



## Magma genesis by rifting of oceanic lithosphere above anomalous mantle: Terceira Rift, Azores

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[1] The Terceira Rift formed relatively recently ( $\sim 1$  Ma ago) by rifting of the old oceanic lithosphere of the Azores Plateau and is currently spreading at a rate of 2–4 mm/a. Together with the Mid-Atlantic Ridge, the Terceira Rift forms a triple junction that separates the Eurasian, African, and American Plates. Four volcanic systems (São Miguel, João de Castro, Terceira, Graciosa), three of which are islands, are distinguished along the axis and are separated by deep avolcanic basins similar to other ultraslow spreading centers. The major element, trace element and Sr-Nd-Pb isotope geochemistry of submarine and subaerial lavas display large along-axis variations. Major and trace element modeling suggests melting in the garnet stability field at smaller degrees of partial melting at the easternmost volcanic system (São Miguel) compared to the central and western volcanoes, which appear to be characterized by slightly higher melting degrees in the spinel/garnet transition zone. The degrees of partial melting at the Terceira Rift are slightly lower than at other ultraslow mid-ocean ridge spreading axes (Southwest Indian Ridge, Gakkel Ridge) and occur at greater depths as a result of the melting anomaly beneath the Azores. The combined interaction of a high obliquity, very slow spreading rates, and a thick preexisting lithosphere along the axis probably prevents the formation and eruption of larger amounts of melt along the Terceira Rift. However, the presence of ocean islands requires a relatively stable melting anomaly over relatively long periods of time.

The trace element and Sr-Nd-Pb isotopes display individual binary mixing arrays for each volcanic system and thus provide additional evidence for focused magmatism with no (or very limited) melt or source interaction between the volcanic systems. The westernmost mantle sources beneath Graciosa and the most radiogenic lavas from the neighboring Mid-Atlantic Ridge suggest a mantle flow from Graciosa toward the Mid-Atlantic Ridge and hence a flux of mantle material from one spreading axis into the other. The Terceira Rift represents a unique oceanic rift system situated within the thickened, relatively old oceanic lithosphere and thus exhibits both oceanic and continental features.

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## 1. Introduction

[2] The mid-ocean ridge (MOR) system represents the largest magmatic feature on Earth with a length of more than 60,000 km. Early studies suggested that the mid-ocean ridges were relatively uniform structures erupting basalts of homogeneous incompatible-element depleted tholeiitic composition [Gast, 1968; Shaw, 1970]. However, more detailed investigations revealed that both tectonic structures and composition of the rocks of the MOR system are highly variable [Dupré and Allegre, 1983; Geshi *et al.*, 2007; Kane and Hayes, 1994]. One important factor that influences the ridge's structure is the spreading rate, which can vary from 160 mm/a (full spreading) at the East Pacific Rise to ultraslow spreading rates such as at the Southwest Indian Ridge (SWIR; 12–16 mm/a [Dick *et al.*, 2003]) or the Arctic Gakkel Ridge (8–13 mm/a [Cochran *et al.*, 2003]). Because the thermal structure and the magma budget of the MOR depend on the spreading rate, the degree and depth of partial melting, the MORB compositions are also affected which is generally reflected by variable major and trace element composition of the basalts [Gast, 1968; Shaw, 1970]. It has also been recognized that MORs are evolving, i.e., the spreading rate varies and ridge segments propagate or become extinct [Kane and Hayes, 1994; MacDonald *et al.*, 1991; Smith *et al.*, 2001]. Thus, it is expected that during the evolution of a spreading segment the compositions of the magmas change as they are affected by variable mantle sources and spreading regimes. Here, we present geochemical data for a segment

of the MOR, which developed from a transform fault into an obliquely ultraslow spreading rift separating old oceanic plateau lithosphere.

[3] The unique setting of the Terceira Rift in the submarine Azores Plateau, with an ultraslow spreading axis above a melting anomaly, allows to address the following fundamental questions: are melting processes and mantle sources homogeneous in the presence of a melting anomaly and, if a heterogeneous mantle is present, are melting processes and mantle source distribution along ultraslow spreading ridges either controlled by rifting, by lithospheric thickness, by mantle temperature or by their composition? This study presents new major element, trace element and Sr-Nd-Pb isotope data from a suite of submarine and subaerial volcanic rocks along the ultraslow spreading Terceira Rift in the Azores. Large scale along-axis geochemical variations suggest deeper melting at the island of São Miguel, farthest from the Mid-Atlantic Ridge, but relatively small along-axis changes in degrees of partial melting. The occurrence of distinct, well defined mantle source compositions with very limited mixing between the magmatic segments implies that, despite the presence of a melting anomaly, the distribution of magmatic activity and mantle sources along axis is mainly controlled by the spreading movement. The occurrence of small-scale heterogeneity also gives evidence that the chemical enrichment of the adjacent Mid-Atlantic Ridge may be the result of mixing between the enriched Graciosa mantle source and a depleted mantle, indicating a flux of

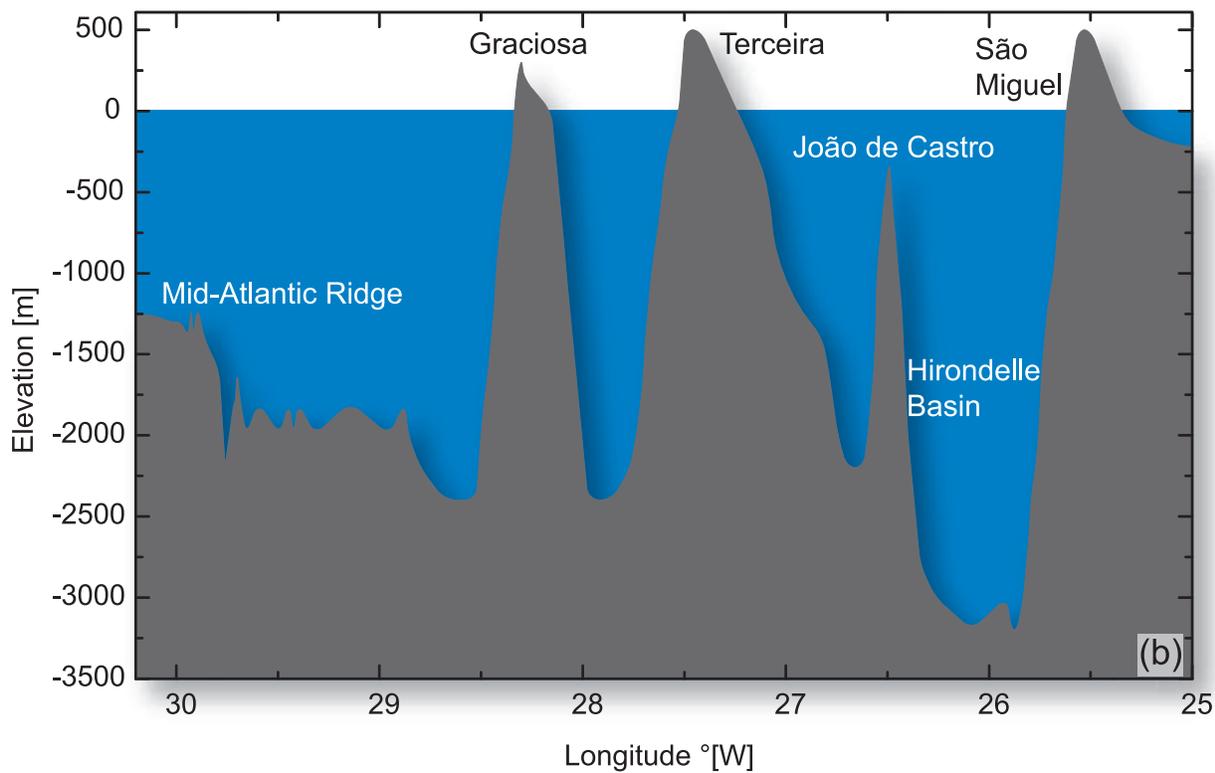
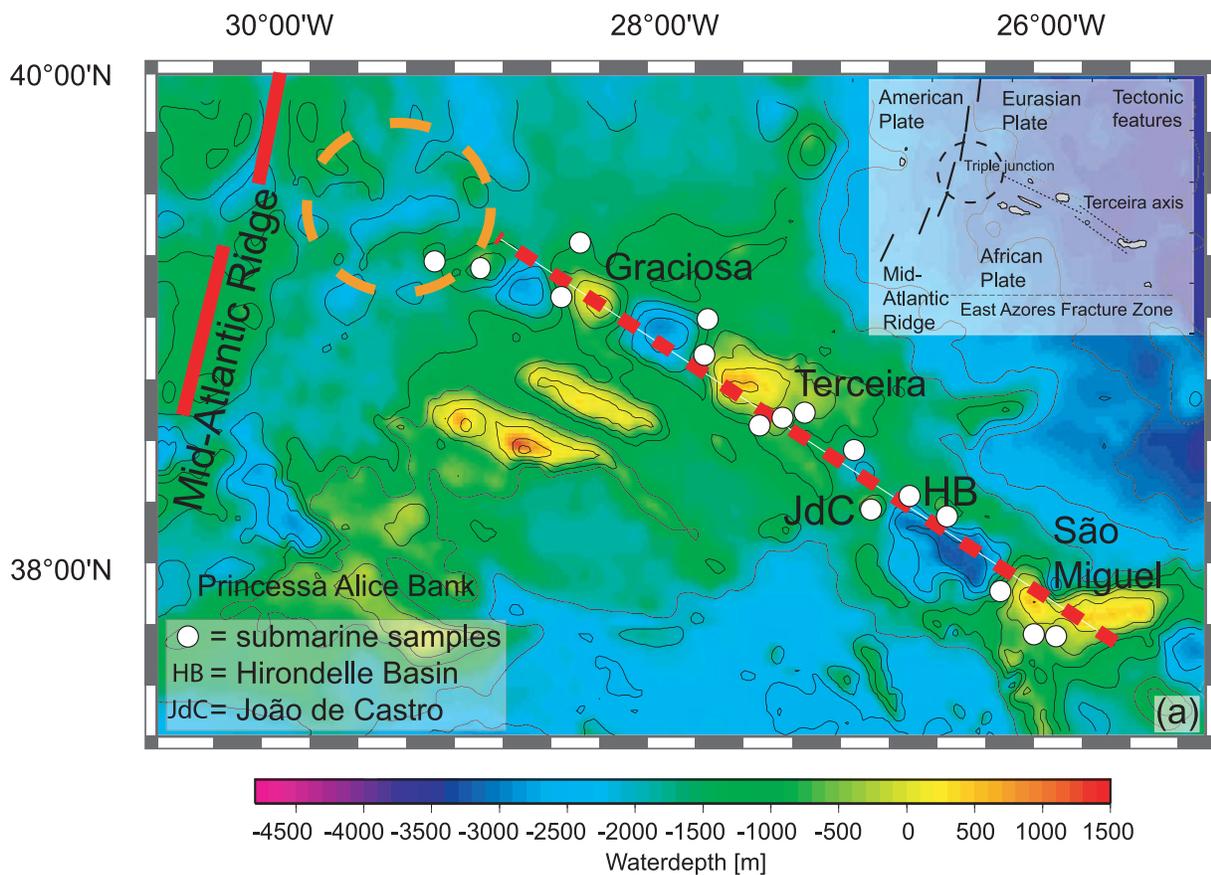


Figure 1

mantle material from one spreading axis into the other.

## 2. Geological Setting

[4] The ultraslow (2–4 mm/a) and obliquely spreading Terceira Rift formed very recently (about 1 Ma ago) [Vogt and Jung, 2004] probably from a transform fault and is rifting old oceanic lithosphere of the northern Azores Plateau (Figure 1a). Together with the Gloria transform fault in the east, the Terceira Rift forms the plate boundary between the African and Eurasian Plates. Three (i.e., São Miguel, Terceira, Graciosa) of the nine volcanic islands of the Azores Archipelago and the submarine João de Castro seamount are situated along the Terceira Rift. Each volcanic center is bordered by deep avolcanic basins (Figure 1b). The volcanically active seamount João de Castro lies between the islands of Terceira and São Miguel and reached subaerial stages in 1638 and 1720 but was eroded soon after [Nunes *et al.*, 2003]. The Azores Plateau probably formed by a melting anomaly in the mantle either as a result of a small thermal plume head [Cannat *et al.*, 1999; Schilling, 1975; White *et al.*, 1979] or of anomalously volatile-enriched mantle, a “wet spot” [Bonatti, 1990; Schilling *et al.*, 1980]. Seismic tomography studies reveal the presence of mantle with anomalously slow seismic velocities beneath the Azores, but a connection to the lower mantle is disputed [Courtilot *et al.*, 2003; King, 2007; Montelli *et al.*, 2004; Ritsema and Allen, 2003]. In fact, the lack of a tail of the mantle anomaly has been interpreted to possibly show a dying plume with a short life time of less than 40 Ma [Silveira *et al.*, 2006]. Lavas from São Miguel and Terceira have relatively primitive He and Ne isotope compositions which indicate the presence of relatively undegassed mantle material beneath the Azores [Madureira *et al.*, 2005]. It was suggested that mantle material from beneath Terceira is flowing into the MAR causing the geochemical anomaly at the spreading center [Moreira *et al.*, 1999b].

