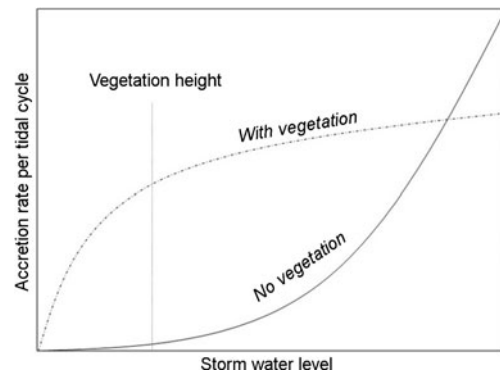
708 Influence of Storms on Accretion

709 The analysis presented in this study focuses on the
 710 historical vertical development of a single salt marsh. It is
 711 shown that an increase of storminess positively correlates
 712 with marsh accretion rates. While many authors have
 713 focused on the destructive, erosional influence of storms
 714 on salt marshes (e.g., Callaghan et al. 2010; Mariotti and
 715 Fagherazzi 2010; van de Koppel et al. 2005), our methods
 716 are not able to reproduce the lateral marsh development and
 717 their implications on accretion processes but rather the
 718 vertical marsh accretion.

719 Vertical accretion (at core location S2) was shown to be
 720 closely linked to past storm patterns with accretion rates of
 721 up to 16 mm year⁻¹ in very stormy years. Linear regression
 722 analysis showed that both storm intensity and storm
 723 frequency are important factors influencing accretion rates.
 724 Storm frequency does not seem to be the driving factor of
 725 accretion if storm levels do not rise higher than 18 cm
 726 above the marsh surface at S2. The importance of
 727 frequency greatly increases above this threshold level
 728 (Fig. 12). We hypothesize that the presence and the height
 729 of the vegetation could be the reason for this threshold.

730 Assuming that the threshold value we found for the core
 731 S2 is true for different marsh elevations as well, we can
 732 identify two different driving factors for marsh accretion
 733 in the lower and the higher parts of the marsh: In the lower
 734 marsh zones, where inundation is frequent and usually
 735 higher than the vegetation height, accretion rates increase
 736 with increasing inundation frequency. Inundation frequency
 737 at low elevations, in turn, is highly correlated with the
 738 MHW level and possibly caused by SLR. In the higher
 739 parts of the marsh, where inundation is less frequent and
 740 heights of the inundation are often lower or just slightly
 741 higher than the vegetation height, we expect the mean
 742 strength of the storms to be more important than the actual
 743 number of storms.

744 These findings are especially relevant when considering
 745 possible implications on storm-related accretion processes.
 746 Assuming that more sediment is in suspension when storm
 747 water levels are high, the solution of the mass balance
 748 equation for incoming and settling sediment (Mudd et al.
 749 2010; Temmerman et al. 2003a) results in an exponential
 750 increase of accretion rates with higher storm water levels
 751 (Fig. 13; Temmerman et al. 2003a). Considering our results,
 752 the effect of storm intensity on salt marsh accretion is
 753 suggested to be stronger at low storm water levels than at
 754 high storm water levels, resulting in a logarithmic rather
 755 than an exponential relationship between storm water levels
 756 and accretion rates (Fig. 13). The height at which the
 757 logarithmic curve starts to flatten is hypothesized to be
 758 connected to the height of the vegetation. Vegetation has a
 759 strong effect on flow velocities and the Reynolds number



760 **Fig. 13** Conceptual scheme of how vegetation is influencing
 761 accretion rates at various storm water levels. The “no vegetation
 762 scenario” does not reflect the resuspension happening in absence of a
 763 vegetation canopy

764 within the vegetation canopy and can therefore influence
 765 the flocculation and break-up processes of the sediment in
 766 suspension (Winterwerp 2002). Low flow velocities within
 767 the vegetation canopy could therefore result in larger floc
 768 sizes and corresponding settling velocities, while high flow
 769 velocities could decrease the floc sizes and decrease the
 770 settling velocities of the sediment particles (Bartholomä et
 771 al. 2009). This would infer that the depth-averaged settling
 772 velocity would decrease with inundation heights that
 773 exceed the vegetation height and result in a flattening of the
 774 above mentioned curve. We must emphasize that this
 775 hypothesis should be further investigated by explicitly
 776 examining flocculation processes over salt marsh surfaces
 777 during storm events.

774 **Conclusions**

775 In accordance with the available literature, ²¹⁰Pb has proven
 776 to be a good tool for determination of sediment accretion
 777 rates on salt marshes. ¹³⁷Cs data and aerial photographs
 778 independently supported the ²¹⁰Pb dating and the derived
 779 accretion rates. However, restrictions and limitation of the
 780 method can be recognized in core S1 and S3. Based on the
 781 results of this study, we make the following conclusions:

- 782 1. While not always coincident with one another, both
 783 storm frequency and intensity seem to affect salt marsh
 784 accretion rates. In very stormy years, accretion rates
 785 can increase 5-fold above mean value. Frequent and
 786 strong storms have shown that they can lead to
 787 accretion rates which are higher than MSLR and
 788 therefore are considered as an important factor for the
 789 ability of marshes to keep pace with eustatic SLR.
- 790 2. We show that eustatic SLR slightly increases accretion
 791 rates during calm wind periods, while a surplus of
 792 accretion is observed in stormy years, leading to a net
 793 increase of marsh elevation relative to mean sea level,

794 accompanied by a slight decrease in elevation relative
 795 to the more rapidly rising mean tidal high water level.
 796 3. There exists a threshold for inundation depth (18 cm)
 797 on the investigated marsh at which the relative
 798 importance of storm frequency and intensity reverses.
 799 Storm intensity seems to be the driving factor for high
 800 accretion rates below an inundation depth of 18 cm,
 801 while storm frequency is more important above this
 802 inundation depth. The influence of vegetation is
 803 suggested to be the reason for this threshold.

804 4. The existence of a threshold for inundation depth
 805 implies that accretion of higher marsh zones is crucially
 806 dependent on the storm intensity, while lower marsh
 807 parts rather depend on the development of storm
 808 frequency and mean high water level, possibly influ-
 809 enced by a mean sea level rise. The different major
 810 driving factors for the marsh zones may lead to changes
 811 in marsh zonation, if storm activity and SLR rate
 812 change in the future.

813 Further investigation, including short-term measure-
 814 ments of accretion rates during storm events and modeling,
 815 is necessary to better understand the singular effects of
 816 storm intensity, storm frequency, and sea level rise on
 817 accretion processes. Also, more marshes in mesotidal
 818 environments need to be investigated to test whether a
 819 similar critical inundation depth threshold exists elsewhere
 820 and the hypothesis of vegetation being responsible can be
 821 verified.

822
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- Q1. The citation “Malvren Mastersizer 2000” (original) has been changed to “Malvern Instruments Ltd 2010”. Please check if appropriate.
- Q2. “Wahl et al. 2011” was cited here but not found in the reference list. Please provide complete bibliographic information.
- Q3. Figure 3 – contains cut off data. Please provide replacement otherwise, please advise if we can proceed with the figure/s as is.
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