Primitive olivine-hosted melt inclusions (MI) are a useful means to estimate the pre-eruptive volatile contents of a volcanic melt. However, post-entrapment processes complicate this approach. In particular, crystallization of the host phase along the wall of the MI and diffusion of H\(^+\) through the host cause CO\(_2\) and potentially S or other volatiles to exsolve from the melt to a separate fluid bubble. Recently, experimental rehydration and Raman spectroscopy have become potential methods for restoring the volatile contents of MI by rehomogenization or through mass balance calculations respectively. In order to compare these two approaches, we have studied MI from a single suite of samples from Klyuchevsky volcano (Kamchatka Arc) that have been treated with both experimental rehydration and analyzed using Raman spectroscopy. The maximum MI CO\(_2\) contents are in agreement (~4000 ppm) regardless of the method used to account for CO\(_2\) in the bubble, but there is significantly more scatter to lower values using the Raman method which can be attributed to uncertainty related to mass balance calculations and carbonate daughter minerals that have formed at the glass-bubble interface. The presence of S- and C-bearing daughter minerals on the surface of the bubble in unheated melt inclusions indicates that to obtain more confident results with Raman spectroscopy, naturally quenched MIs should be also shortly reheated to dissolve most or all the crystals at the glass-bubble interface. Concerning H\(_2\)O, MI from the unheated tephra samples contain less H\(_2\)O than rehydrated MI in lavas. Determining the original H\(_2\)O content of rehydrated MI is difficult because the H\(_2\)O concentration in the glass is controlled by the conditions during the rehydration experiment. Thus reconciling the initial H\(_2\)O content in primitive arc MIs (and degree of H\(_2\)O loss) still remains a challenging task.

**Authors**

Lowell Moore *
Virginia Polytechnic Institute and State University

Nikita Mironov
Vernadsky Institute of Geochemistry and Analytical Chemistry RAS

Maxim Portnyagin
Vernadsky Institute of Geochemistry and Analytical Chemistry RAS

GEOMAR Helmholtz Centre for Ocean Research Kiel

Esteban Gazel
Virginia Polytechnic Institute and State University

Robert J Bodnar
Virginia Polytechnic Institute and State University
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