[5] GPS and laser measurements show that the islands lie in an extensional regime [Miranda *et al.*, 1998]. On the basis of relative plate motions using the NUVEL-1A model, Vogt and Jung [2004] determined spreading rates of 2–4 mm/a for the recent plate boundary between the Eurasian and African Plates. Such ultraslow extensional movements are comparable to those observed at continental rifts like, for example, the East African Rift [Corti, 2008; Ebinger *et al.*, 1993]. Extensional tectonics in the Terceira Rift are also revealed by magnetic anomalies [Searle, 1980], geometric modeling [Krause and Watkins, 1970], and by extensional tectonic structures on some islands, e.g., at the western end of São Miguel [Beier *et al.*, 2006]. The Terceira Rift and the MAR form the ridge-ridge-ridge plate triple junction between the three bordering plates (Figure 1). Although the precise location of this triple junction is not well defined, focal earthquake mechanisms infer the locus west of Faial and/or Graciosa [Grimison and Chen, 1986, 1988; Udias *et al.*, 1976].

## 3. Methods

[6] The submarine samples were obtained during two cruises with the German research vessel POSEIDON in 1997 (POS 232) and in 2002 (POS 286). The islands of São Miguel, Terceira, and Graciosa were sampled during three field trips between 2001 and 2003.

[7] Most submarine samples dredged along the Terceira Rift are fresh and only few are slightly hydrothermally altered. Volcanic glasses were dredged west of São Miguel, at João de Castro and west of Graciosa. Representative samples have been studied petrographically. Wherever possible glass was separated, hand-picked, washed, and used for the geochemical analyses. Fresh cores were cut from samples without glass, coarse crushed, washed thoroughly in deionized water, and then fine crushed in an agate ball mill.

[8] Major element analyses on whole rocks were carried out on a Philips 1400 XRF spectrometer at

**Figure 1.** (a) Bathymetric chart of the northern Azores platform and the Terceira Rift according to Smith and Sandwell [1997]. Orange dashed circle marks the assumed position of the plate triple junction and the dotted line is the estimated general trend of the Terceira Rift. Subaerial samples were taken on each of the three islands (for detailed sample locations see Table S1 and for São Miguel samples see Beier *et al.* [2006] and Beier *et al.* [2007]). Inset shows the major tectonic features in the Azores; the East Azores Fracture Zone (EAFZ) is the seismically inactive former plate boundary between the Eurasian and African Plates. The recent plate boundary is thought to be in the vicinity of the Terceira Rift. (b) Depth profile along the Terceira Rift showing a segmentation pattern as also observed on other slow spreading ridges, such as the Arctic Gakkel Ridge [Michael *et al.*, 2003].

the Institut für Geowissenschaften, Universität Kiel using fused glass beads. Results for all samples and international rock standards are presented in Table S1 in the auxiliary material and show that precision and accuracy are better than 0.8% ( $2\sigma$ ) and 1% ( $2\sigma$ ), respectively.<sup>1</sup> The major element analyses of glasses were determined on a JEOL JXA8900 Superprobe electron microprobe at the Institut für Geowissenschaften, Universität Kiel. SiO<sub>2</sub>, TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, FeO<sup>T</sup>, MnO, MgO, CaO, Na<sub>2</sub>O, K<sub>2</sub>O, P<sub>2</sub>O<sub>5</sub>, Cr<sub>2</sub>O<sub>3</sub>, and, in some cases, also F, Cl, and NiO were measured. The EMP operated with an accelerating voltage of 15 kV, a beam current of 12 nA and a defocused beam (12 μm). Counting times were set to 20 and 10 s for peaks and backgrounds, respectively.

[9] Trace element analyses were carried out using an Agilent 7500c/s Quadrupole Inductively Coupled Plasma Mass Spectrometer (ICP-MS) at the Institut für Geowissenschaften, Universität Kiel. The samples were prepared following procedures in the work of *Garbe-Schönberg* [1993]. Trace element analyses of the samples along with international rock standards are reported in Table S1 and indicate a standard deviation of the precision and accuracy of <5% and <8% ( $2\sigma$ ), respectively, based on repeated standard measurements.

[10] Sr-Nd-Pb isotope analyses were performed at the Max-Planck-Institut für Chemie in Mainz (MPI) and at IFM-GEOMAR in Kiel. In both labs ~150–200 mg of sample grains were leached in hot 6N HCl for 2 h, ultrasonicated 30 min, and then dissolved using standard digestion procedure described by *Eisele et al.* [2002] and *Abouchami et al.* [2000a]. At the MPI Sr and Nd isotopes were measured on a Finnigan MAT 261 (TIMS) and a Nu Plasma (HR MC-ICP-MS), respectively, while at IFM-GEOMAR, Sr and Nd compositions were both measured on a TRITON TIMS. Sr and Nd isotope ratios on all instruments were obtained in static mode and mass bias corrected relative to  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$  and  $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$ . In Mainz, standard runs of NIST SRM-987 gave  $0.710299 \pm 26$  (2SD,  $n = 16$ ) and  $0.710273 \pm 5$  ( $n = 8$ ) in Kiel. Sr isotope analyses were normalized to a common value of 0.710250 for NIST SRM-987. La Jolla Nd standard measured on the Nu Plasma MC-ICP-MS in Mainz yielded a value of  $0.511862 \pm 24$  ( $n = 14$ ). The data obtained in Mainz were mass fractionation corrected using the dual normalization from

exponential to generalized power law [*Galer and Abouchami*, 2004]. In Kiel, the in-house Nd monitor SPEX yielded  $^{143}\text{Nd}/^{144}\text{Nd} = 0.511710 \pm 5$  ( $n = 5$ ) corresponding to a La Jolla value of  $0.511845 \pm 6$  ( $n = 161$ ). The Nd isotope ratios in Kiel were normalized to a common value of La Jolla 0.511858. Procedural blanks in both laboratories were generally better than 0.2 ng and 0.1 ng for Sr and Nd, respectively.

[11] High precision Pb isotope analyses were carried out at the Max-Planck-Institut für Chemie in Mainz, using the triple spike technique [*Galer and Abouchami*, 1998; *Galer*, 1999]. Samples were loaded onto Re filaments with a silica gel H<sub>3</sub>PO<sub>4</sub> activator and unspiked and spiked sample aliquots were measured on a TRITON TIMS in static multi-collection mode. The mass bias correction estimated from the two runs follows the method outlined by *Galer* [1999]. On the basis of sample duplicate analyses, the external reproducibility is ~160 ppm for  $^{206}\text{Pb}/^{204}\text{Pb}$ ,  $^{207}\text{Pb}/^{204}\text{Pb}$ , and  $^{208}\text{Pb}/^{204}\text{Pb}$ . Standard runs of NIST SRM 981 ( $n = 32$ ) gave average values of  $16.9434 \pm 25$ ,  $15.5010 \pm 24$ , and  $36.7304 \pm 63$  for  $^{206}\text{Pb}/^{204}\text{Pb}$ ,  $^{207}\text{Pb}/^{204}\text{Pb}$ , and  $^{208}\text{Pb}/^{204}\text{Pb}$  respectively. A subset of Pb samples was analyzed on a Finnigan MAT 262 TIMS at IFM-GEOMAR using an external mass bias correction based on repeated NIST SRM 981 measurements. The long-term reproducibility in this lab for NIST SRM 981 ( $n = 189$ ) is  $^{206}\text{Pb}/^{204}\text{Pb} = 16.899 \pm 7$ ,  $^{207}\text{Pb}/^{204}\text{Pb} = 15.437 \pm 9$ ,  $^{208}\text{Pb}/^{204}\text{Pb} = 36.525 \pm 29$ . These values were normalized to the Mainz NIST SRM 981 triple spike values to obtain mass bias factors applied to the sample data. Although conventional Pb isotope data cannot resolve small-scale variations, the large Pb isotopic variation in this particular sample set allows to merge both data sets. Representative major element, trace element, and Sr-Nd-Pb isotope data are given in Table 1. A comprehensive data set is available in Table S1.

## 4. Results

### 4.1. Major and Trace Elements

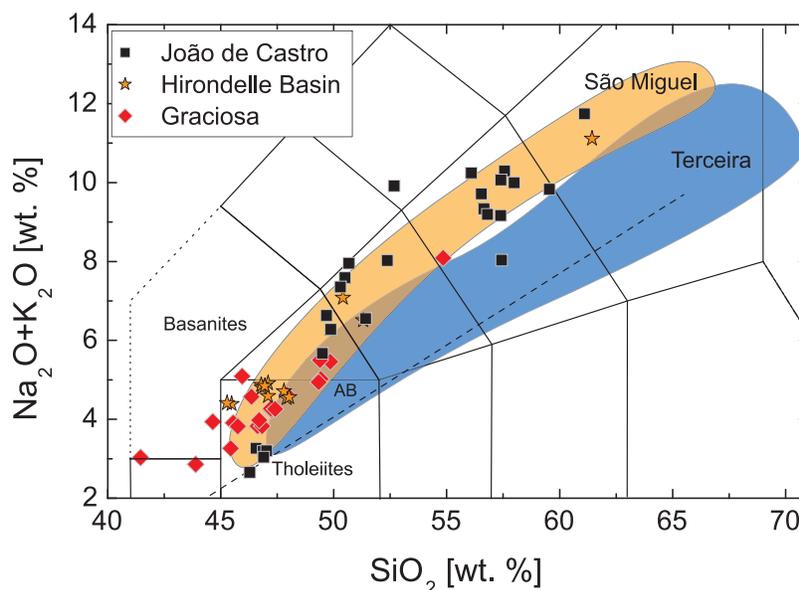
[12] The Terceira Rift lavas are subdivided into four groups representing the volcanic centers of São Miguel, João de Castro (including lavas from the neighboring Hirondele Basin, Figure 1), Terceira, and Graciosa and range from alkali basalts to trachytes on a total alkali versus silica (TAS) diagram (Figure 2). The two major islands of São Miguel and Terceira display different trends with the São Miguel lavas being more alkaline

<sup>1</sup>Auxiliary materials are available in the HTML.

**Table 1 (Sample).** Selected Major Element, Trace Element, and Sr-Nd-Pb Isotope Data of Whole Rocks and Glasses From the Terceira Rift<sup>a</sup> [The full Table 1 is available in the HTML version of this article at <http://www.g-cubed.org>]

Sample	513DS-1	515DS-1	523DS-1	524DS-2	525DS-2	558DS-1	558DS-8	542DS-1	244 DS-1	245 DS-4
Location	W slope of São Miguel	W slope of São Miguel	Banco João de Castro	Banco João de Castro	Banco João de Castro	Banco João de Castro	Banco João de Castro	SW Graciosa	W of Graciosa	W of Graciosa
Latitude (°N)	37°51.933N	37°51.874N	38°10.455N	38°11.287N	38°11.746N	38°13.725N	38°13.725N	39°05.966N	39°10.026N	39°08.184N
Longitude (°W)	25°56.277W	25°59.887W	26°37.895W	26°36.761W	26°35.852W	26°39.068W	26°39.068W	28°16.412W	28°17.959W	28°14.311W
Whole Rock/GL	Whole Rock	Whole Rock	Whole Rock	Glass	Whole Rock	Glass	Glass	Whole Rock	Glass	Whole Rock
TAS classification	Basalt	Trachyandesite	Trachyandesite	Basaltic Trachyandesite	Basalt	Trachyandesite	Basaltic Trachyandesite	Basalt	Basalt	Basalt
Melting depths (GPa)	2.65				2.67		1.57	3.17	3.08	3.57
Melting depths (wt. %)										
SiO <sub>2</sub>	46.95	59.29	55.58	52.39	46.89	56.70	49.42	45.76	45.98	44.85
TiO <sub>2</sub>	2.74	1.28	1.87	2.56	2.01	1.56	3.29	2.73	4.66	2.73
Al <sub>2</sub> O <sub>3</sub>	12.33	18.69	17.64	17.67	10.45	17.75	16.33	15.10	14.60	15.07
Fe <sub>2</sub> O <sub>3</sub>	11.06	5.93	6.90	9.41	10.55	6.66	9.25	10.54	12.35	10.74
MnO	0.17	0.19	0.18	0.20	0.15	0.25	0.19	0.16	0.15	0.17
MgO	12.13	1.58	2.24	3.39	15.78	1.95	3.73	9.67	5.88	9.67
CaO	10.97	3.53	4.64	6.91	11.73	3.50	7.82	11.09	11.62	10.11
Na <sub>2</sub> O	2.56	5.99	6.22	4.07	2.10	5.16	4.10	2.85	2.95	2.90
K <sub>2</sub> O	1.38	4.32	3.93	3.25	1.09	4.78	3.12	0.91	2.14	0.95
P <sub>2</sub> O <sub>5</sub>	0.48	0.38	0.59	0.76	0.33	0.40	0.90	0.39	0.82	0.49
LOI	-	-	-	-	-	-	-	-	-	1.87
Total (ppm)	100.77	101.18	99.79	100.63	101.08	98.71	98.15	99.20	101.16	97.68
Sc	24.1	3.16	7.53	11.2	38.6	2.97	3.41	28.2	30.2	29.5
Cr	996	2.25	1.49	5.01	845	6.57	1.82	465	132	433
Co	44.6	4.95	7.85	16.3	59.4	6.67	6.77	38.9	48.8	45.4
Ni	319	1.75	-	5.94	376	0.10	-	182	100	183
Cu	70.5	3.29	5.37	16.3	162	4.40	3.95	42.2	45.9	46.1
Zn	90.0	92.1	103	112	84.8	111	107	103	114	91.9
Mo	1.90	0.97	4.63	3.65	1.29	5.42	4.93	1.39	-	-
Rb	28.5	1090	101	77.8	27.2	118	108	15.4	22.3	20.1
Sr	622	543	593	776	391	847	766	491	684	521
Y	19.3	33.7	41.8	36.8	18.2	38.2	36.6	20.3	26.1	26.4
Zr	259	607	447	296	146	376	476	172	279	222
Nb	48.0	130	96.8	81.0	31.8	99.2	87.4	33.7	57.5	45.6
Cs	0.32	0.26	0.82	0.61	0.23	0.92	0.86	0.19	0.21	0.26
Ba	404	-	976	792	290	1442	1165	280	275	263

<sup>a</sup> Major element, trace element, and isotopic data from São Miguel are presented by *Beier et al.* [2006]. The whole-rock (WR) major element data were determined by XRF and the trace element data were determined by ICP-MS. Melting depths were calculated using fractionation corrected SiO<sub>2</sub> contents and the equation of *Hause* [1996]. The <sup>232</sup>Th/<sup>238</sup>U ratios have been calculated from the trace element concentrations.



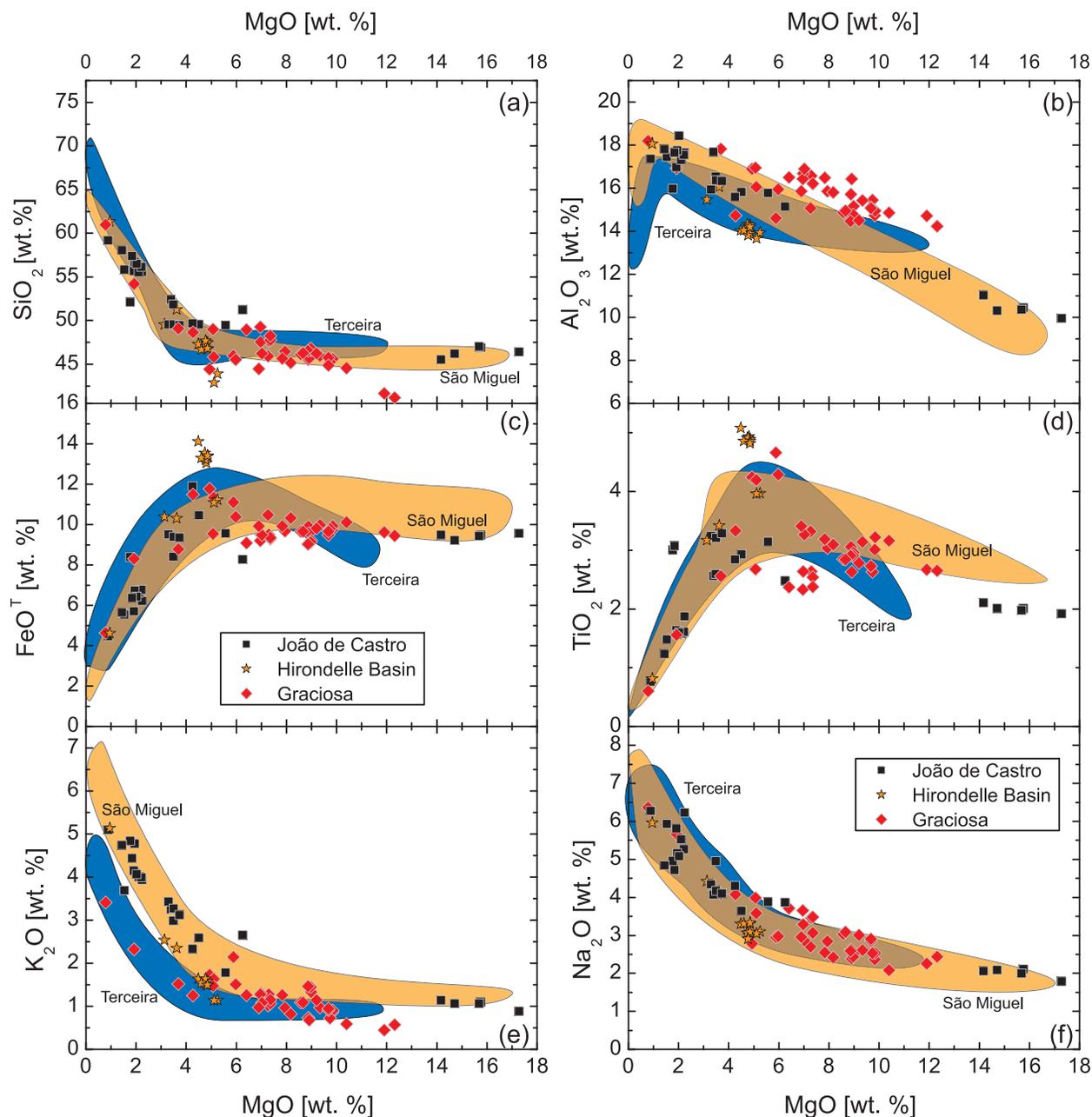
**Figure 2.** Total alkalis ( $\text{Na}_2\text{O}+\text{K}_2\text{O}$ ) versus  $\text{SiO}_2$  on a volatile-free basis according to *Le Maitre* [1989]. The separation line between alkaline (AB) and tholeiitic compositions was taken from *Macdonald* [1968]. The fields of São Miguel and Terceira include subaerial and submarine samples. Major element data from São Miguel are from *Beier et al.* [2006]. On the basis of bathymetry and geochemistry, samples have been subdivided into four groups for clarity: lavas from/around São Miguel, from/around Terceira, from/around Graciosa, and from the João de Castro seamount (including lavas from the neighboring Hirondele Basin that are chemically comparable).

(Figure 2), due to their slightly higher  $\text{K}_2\text{O}$  contents (see below). A few samples also plot in the basanite and phonotephrite fields. Most samples from the João de Castro seamount overlap the field of the São Miguel lavas whereas the Graciosa samples are relatively primitive (most have  $>6$  wt.% MgO) overlapping the Terceira and São Miguel fields (see Figure 3). Here, we will concentrate on the more primitive lavas with MgO contents higher than 5 wt.% in order to determine magma generation and source processes in the mantle.

[13] The  $\text{SiO}_2$  contents in the lavas from São Miguel, João de Castro, Terceira, and Graciosa are relatively constant at  $46 \pm 3$  wt.% in the range between 15 and 5 wt.% MgO (Figure 3a). The  $\text{FeO}^T$ ,  $\text{Na}_2\text{O}$  (Figures 3c and 3f), and CaO (not shown) contents slightly increase with decreasing MgO in the primitive lavas from the four volcanic systems of the Terceira Rift and follow narrow trends (Figure 3) which resemble the well-defined trend of lavas from Sete Cidades volcano on São Miguel [*Beier et al.*, 2006]. In terms of the  $\text{Al}_2\text{O}_3$  contents we find that Graciosa has significantly higher concentrations compared to the other three volcanic systems which lie on one trend of increasing  $\text{Al}_2\text{O}_3$  with decreasing MgO (Figure 3b). The most primitive lavas with  $\text{MgO} > 8$  wt.% also display differences in  $\text{FeO}^T$ , with Graciosa, Terceira and João de Castro having

slightly lower  $\text{FeO}^T$  contents at a given MgO than São Miguel lavas.  $\text{TiO}_2$  concentrations increase from 18 to about 5 wt.% MgO and are generally lower in basalts from Graciosa and Terceira compared to lavas from São Miguel and João de Castro. The most significant differences among the Terceira Rift lavas exist for  $\text{K}_2\text{O}$  with the São Miguel and João de Castro lavas having higher  $\text{K}_2\text{O}$  contents at a given MgO than lavas from the western two volcanic systems Terceira and Graciosa ( $>9$  wt.% for Graciosa specifically, Figure 3e). Similarly, the São Miguel and João de Castro samples are also more enriched in other highly incompatible elements like Rb, Ba, and Ce.

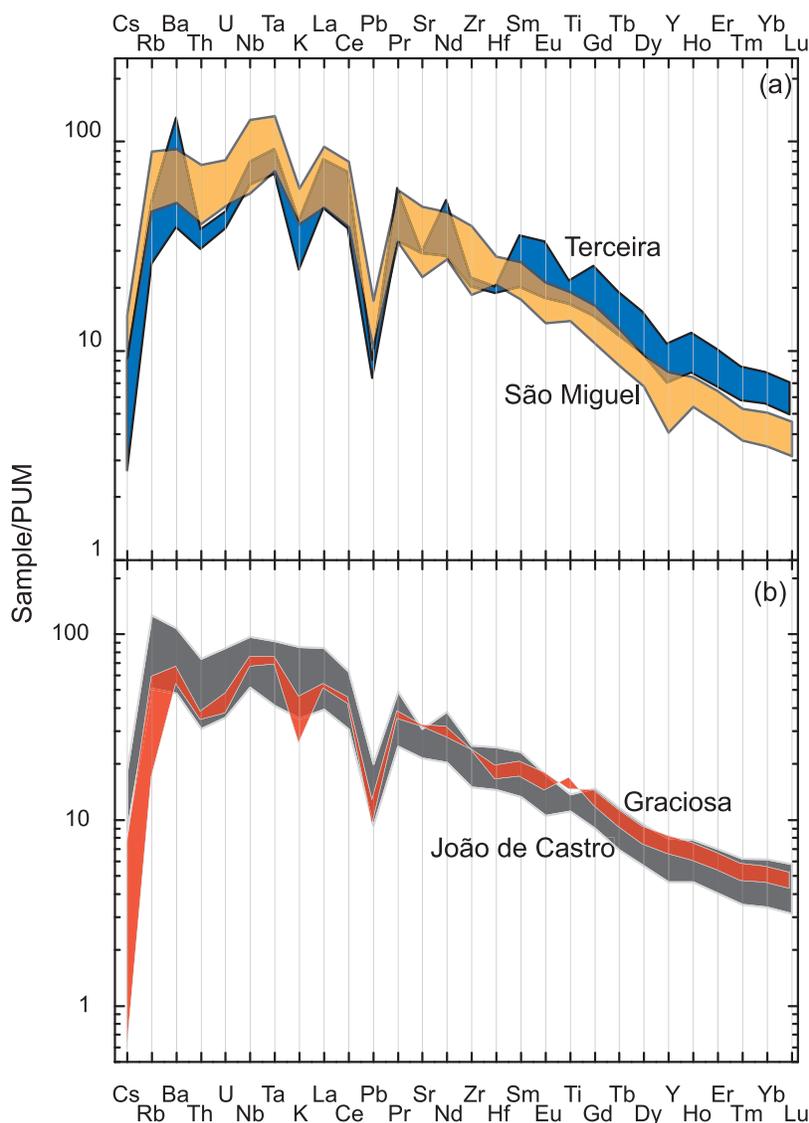
[14] The trace element patterns of the primitive lavas ( $\text{MgO} > 5$  wt.%) from the four volcanic systems of the Terceira Rift lavas are relatively similar with Th, U, K, and Pb troughs and peaks in Rb, Ba, Nb, Ta, and La. The most notable difference is that the São Miguel and João de Castro lavas have higher enrichments of the light REE relative to the heavy REE contents than lavas from the two western Terceira Rift volcanic systems (Figures 4 and 5a). For example, the eastern two structures generally have  $(\text{Ce}/\text{Yb})_N \geq 9$  in contrast to the Terceira and Graciosa lavas which have  $(\text{Ce}/\text{Yb})_N \leq 9$  (see discussion below and Figure 8). The samples from Terceira, Graciosa and João de



**Figure 3.** Major element data versus MgO of the Terceira Rift samples. (a–f) Fields show respective trends of the São Miguel and Terceira data. The most notable differences are observed in the FeO<sup>T</sup> and K<sub>2</sub>O contents (Figures 3c and 3e). Major element data from São Miguel are from *Beier et al.* [2006].

Castro seamount have relatively narrow variations of (Ce/Yb)<sub>N</sub> of 6–9, 7–10, and 9–10, respectively, but the São Miguel lavas vary between 9 and 17, i.e., nearly by a factor of two. Although only three samples from João de Castro have >5 wt.% MgO they resemble São Miguel lavas in terms of (Ce/Yb)<sub>N</sub> but have relatively low (Dy/Yb)<sub>N</sub> similar to basalts from the western islands. In contrast, São Miguel basalts have the highest (Dy/Yb)<sub>N</sub> but the Yb contents in the near-primary magmas are comparable in all lavas suites (1.5–1.7 ppm). Graciosa

and Terceira both generally exhibit peaks in Ti (with the exception of very few samples for Terceira which have a slight Ti trough). Lavas from the João de Castro seamount and many samples from São Miguel also have higher Rb/Nb ratios than the rocks from Graciosa and Terceira (Figure 5). On the other hand, the variations in the Heavy Rare Earth Elements (HREE) and Nb/Zr are not as clear although primitive São Miguel lavas show higher (Dy, Tb/Yb)<sub>N</sub> ratios than the other lava groups (Figure 5). Although they overlap, the



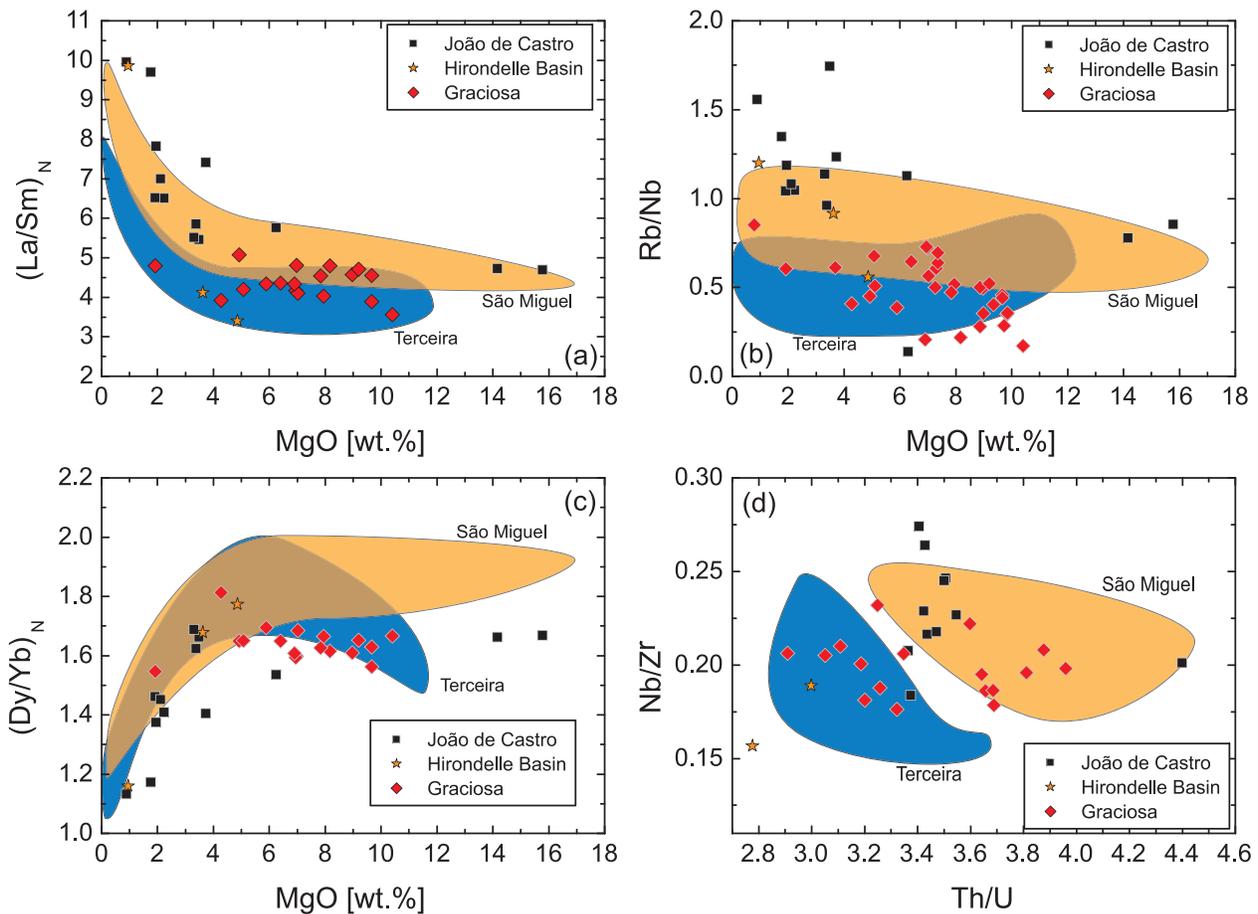
**Figure 4.** Primitive mantle [McDonough and Sun, 1995] normalized trace element pattern of the Terceira Rift lavas. (a) The islands of São Miguel and Terceira and (b) samples from João de Castro and Graciosa.

Terceira lavas tend toward slightly lower Nb/Zr and Th/U (Figure 5) than the lavas from São Miguel and João de Castro, implying different mantle source compositions along the Terceira Rift.

#### 4.2. Radiogenic Isotope Compositions

[15] The large variation of Sr, Nd, and Pb isotope compositions in the lavas from São Miguel is long known and the origin of their distinct mantle sources has been discussed elsewhere [Beier et al., 2007; Widom et al., 1997] and thus will not be discussed in detail here. Lavas from the other three volcanic systems of the Terceira Rift show relatively constant Sr isotope ratios between 0.7033 and 0.7036 but highly variable and distinct  $^{143}\text{Nd}/^{144}\text{Nd}$  (Figure 6a). Thus, the lowest

$^{143}\text{Nd}/^{144}\text{Nd}$  occur in rocks from João de Castro, slightly higher  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios at Graciosa, and the most radiogenic  $^{143}\text{Nd}/^{144}\text{Nd}$  in lavas from Terceira (Figure 6a). Importantly, each volcanic system of the Terceira Rift forms a distinct trend in Sr-Nd-Pb isotope space (Figure 6). The São Miguel lavas show the largest Sr-Nd-Pb isotopic variations and resemble some Graciosa lavas at unradiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$ . The Pb-Pb isotope systematics reveals complex variations for each volcanic system where they form distinct arrays with different slopes (Figures 6d and 6e). São Miguel lavas show the largest Pb isotopic variations with  $^{206}\text{Pb}/^{204}\text{Pb}$  ranging from 19.3 to 20.2 that form steep arrays toward high  $^{207}\text{Pb}/^{204}\text{Pb}$ . João de Castro and Terceira/Graciosa form distinct Pb



**Figure 5.** Trace element (a) La/Sm (chondrite normalized), (b) Rb/Nb, and (c) Dy/Yb (chondrite normalized) ratios versus wt. % MgO of the Terceira Rift lavas, and (d) Nb/Zr versus Th/U ratios of the Terceira Rift lavas. Samples from Terceira and São Miguel are shown as fields to clarify the most notable differences among the Terceira Rift lavas. Chondrite composition from *McDonough and Sun* [1995].

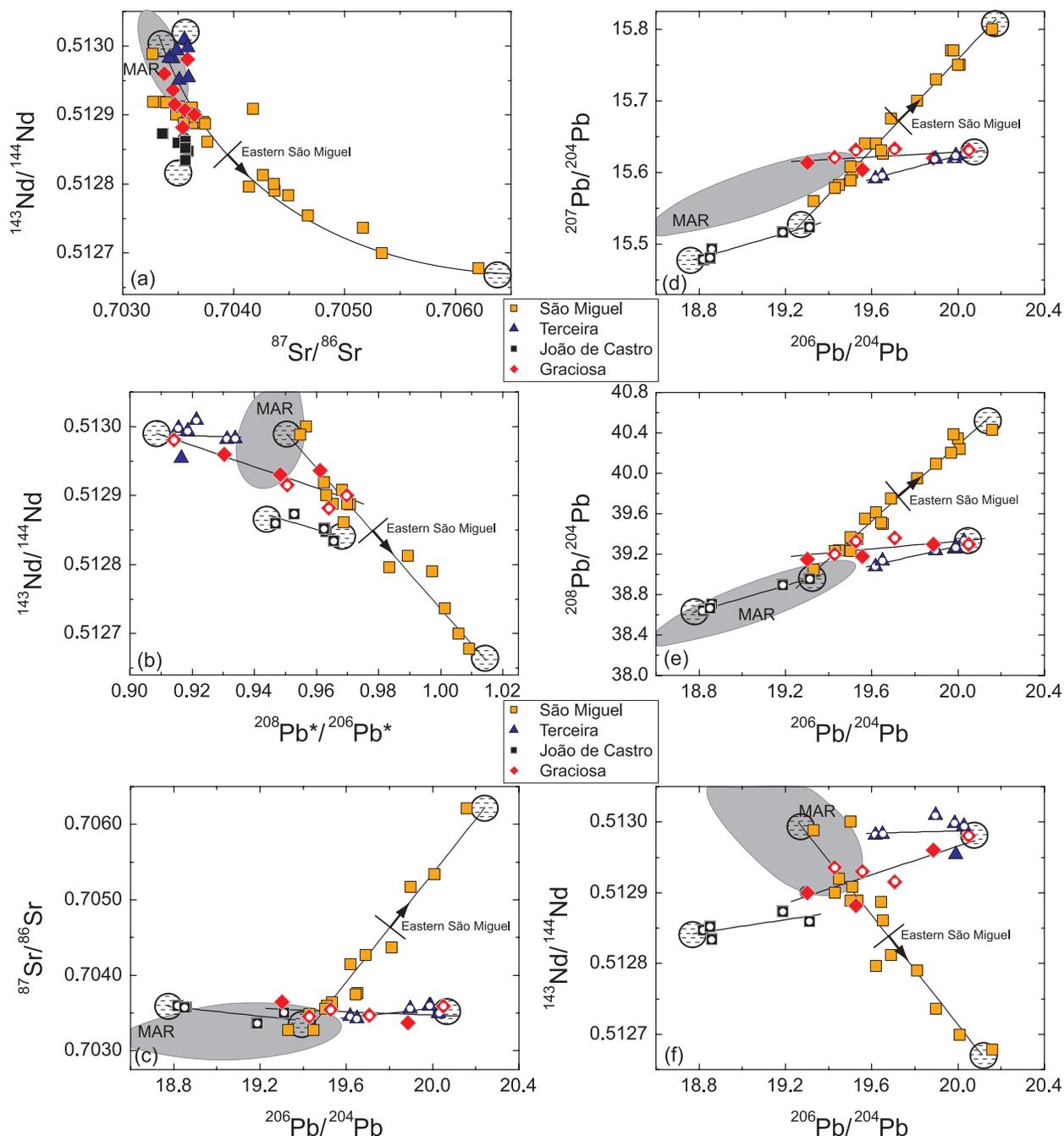
isotopic trends. Lavas from João de Castro have the lowest Pb isotope ratios of the Terceira Rift volcanoes and also the lowest  $^{143}\text{Nd}/^{144}\text{Nd}$  (Figure 6f). The low  $^{206}\text{Pb}/^{204}\text{Pb}$  and Sr isotope compositions of lavas from João de Castro resemble North Atlantic MORB but have significantly lower  $^{143}\text{Nd}/^{144}\text{Nd}$ , and also lower  $^{207}\text{Pb}/^{204}\text{Pb}$  for a given  $^{206}\text{Pb}/^{204}\text{Pb}$  than MORB. Graciosa lavas have slightly more radiogenic  $^{208}\text{Pb}/^{204}\text{Pb}$  and  $^{207}\text{Pb}/^{204}\text{Pb}$  ratios at a  $^{206}\text{Pb}/^{204}\text{Pb}$  range comparable to Terceira but converge at the highest  $^{206}\text{Pb}/^{204}\text{Pb}$  (Figures 6d and 6e). On a  $^{143}\text{Nd}/^{144}\text{Nd}$  diagram all lavas of the Terceira Rift (except São Miguel) lie on a broad positive correlation (Figure 6f), while São Miguel forms a well correlated array, orthogonal to the other Terceira Rift lavas. Importantly, the trends of the different volcanic systems of the Terceira Rift as well as the MAR converge at a composition with  $^{87}\text{Sr}/^{86}\text{Sr} \sim 0.7035$ ,  $^{143}\text{Nd}/^{144}\text{Nd}$

$\sim 0.5129$ , and  $^{206}\text{Pb}/^{204}\text{Pb}$  of 19.5 which has been associated with a mantle component termed FOZO [Hart *et al.*, 1992], recently redefined by Stracke *et al.* [2005], or “C” [Hanan and Graham, 1996]. This composition is represented most clearly by some lavas from Graciosa and those from the western end of São Miguel but appears to be inherent in all Terceira Rift magmas. This mantle source also affects the adjoining MAR from  $\sim 37^\circ$  to  $40^\circ\text{N}$  [Dosso *et al.*, 1999] leading to enriched incompatible element and Sr-Nd-Pb isotope compositions in MORB from this region.

## 5. Discussion

### 5.1. Magma Generation Along the Terceira Rift

[16] Because most of the lavas of the different Terceira Rift volcanoes lie on distinct narrow



**Figure 6.** Isotope systematics of the Terceira Rift lavas. MAR indicates samples from the Mid-Atlantic Ridge from *Dosso et al.* [1999]. Lines indicate linear arrays of São Miguel, Terceira, Graciosa, and João de Castro, respectively. Pb triple spike analyses are marked with a white circle. The dashed circles represent possible mantle source end-members for each mixing array. Arrow marks the trend toward the eastern São Miguel lavas discussed by *Beier et al.* [2007].

trends of major and trace elements versus MgO contents (Figures 2–5) we conclude that each suite of rocks represents a liquid line of descent. In general, these trends resemble those observed for the well-studied Sete Cidades volcano on São Miguel and thus we follow the arguments of *Beier*

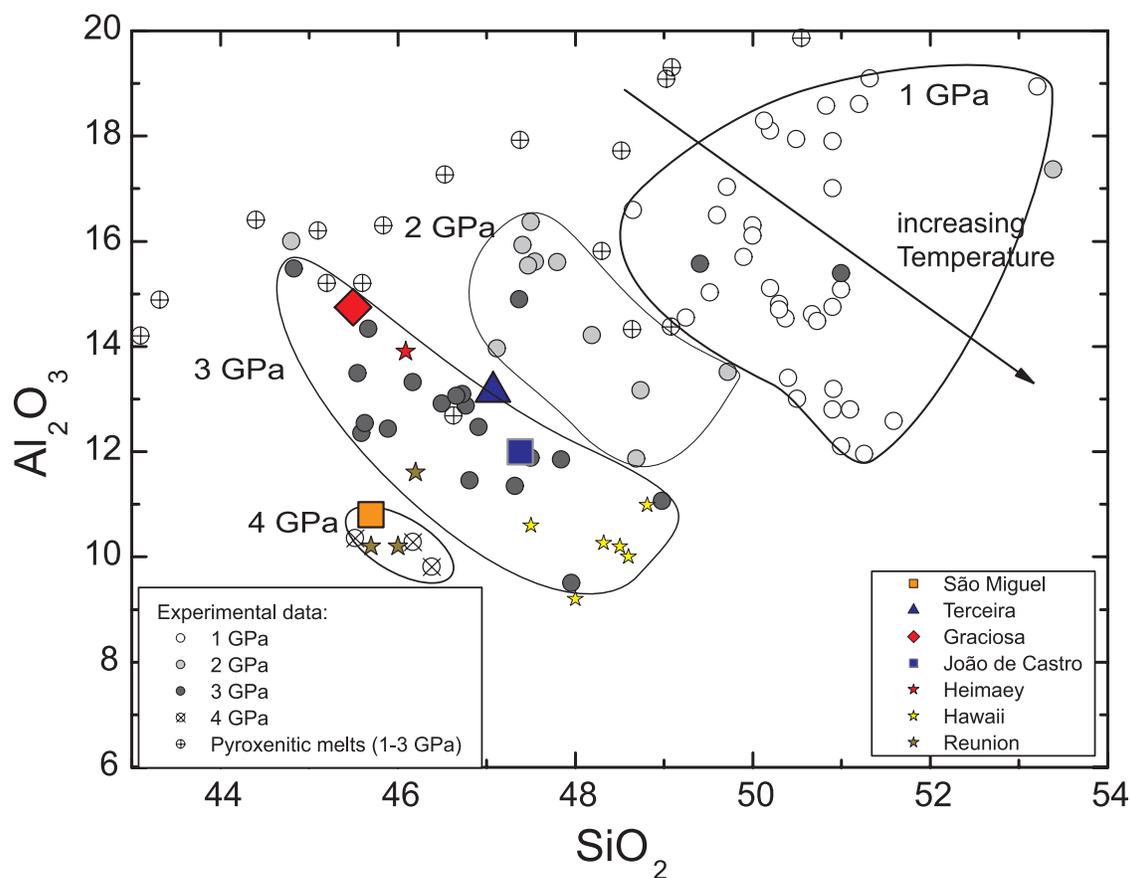
*et al.* [2006] suggesting that the Terceira Rift primary magmas may have about 12.5 wt.% MgO. Lavas with >12.5 wt.% MgO are considered to be the result of the accumulation of olivine and clinopyroxene because many of these lavas contain olivine and clinopyroxene xenocrysts (see also

**Table 2.** Estimated Primitive Magma Compositions From João de Castro, Graciosa, and Terceira Calculated From the Major Element Trends in Figure 3<sup>a</sup>

Island	São Miguel	João de Castro	Graciosa	Terceira
SiO <sub>2</sub>	45.7	47.6	45.6	47.3
TiO <sub>2</sub>	2.7	2.6	2.1	2.1
Al <sub>2</sub> O <sub>3</sub>	10.8	12.0	14.9	13.8
FeO	10.2	9.5	9.5	9.3
MnO	0.2	0.2	0.2	0.2
MgO	12.7	12.3	12.5	11.8
CaO	12.0	11.3	11.0	11.3
Na <sub>2</sub> O	2.2	2.6	2.1	2.4
K <sub>2</sub> O	1.1	1.4	0.8	1.0
P <sub>2</sub> O <sub>5</sub>	0.4	0.4	0.5	0.3
Total	98.0	99.9	99.2	99.5

<sup>a</sup>Primary melt for São Miguel is sample SM0140 from *Beier et al.* [2006]. Only lavas with MgO > 5 wt.% have been taken into consideration to avoid extended crystal fractionation. Primary melts have been estimated to have Mg# of 72 in equilibrium with Fo89 [Roeder and Emslie, 1970].

discussion in the work of *Beier et al.* [2006]) while lavas with <5 wt.% MgO do show evidence for the extensive crystallization of plagioclase and Fe-Ti oxides similar to Sete Cidades volcano on São Miguel. Thus, for each volcanic system we determine an approximate primary magma composition from the major element trends in Figure 3 for all lavas with >5 wt.% MgO. We have estimated the primary magma composition to be in equilibrium with mantle olivine (Fo89), i.e., the primary magmas have Mg# ~72 [Niu and O'Hara, 2008; Roeder and Emslie, 1970]. In Table 2 the compositions of these estimated primary magma compositions for each volcanic system of the Terceira Rift are listed. These compositions are compared to primary magmas from other oceanic and intraplate and rift lavas and the degree and depth of partial melting can be modeled using both major elements and trace element ratios (Figure 7). The major elements SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub> and FeO<sup>T</sup> can



**Figure 7.** Estimated primitive Al and Si compositions (see Table 2) from São Miguel, João de Castro, Terceira, and Graciosa. Lavas from Heimaey, Reunion, and Hawaii are shown for comparison [Clague *et al.*, 1991; Eggins, 1992; Mattsson and Oskarsson, 2005; Sobolev and Nikogosian, 1994]. Experimental data (peridotite) are from Baker and Stolper [1994], Hirose and Kushiro [1993], Jaques and Green [1980], Kushiro [1996], Takahashi [1986], Takahashi and Kushiro [1983], and Walter [1998]. Pyroxenitic melt experiments are from Hirschmann *et al.* [2003] and Kogiso *et al.* [1998].

provide information about variations in melting depth using experimental data from dry peridotite [e.g., *Hirose and Kushiro*, 1998] and pyroxenite [e.g., *Hirschmann et al.*, 2003; *Kogiso et al.*, 1998]. The relatively high  $\text{Al}_2\text{O}_3$  and low  $\text{SiO}_2$  of the Graciosa magma suggests melting of peridotite at pressures of about 3 GPa (i.e., an average depth of  $\sim 90$  km) and relatively low temperatures whereas Terceira magmas formed at higher degrees of melting and probably also higher temperatures but comparable pressure. Interestingly, the Graciosa magmas resemble primitive basalts from the Heimaey volcanoes [*Mattsson and Oskarsson*, 2005], i.e., a propagating rift into  $>3$  Ma-old Icelandic crust and thus from a comparable young setting [see *Mattsson and Hoskuldsson*, 2003, and references therein].

[17] The São Miguel primary magmas have lower  $\text{Al}_2\text{O}_3$  and  $\text{SiO}_2$  than the Terceira magmas which implies higher pressures of partial melting (Figure 7). Lower concentrations of  $\text{Al}_2\text{O}_3$  in the São Miguel basalts could also give evidence for the presence of increased amounts of residual garnet [*Herzberg*, 1995; *Herzberg and O'Hara*, 1998]. The amount of garnet will either be controlled by a change of mantle lithology, e.g., the presence of garnet pyroxenite veins, or by increasing depth of melting and, hence, increased amounts of residual garnet. If the higher HREE ratios at São Miguel would be solely attributed to the presence of garnet pyroxenite veins (see discussion in the work of *Hirschmann and Stolper* [1996]), the  $\text{Al}_2\text{O}_3$  contents would be significantly increased, because a garnet peridotite lithology contains approximately 8 wt.% garnet [*Salters and Longhi*, 1999; *Salters et al.*, 2002], whereas pyroxenite may contain up to 20 wt.% garnet [*Hirschmann and Stolper*, 1996]. Partial melting of garnet pyroxenite would thus lead to high  $\text{Al}_2\text{O}_3$  contrary to the observed low concentrations of the Terceira Rift magmas (Figure 7). We therefore conclude that the Graciosa, João de Castro, and some of the Terceira magmas have formed at lower pressures possibly ranging into the spinel stability field whereas those from São Miguel were generated in the garnet stability field at pressures of  $\sim 4$  GPa (Figure 7). This is in agreement with thickening of the lithosphere toward the east, i.e., with increasing distance from the MAR. It also implies that the oceanic lithosphere is largely intact and that neither rifting nor reheating by the Azores plume did thin it significantly. The ages of the oceanic crust beneath the Terceira Rift volcanoes ranges from about 10 Ma beneath Graciosa [*Cannat et al.*,

1999] to about 45 Ma beneath São Miguel [*Searle*, 1980]. Consequently, the depth of the  $1300^\circ\text{C}$  isotherm for a normal oceanic spreading center increases from 35 to 80 km, respectively, which is in agreement with melting in the garnet stability field beneath the eastern volcanoes. In contrast, a significant part of the melting column beneath Graciosa and Terceira lies in the spinel stability field because the spinel-garnet transitions occur between 70 to 80 km depending on the mantle lithology [*Robinson and Wood*, 1998]. The fact that the lithosphere of the Azores Plateau formed by increased amounts of partial melting due to the presence of a melting anomaly could have additionally thickened the lithosphere, however, the increasing lithospheric thickness from the MAR and the consistency between the estimated depth of the  $1300^\circ\text{C}$  isotherm and geochemical data suggests that the influence of the melting anomaly may indeed be small.

[18] The degree of partial melting beneath the Terceira Rift can be estimated using  $\text{TiO}_2$  rather than  $\text{Na}_2\text{O}$  because experiments have shown that Na partitioning in clinopyroxene is pressure-dependent with  $D_{\text{Na}}$  increasing with increasing pressure [*Blundy et al.*, 1995]. In contrast, Ti is not affected by variable pressures. On the basis of variable  $\text{TiO}_2$  in the Azores magmas it has been argued that São Miguel basalts formed by lower degrees of melting than magmas from the western volcanoes like those from Pico Island [*Prytulak and Elliott*, 2007]. Indeed, we find a variation of  $\text{TiO}_2$  contents with the São Miguel and João de Castro primary magmas having high  $\text{TiO}_2$  of about 2.7 compared to about 2.1 in the Graciosa and Terceira magmas, respectively (Figure 8a). On the other hand, the estimated primary magmas from both eastern and western volcanoes have comparable  $\text{Na}_2\text{O}$  between 2.1 and 2.6 wt.%. A slightly lower degree of partial melting beneath São Miguel and João de Castro seamount is supported by their higher enrichment of light and middle REE relative to Yb (Figures 4 and 8). Alternatively, the relatively enriched Ti contents of the São Miguel and Graciosa lavas at similar Na contents as the remaining Terceira Rift lavas could be due to an increased Ti concentration in the mantle source relative to Na. If an increased Ti content in the mantle source would be responsible for this signature, Ti has to be enriched by a factor of  $\sim 1.5$  relative to the Na concentration to produce the observed trend assuming the same modal composition for all Terceira Rift mantle sources. An enrichment of Ti has also been proposed by *Prytulak and Elliott* [2007] who argue



for the presence of an enriched component from recycled oceanic crust in peridotitic mantle. We agree that a possible mechanism to explain the combined higher Ti at slightly lower Na contents

of the São Miguel and possibly also Graciosa lavas may likely be the presence of an alkaline component, e.g., recycled, subducted oceanic crust which will have a higher Ti but less increased Na contents.

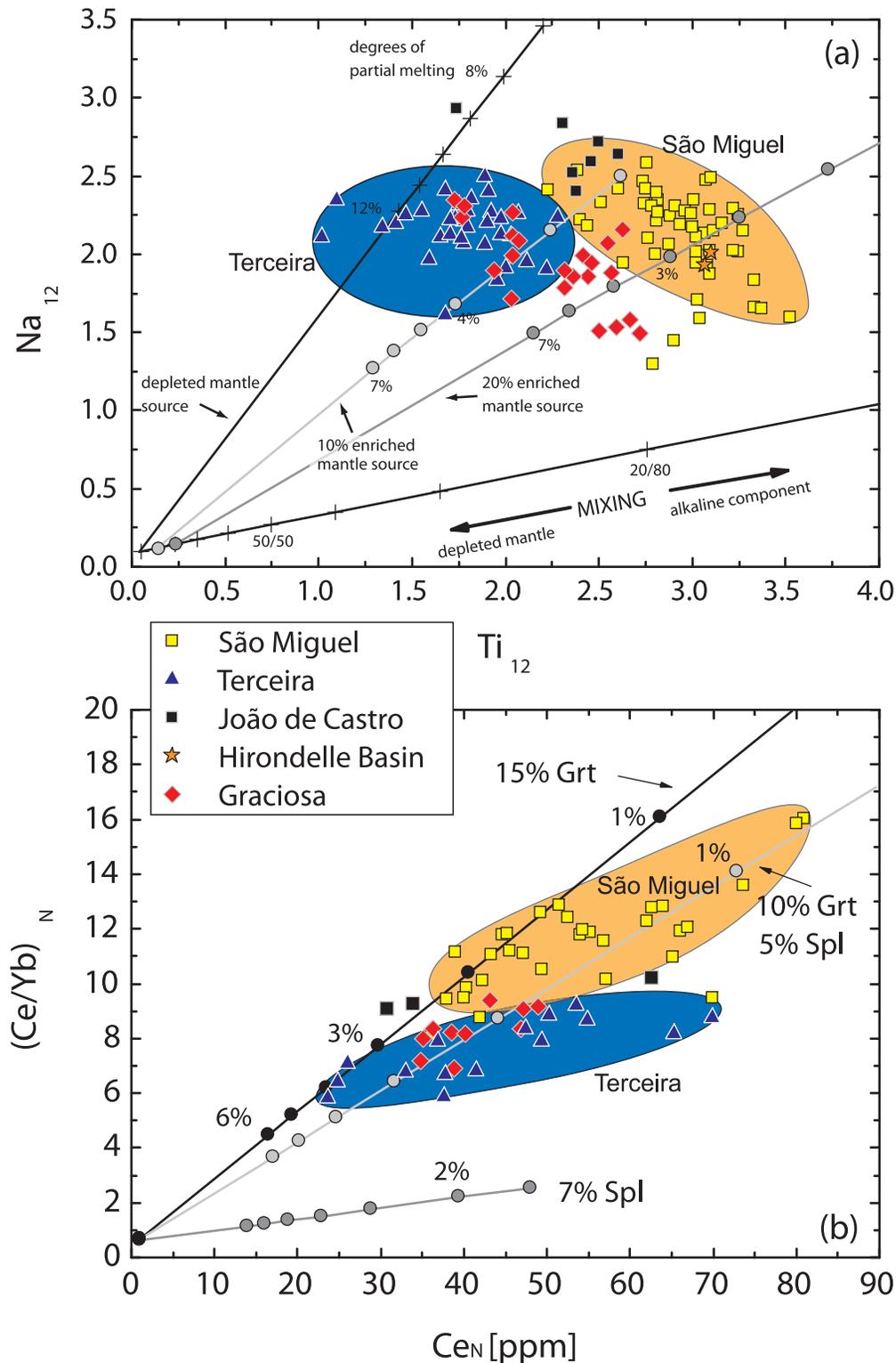
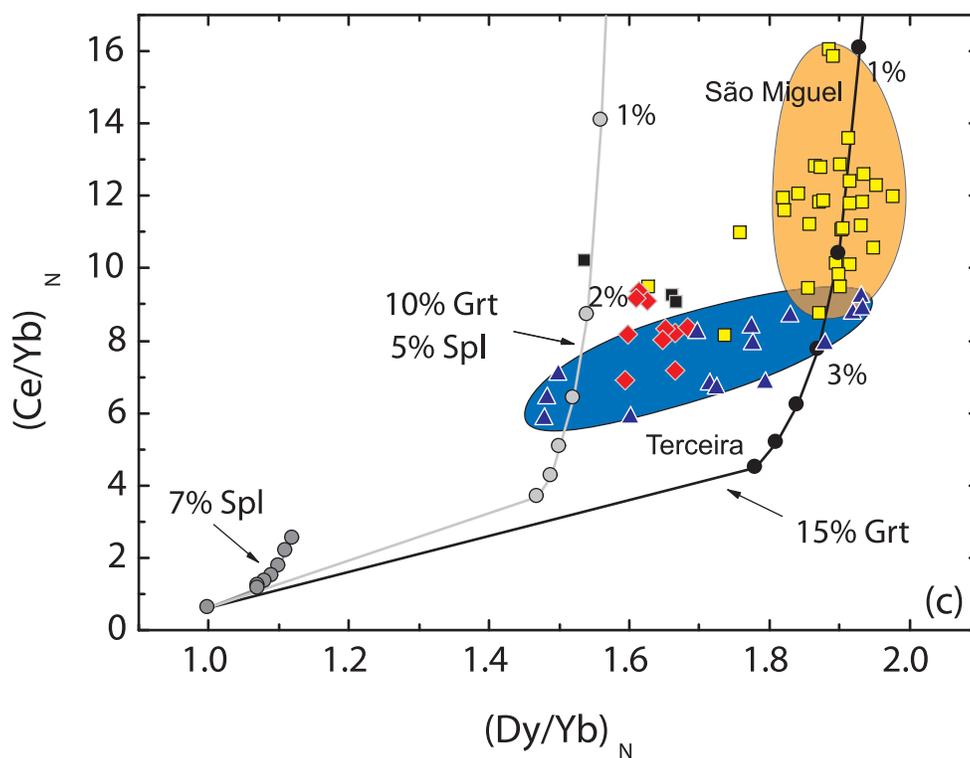


Figure 8



**Figure 8.** (continued)

In the work of *Beier et al.* [2007] we have argued for the presence of recycled oceanic crust in the São Miguel lavas. Whereas recycled oceanic crust will be mainly of pyroxenitic and/or eclogitic composition [see *Stracke et al.*, 1999, and references therein], most major elements and REE will be largely unaffected by changes in mantle lithology, i.e., the changes observed in Ti are too large to be solely explained by varying amounts of pyroxenite beneath the Terceira Rift. However, to ensure the best possible fit we have modeled the trace element and REE systematics (see below) considering the presence of pyroxenite.

[19] Modeling the Na and Ti contents and REE contents of the Terceira Rift lavas (Figure 8a) using the batch-melting equation [*Shaw*, 1970] shows that the Ti-Na compositional variability of the axis lavas may indeed best be explained by mixing small amounts (~10–20%) of an alkali basaltic

component mixed with a depleted upper mantle leaving the REE concentrations relatively unaffected (Figures 8b and 8c, Table 3). The presence of an alkaline component, i.e., an enriched, recycled oceanic crust, in the São Miguel mantle sources has been previously suggested based on quantitative trace element and Sr-Nd-Pb-Hf isotope modeling [*Beier et al.*, 2007; *Elliott et al.*, 2007].

[20] Estimates on the degree of partial melting using REE modeling are generally similar to those derived from the Ti-Na systematics (Figure 8a). The Ce/Yb versus Ce systematics are positively correlated in the Terceira Rift samples indicating a slightly increasing degree of partial melting from São Miguel (lowest) over Graciosa (intermediate) to highest degrees of partial melting beneath Terceira with the lowest Ce/Yb ratio (Figure 8b) as already inferred from the variations in La/Sm ratios (Figure 5). Because the MREE and HREE are most

**Figure 8.** (a) Fractionation-corrected (12 wt.% MgO)  $Ti_{12}$  versus  $Na_{12}$  concentrations. The negative correlation of the Terceira Rift samples can best be explained by mixing an alkaline mantle source (e.g., an alkali basalt similar to sample SM18-8-97-5 from *Beier et al.* [2006]) and a depleted mantle source (Table 3, residual pyrolite source after 10% degree of partial melting). Degrees of partial melting correspond to the melting degrees inferred from the trace element ratios. (b and c) Primitive mantle normalized trace element ratios of  $(Ce/Yb)_N$  versus  $(Ce)_N$  (Figure 8b) and primitive mantle normalized  $(Dy/Yb)_N$  (Figure 8c). Tick marks of melting curves represent 0.5%, 1%, 2%, 3%, 5%, and 6% batch partial melting [*Beier et al.*, 2007] of an enriched mantle source with a trace element composition similar to those implied for the source of the western São Miguel lavas (Table 3).

**Table 3.** Melting Conditions for Figure 8<sup>a</sup>

	Pyrolite Residual After 10% Melting	Alkaline Component	10% Enriched Source (Figure 8a)	20% Enriched Source (Figure 8a)	Enriched Source (Figures 8b and 8c)
Ti	419	35,908	906	1,512	-
Na	779	4,076	912	1,079	-
Ce	-	-	-	-	1.50
Yb	-	-	-	-	0.69
Dy	-	-	-	-	0.99

	50:50 Peridotite/Pyroxenite Mix, Spinel/Garnet Transition	50:50 Peridotite/Pyroxenite Mix, Garnet Stability
Olivine	30%	30%
Orthopyroxene	25%	25%
Clinopyroxene	30%	30%
Spinel	5%	-
Garnet	10%	15%

<sup>a</sup>Note that Ce, Yb, and Dy have been normalized in Figure 8. Ti, Na, Ce, Yb, and Dy are measured in ppm. Pyrolite composition from *McDonough and Sun* [1995]. Residual calculated for 10% degree of partial melting of pyrolite. Enriched source composition for Ce, Yb, and Dy from western São Miguel [*Beier et al.*, 2007]. Representative enriched alkaline component is sample SM18-8-97-5 from *Beier et al.* [2006]. Partition coefficients for Ti were calculated by *McKenzie and O'Nions* [1991] after analyses from *Irving and Frey* [1978], *Harte et al.* [1987], *Stolz and Davies* [1988], and *Galer and O'Nions* [1989]. Na partition coefficients are from *Leeman and Scheidegger* [1977] for olivine, from *Blundy et al.* [1995] for Cpx, from *Onuma et al.* [1968] for Opx, and from *Putirka* [1998] for Grt. The partition coefficients for Ce, Yb, and Dy are taken from *McKay* [1986] for olivine, *Kelemen et al.* [1993] and *Dick and Kelemen* [1991] for orthopyroxene, *Hart and Dunn* [1993] for clinopyroxene, and *Stosch* [1982] for spinel. The garnet partition coefficients are from *Johnson* [1994]. Primitive upper mantle from *McDonough and Sun* [1995].

sensitive to varying amounts of garnet, we will also use the even more incompatible trace elements to confirm the observations made by the La/Sm ratios. Our model suggests that the elevated Dy/Yb of the São Miguel and some Terceira lavas formed by partial melting under the influence of higher amounts of garnet (~15%) whereas the lower Dy/Yb of the western islands of Graciosa, some lavas from Terceira, and lavas from João de Castro (Figure 8c) formed by melting of mantle containing ~10% residual garnet and 5% spinel, i.e., at the transition from garnet to spinel stability field. It has to be noted though that the lavas from Terceira have both low and high Dy/Yb ratios (Figure 8c). To account for the possible presence of pyroxenite in the Azores mantle sources we assume equal amounts of pyroxenite and peridotite (see figure caption of Figure 8 for details) to ensure the best possible fit. While the REE ratios (Figures 8b and 8c) and the presence of recycled oceanic crust imply that pyroxenitic veins may be present beneath the Azores, the Al<sub>2</sub>O<sub>3</sub> contents at given SiO<sub>2</sub> contents (Figure 7) are too low to be solely explained by pyroxenite melting. The discrepancy between major and trace elements may possibly be best explained by reequilibration of the melts with the surrounding mantle leaving the trace elements relatively unaffected while the major elements are reequilibrated. If the modal source compositions

are largely similar then the magmas from São Miguel and João de Castro have formed by slightly smaller degrees of partial melting (1–2%) than the melts beneath the western Terceira Rift ranging from 2 to 4% (inferred from the REE ratios in Figures 8b and 8c).

[21] We conclude that the magmas generated beneath Graciosa, Terceira, and João de Castro have been generated in the garnet/spinel transition zone in contrast to the magmas beneath São Miguel which formed in the garnet stability field only. The degrees of partial melting along the axis indicate that, in general, the eastern volcanoes (São Miguel particularly) have slightly lower degrees of partial melting than the western islands (Terceira and Graciosa); however, these changes are small. Along the oblique spreading segments of the SWIR decreasing degrees of partial melting have been associated with an increasing lithospheric thickness [*Standish et al.*, 2008].

[22] The increasing depth of partial melting toward São Miguel is consistent with an increasing lithospheric thickness with distance from the MAR [e.g., *Cazenave*, 1984]. The age of the lithosphere in the vicinity of the Princessa Alice bank has been estimated to be 10 Ma [*Cannat et al.*, 1999] whereas the lithosphere south of the island of Terceira may be 36 Ma old and the junction

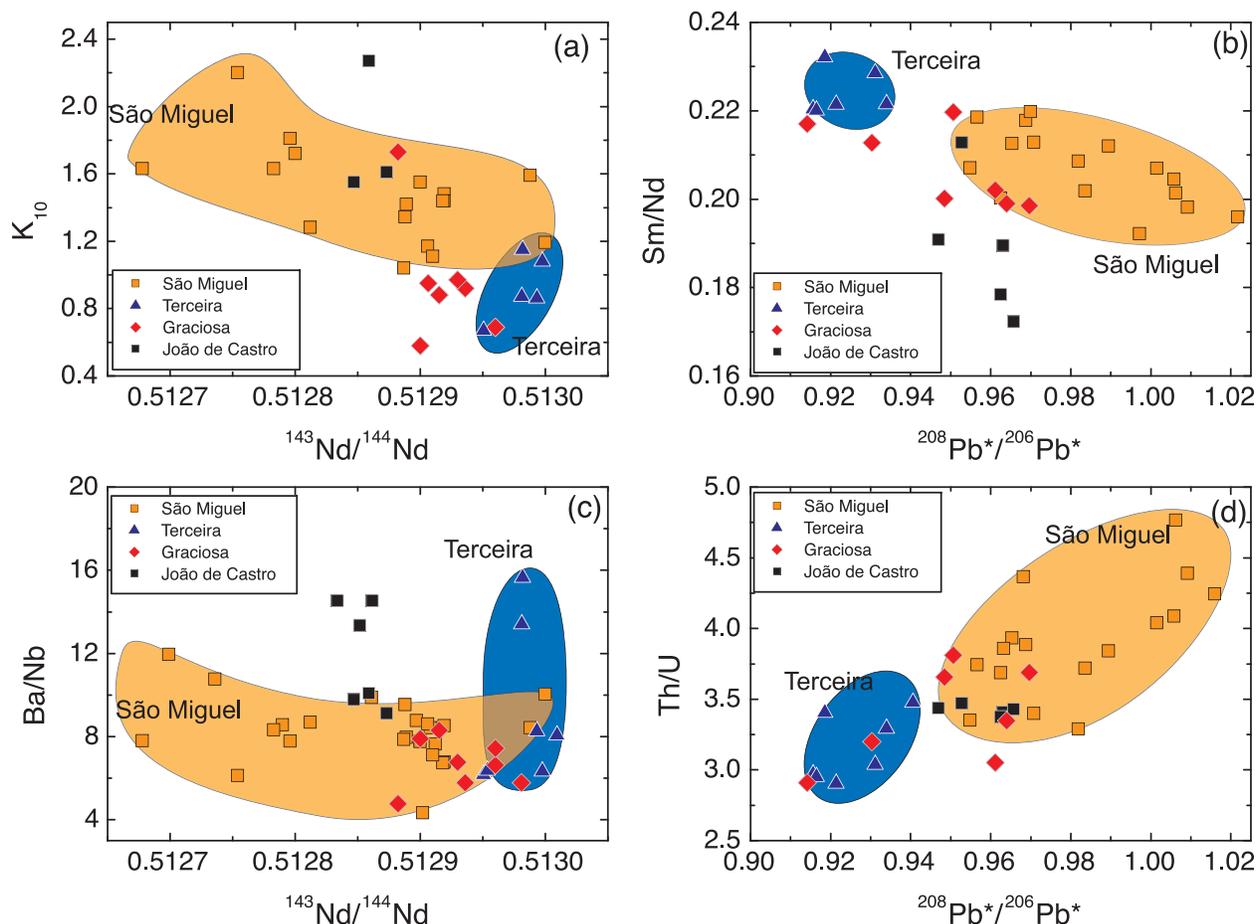
between the Terceira Rift and the lateral Gloria fault east of the Azores plateau has been estimated to be 53 Ma [Searle, 1980]. Based on these ages the thickness of the oceanic lithosphere can be calculated yielding thicknesses of 36, 68, and >81 km, respectively [Stein and Stein, 1992]. However, the formation of alkali basaltic melts beneath the Terceira Rift implies relatively higher pressures of melting than on other ultraslow spreading axes where tholeiitic basalts form by low degree melting at shallow depth, e.g., at the SWIR (N-MORB to E-MORB ~7–14% degree of partial melting depending on mantle lithology, [Standish *et al.*, 2008]) or the Arctic Gakkel Ridge (~5%) [Hellebrand and Snow, 2003]. Higher pressures of melting at the Terceira Rift compared to the other ultraslow spreading axes are a result of young (<6 Ma) rifting of lithosphere formed at the MAR. This lithosphere is probably still largely intact and not yet thinned by the very slow extension (~4 km/Ma). Airy compensation models suggest that the lithosphere beneath the Azores is indeed thickened rather than thinned by mantle melting [Grevemeyer, 1999]. Thus, although the mantle beneath the Azores may be either relatively hot and/or volatile-rich material and begins to melt at great depth, the lithospheric lid above the melting zone prevents increased degrees of partial melting.

## 5.2. Mantle Sources and Mixing Along the Terceira Rift

[23] The linear Pb isotope arrays of Graciosa, João de Castro and Terceira either reflect true isochrons or binary mixing lines (Figures 6d and 6e). The isotopic array of São Miguel has been modeled to represent a mixing array [Beier *et al.*, 2007]. The results from reduced chi-square regression lines of the triple spike data indicate a slope of  $0.087 \pm 0.005$  for the João de Castro lavas which is comparable to the slope at Terceira ( $0.084 \pm 0.004$ ) in the  $^{207}\text{Pb}/^{204}\text{Pb}$  versus  $^{206}\text{Pb}/^{204}\text{Pb}$  space (Figure 6d). If these slopes represent isochrons the corresponding ages would be  $1.37 \pm 0.12$  and  $1.31 \pm 0.09$  Ga, respectively. A calculated regression line from Graciosa gives a slope of  $0.009 \pm 0.006$ , representing a zero-age isochron slope giving evidence that the Graciosa samples most likely represent binary mixing rather than an isochron. Whether the  $^{207}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$  regression lines of João de Castro and Terceira represent mantle isochrons can be tested using the information derived from  $^{208}\text{Pb}/^{204}\text{Pb}$  versus  $^{206}\text{Pb}/^{204}\text{Pb}$  relationship on the source  $\kappa$  ( $^{232}\text{Th}/^{238}\text{U}$ ) value  $\kappa_{\text{ISO}}$ . Hence, the inferred source

$\kappa$  ( $^{232}\text{Th}/^{238}\text{U}$ ) values from the measured Th and U ( $\kappa_{\text{TE}}$ ) concentrations and from the  $^{208}\text{Pb}/^{206}\text{Pb}$  ratios ( $\kappa_{\text{ISO}}$ ) should be correlated, i.e., the  $\kappa_{\text{TE}}$  should be equal or higher than the  $\kappa_{\text{ISO}}$  if due to melting [Beattie, 1993]. In contrast to the prediction the  $\kappa_{\text{TE}}$  values of João de Castro and Terceira are considerably lower than the  $\kappa_{\text{ISO}}$  which is inconsistent with the expected Th/U fractionation during melting [Abouchami *et al.*, 2000a]. Therefore, we suggest that the linear arrays in the  $^{207}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$  and  $^{208}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$  isotope systems do not represent mantle isochrons as a result of mantle melting but more likely reflect mixing lines between two distinct mantle sources on each array.

[24] Two different magma groups along the Terceira Rift can be distinguished based on their  $\text{K}_2\text{O}$  contents with Terceira and Graciosa lavas having lower  $\text{K}_2\text{O}$  than those from São Miguel and João de Castro at a given MgO content (Figure 3). These two groups can be also distinguished in terms of Nb/Zr and Th/U (Figure 5) suggesting that the eastern volcanoes of the Terceira Rift have more enriched mantle sources than the western volcanoes. A broad negative correlation between the  $^{143}\text{Nd}/^{144}\text{Nd}$  isotope ratios and the K concentrations (and ratios such as K/Ti, K/U) confirms that these variations are mainly the result of mantle source signatures. However, each single mixing array (Graciosa, São Miguel, João de Castro, and Terceira) in the Pb-Pb isotope spaces gives evidence for two component mixing between discrete mantle end-members (dashed circles in Figures 6d and 6e). All mixing arrays except João de Castro appear to converge to a common Azores composition that is reflected by the unradiogenic Graciosa samples. The source of the João de Castro seamount lavas mixes with a relatively unradiogenic mantle source different from basalts from the Mid-Atlantic Ridge as indicated by their lower  $^{207}\text{Pb}/^{204}\text{Pb}$  at a given  $^{206}\text{Pb}/^{204}\text{Pb}$  and less radiogenic  $^{143}\text{Nd}/^{144}\text{Nd}$  (Figures 6d–6f). At Graciosa and Terceira the common end-member is characterized by radiogenic Pb isotope ratios and high Nd isotope ratios, in that sense comparable to the HIMU ocean island lavas (high  $\mu$  = high  $^{238}\text{U}/^{204}\text{Pb}$ ) such as St. Helena [e.g., Zindler and Hart, 1986]. The linear array at Graciosa covers a comparable  $^{206}\text{Pb}/^{204}\text{Pb}$  range as the Terceira lavas but the shallower Graciosa array results from slightly more radiogenic  $^{208}\text{Pb}/^{204}\text{Pb}$  and  $^{207}\text{Pb}/^{204}\text{Pb}$  ratios which is also reflected in higher  $^{208}\text{Pb}^*/^{206}\text{Pb}^*$  ratios at Graciosa than at Terceira (Figure 6b). The radiogenic Pb end-member compositions of Terceira and Graciosa suggest interac-



**Figure 9.** Mantle source systematics along the Terceira Rift showing variations in (a and c) fractionation-corrected (10 wt. % MgO) K-contents, Sm/Nd, Ba/Nb, and Th/U ratios versus Nd isotope ratios and (b and d)  $^{208}\text{Pb}^*/^{206}\text{Pb}^*$  ratios. São Miguel and Terceira are shown as fields for clarity.

tion between the two systems but their relatively unradiogenic Pb end-member compositions indicate that the two unradiogenic mantle sources are slightly distinct, i.e., the common source of the Terceira Rift lavas shows some variation.

[25] The isotope ratios of Sr and Nd are broadly correlated with the La/Sm, Rb/Nb, Ba/Rb, Th/Nb, and Th/U ratios (not shown), i.e., generally the lavas with the highest  $^{143}\text{Nd}/^{144}\text{Nd}$  from Terceira also have the highest Sm/Nd but lavas from São Miguel have both variable Sm/Nd and  $^{143}\text{Nd}/^{144}\text{Nd}$ , suggesting that partial melting processes probably affected the Sm/Nd ratio. Although the linear arrays of the Graciosa, Terceira, and western São Miguel samples meet at similar ratios in the isotope spaces, the combined trace element (K, Th/U, Ba/Nb, Nb/Zr) and Sr-Nd-Pb isotope systematics (Figure 9) show that each volcano along the axis contains its own isolated mantle source implying significant heterogeneity in

the mantle beneath the Azores on a scale of about 100 km. The limited mixing between different sources has also been observed on a much smaller scale of about 20 km between the western and eastern volcanoes at São Miguel [Beier *et al.*, 2007; Haase and Beier, 2003].

[26] Our new data reveal that rather than the three end-members in the Azores lavas suggested by previous authors [e.g., Moreira *et al.*, 1999b], at least four end-members are required to explain the isotopic variation of the Terceira Rift lavas. The three end-members were previously believed to be MORB, a plume component represented by Terceira lavas, and an enriched mantle component represented by lavas from eastern São Miguel [e.g., Beier *et al.*, 2007; Elliott *et al.*, 2007]. The isotopic variation observed in each volcanic system indicates that at least two sources are present beneath each structure. Some of the lavas from Graciosa with  $^{87}\text{Sr}/^{86}\text{Sr}$  of  $\sim 0.7035$ ,  $^{143}\text{Nd}/^{144}\text{Nd}$

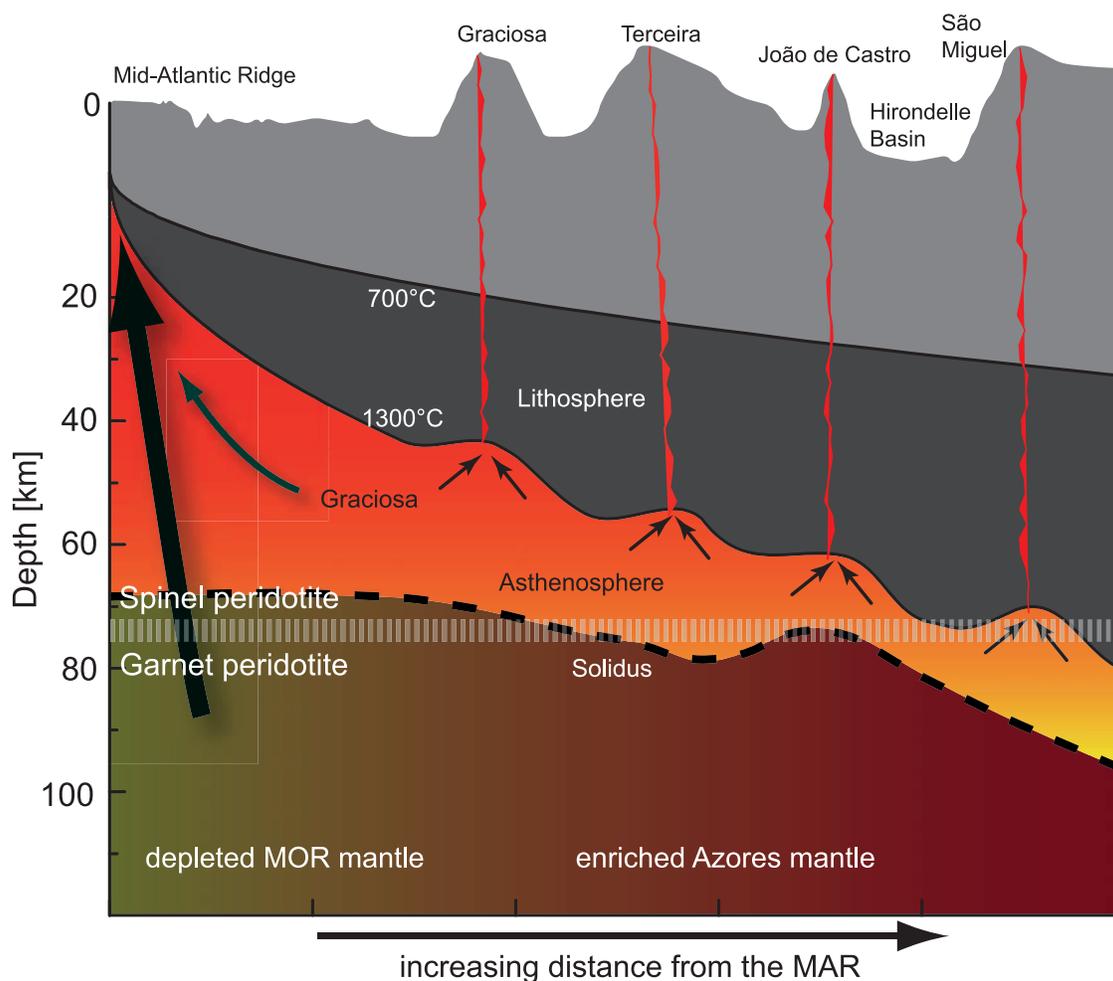
of  $\sim 0.5129$ , and  $^{206}\text{Pb}/^{204}\text{Pb}$  of  $\sim 19.4$  to  $19.6$  appear to represent a possible end-member common in all other volcanic systems as well as MORB and thus could be an average “Azores plume component.” Interestingly, this source also resembles the composition suggested for the FOZO or “C” component which is believed to be abundant in the mantle [Hanan and Graham, 1996; Hart et al., 1992; Stracke et al., 2005]. While the origin of this common component in the Azores is beyond the scope of this work, the data clearly require the Azores mantle anomaly to be heterogeneous on a scale of tens of kilometers similar to findings in mantle plumes like Hawaii [Abouchami et al., 2000b, Eisele et al., 2003]. The low degrees of partial melting ( $<5\%$ ) of the Terceira Rift magmas may be responsible for the preferred sampling of small heterogeneities as a result of the increased viscosity of low degree melts at relatively lower temperatures compared to higher degree melts at higher temperatures [Bourdon et al., 2006; Scarfe and Cronin, 1986].

### 5.3. Influence of the Terceira Rift Mantle on Mid-Atlantic Ridge Basalts

[27] The existence of a geochemical anomaly in the lavas erupting at the Mid-Atlantic Ridge close to the Azores is long known [e.g., Bougault and Treuil, 1980; Schilling, 1975; White et al., 1975; White et al., 1976] and has been attributed to the influx of enriched and hot mantle that originates from a deep mantle plume beneath the Azores. Relatively primitive He isotope ratios ( $^4\text{He}/^3\text{He} \sim 64,000$ ) from Terceira and the adjacent MAR have been interpreted to result of plume-ridge interactions between the plume center located beneath Terceira and the MAR [Moreira et al., 1999a]. However, we have shown that the lavas from Terceira and Graciosa resemble each other in  $^{206}\text{Pb}/^{204}\text{Pb}$  compositions but significant differences exist in  $^{207}\text{Pb}/^{204}\text{Pb}$ ,  $^{208}\text{Pb}/^{204}\text{Pb}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios (Figure 6). Thus, the Terceira lavas have too high  $^{143}\text{Nd}/^{144}\text{Nd}$ , and too low  $^{207}\text{Pb}/^{204}\text{Pb}$  to represent the mixing end-member for Mid-Ocean Ridge Basalts (MORB) from the adjacent MAR (Figure 6d–6f) and rather, the material influencing the spreading axis has a composition comparable to the Graciosa mantle source. In fact, despite slight offsets, the unradiogenic Graciosa lavas lie closest to the point of convergence of the trends of all Terceira Rift and MAR lavas suggesting that they could possibly represent a common mantle end-member inherent in all Terceira Rift magmas as well as MORB close to the Azores. The

$^{143}\text{Nd}/^{144}\text{Nd}$  and  $^{206}\text{Pb}/^{204}\text{Pb}$  similarities between the MAR and Graciosa suggest mixing of the Graciosa mantle source into the MAR mantle, and thus indicates a mantle flux from Graciosa toward the MAR. This seems also evident from the bathymetry in Figure 1, i.e., there is no bathymetrical evidence for active volcanism between the dredged seamount west of Graciosa and the MAR. If the melts are focused along the Terceira Rift (see discussion below) then they will likely be also focused into the MAR and/or toward Graciosa, respectively. As no He isotope measurements are known from Graciosa and the origin of the primitive He isotopic composition are still a matter of debate, we suggest that the enriched MAR signature is a result of mantle flux from Graciosa to the MAR. We speculate that Graciosa’s lavas with the lowest Pb isotope ratios will also have the most primitive He isotope ratios; however, this has to be tested.

[28] Because Graciosa is the island closest to the MAR its influence on the spreading axis is not surprising. However, we note that no plume center can be defined in the Azores because five of the islands east of the MAR have erupted in historical times (São Miguel, Terceira, Graciosa, Faial, Pico) and with one exception all islands (Santa Maria) east but also west of the MAR are very young (2–0.1 Ma) [Abdel Monem et al., 1975; Calvert et al., 2006; Féraud et al., 1980; Féraud et al., 1981; Johnson et al., 1998; Madeira et al., 1995; McKee and Moore, 1992; Snyder et al., 2007], implying that there is no age progression. Because the young volcanism along the Terceira Rift is a result of extension it appears possible that the center of an actively ascending deep mantle plume lies beneath Graciosa-São Jorge-Faial-Pico, the islands closest to the MAR which could also explain the volcanism of the latter two islands. However, such a model cannot explain why two very young volcanic islands Corvo and Flores formed west of the MAR because the relatively deep and thin crust of the MAR implies that there is most likely no plume rising beneath the MAR [Cannat et al., 1999]. Some tomographic models suggest that there is no connection of the Azores mantle anomaly into the deep mantle [Courtilot et al., 2003; Ritsema and Allen, 2003] and thus this material may represent old material from a plume head arriving beneath the lithosphere some 10 to 5 Ma ago but which is no longer active, hence there is no evidence for an ascending tail of a mantle plume [Ritsema and Allen, 2003; Silveira et al., 2006]. In this case the fossil plume head material may flow into the MAR spreading axis and the Terceira Rift



**Figure 10.** Sketch illustrating the processes dominating the Terceira Rift evolution. The melting regions of Graciosa, João de Castro, and Terceira are situated within the spinel/garnet transition zone. The São Miguel melts are generated within the garnet stability field. The mixing relationships described in the main text suggest only very limited mixing between each islands source. The occurrence of a typical avolcanic-volcanic segmentation pattern has its origin in focused magmatism comparable to the SWIR. The isotope systematics suggest a mantle flux from Graciosa toward the adjoining MAR. Not to scale.

because it is less viscous and less dense than the upper mantle material and because of the suction of the diverging plates. The presence of 5–10 Ma-old mantle plume head material beneath the whole Azores Platform and west of the MAR may also explain the young volcanism of the two islands west of the MAR by passive melting of enriched material within the mantle anomaly.

#### 5.4. Comparison to Ultraslow Ridges and Continental Rifts

[29] The Terceira Rift is characterized by an avolcanic-volcanic segmentation pattern (Figure 1b) that has been also observed along the oceanic Southwest Indian (SWIR) [Sauter et al., 2004a] and Arctic Gakkel Ridges [Michael et al., 2003]

and the continental Ethiopian rift [Corti, 2008]. These ridges are defined by large magmatic segments with a thicker lithosphere bordered by deeper avolcanic basins with a relatively thin lithosphere. The avolcanic segments at the SWIR are also characterized by either the presence of sediments or peridotitic rocks, a feature that we can only suspect in the Azores. The magmatic segments of the oceanic ridges typically consist of large volcanic structures, which, in the case of the Terceira Rift, form islands. As Vogt and Jung [2004] pointed out, the extreme topographic variation at the Terceira Rift is most likely the result of the presence of anomalous mantle beneath the Azores which led to the formation of a shallow oceanic plateau which is then divided by the

Terceira Rift. A comparable segmentation pattern is also observed along the Mid-Atlantic Ridge from 25°N to 48°N [Magde and Sparks, 1997], where neither temperature differences nor viscosity changes have been found to affect the segmentation pattern. Instead, the ultraslow spreading may lead to an increased cooling beneath the avolcanic, hence cooler sections, forcing melts to laterally move along ridge, which produces the observed segmentation pattern [Magde and Sparks, 1997].

[30] The degree of partial melting beneath the Terceira Rift is comparable or lower than at other ultraslow ridges but melting occurs much deeper, producing alkaline magmas in the Azores in contrast to mostly tholeiitic basalts at the Gakkel Ridge and the SWIR [e.g., Michael et al., 2003; Standish et al., 2008]. The process of conductive cooling becomes important at ultraslow spreading rates (<20 mm/a) leading to a decrease of melting at shallow depth and lower degrees of partial melting [Reid and Jackson, 1981]. As pointed out above, the smaller degrees of partial melting at the Terceira Rift are probably a result of the thicker surrounding lithosphere and the relatively recent transition from a transform fault to spreading (<6 Ma from <sup>40</sup>Ar/<sup>39</sup>Ar age determinations) which lead to a thick lithospheric lid and deep melting. In contrast, spreading has been established for a long time at the Gakkel Ridge and SWIR leading to a thin lithosphere and shallow melting. The obliquity of ultraslow spreading rifts also has an important impact onto melting processes; i.e., an increasing obliquity leads to decreasing effective spreading rates, lower upwelling velocities, and smaller melt fractions [Okino et al., 2002; Standish et al., 2008]. The spreading rate along the Terceira Rift increases from the east (3.7 mm/a) to the western edge (4.5 mm/a) and obliquity decreases from São Miguel (61°) to Graciosa (40°) [Vogt and Jung, 2004]. We suggest that the smaller degrees of partial melting at the Terceira Rift compared to other ultraslow spreading rifts are a result of a combination of a thicker, relatively older lithosphere with a very slow spreading rate and a higher obliquity. The formation of islands along the axis, despite the smaller degrees of partial melting, results from the large relief amplitude of 2–4 km [Vogt and Jung, 2004] on a shallow plateau and is most likely the result of the presence of enriched and possibly hot mantle plume material beneath the rifted Azores Plateau. This material generates larger volumes of melts over longer time periods than the depleted mantle present beneath other ultraslow spreading axes. The distinct variations of melting degrees between the Terceira

Rift volcanoes give either evidence for a limited melt production beneath the bathymetric basins or highly focused magmatism along-axis like it has been proposed for the SWIR [Sauter et al., 2004b; Standish et al., 2008] and the Ethiopian rift [Corti, 2008]. The occurrence of focused melts may also lead to increased melt volumes in the volcanic segments comparable to the oblique segment of the SWIR [Standish et al., 2008]. As a result of the Oceanic Plateau situated in the vicinity of the spreading axis the Terceira Rift may indeed share many similarities with continental rifts (e.g., East African Rift system) such as low degrees of partial melting, oblique rifting, and segmentation patterns.

[31] The presence of well defined, distinct mantle sources beneath each island/seamount inferred from trace elements and Nd and Pb isotopes suggests that focused magmatism occurs along the Terceira Rift as has been proposed from geophysical observations along the SWIR and Ethiopian rift. The focusing most likely occurs in distinct mantle diapirs which underlie each volcanic system but lack beneath the avolcanic basins. Within the diapirs the melts are focused to the surface and may mix within each diapir but not among different diapiric structures [Crane, 1985; Okino et al., 2002]. The diapiric melts that have lower densities and viscosities than the surrounding mantle move toward regularly spaced established gravitational instabilities of the partially molten mantle by porous flow avoiding mixing between the segments [Lin et al., 1990; Schouten et al., 1985; Whitehead et al., 1984]. The initial establishment of gravitational melt instabilities is mainly controlled by the continuity and thickness of the underlying melt layer [Crane, 1985; Michael et al., 2003], whereas the spacing of magmatic centers is mainly controlled by the effective spreading rate [Schouten et al., 1985]. Slower spreading rates are correlated to smaller distances between the magmatic segments consistent with the relatively small-scale segmentation pattern (10–30 km) observed along the Terceira Rift and other slow spreading rifts (e.g., Gakkel Ridge, SWIR) compared to faster spreading ridges such as the MAR (50–80 km). The segmentation pattern along the Terceira Rift is also comparable to propagating rifts such as the southern Iceland rift [Tentler, 2005] or continental rifts such as the East African Rift system [Wright et al., 2006]. While a magmatic origin of the segmentation pattern beneath continental rifts is a matter of active debate (i.e., solely tectonic influences versus magma intrusions [Corti, 2008 2347; Wright et al., 2006]) the occurrence of segmenta-

tion patterns in oceanic rifts will most likely be controlled by the availability of melts [Tentler, 2005]. For the specific case of Iceland, the obliquity and segmentation pattern are likely to be controlled by the emplacement of dykes.

[32] Summarizing, the Terceira Rift represents an ultraslow spreading rift which consists of both continental (e.g., spreading rate, lithospheric thickness) and oceanic features (e.g., melt availability, segmentation pattern/focused magmatism). The Terceira Rift thus represents the first known ultraslow spreading rift within old (>10 Ma) oceanic lithosphere. It differs from other oceanic rifts mainly as a result of the lithospheric age and the presence of a melting anomaly and thus is a unique example of oceanic rifting. On the basis of the occurrence of both continental and oceanic features, one could speculate that the Terceira Rift may represent the earliest stages of a rift system and may later develop into an oceanic spreading center.

## 6. Conclusions

[33] We conclude that the melting depth along the Terceira Rift does not vary systematically beneath Graciosa, Terceira, and João de Castro, where melts are generated in the spinel/garnet transition zone (Figure 10). However, deeper melting in the garnet stability field is observed at São Miguel, the island most distant from the MAR and possibly with the thickest lithosphere. The degrees of partial melting are smaller at São Miguel and João de Castro compared to the other volcanic systems along the axis. Compared to other very slow spreading axes the generally low degrees of partial melting at the Terceira Rift probably result from a combination of a thick lithospheric lid and a relatively young (<1 Ma) and ultraslow spreading movement combined with a high obliquity (40° to 61°). Although the Terceira Rift has lower degrees of partial melting compared to the SWIR or Gakkell Ridge, the presence of an anomalous upper mantle generates enough melts over relatively long time periods to enable subaerial volcanism. The incompatible trace element ratios (e.g., Th/U, Nb/Zr) and combined Sr-Nd-Pb isotopes suggest that every volcanic system along the Terceira Rift is situated on a single binary mixing trend without evidence of mixing in between. The limited mixing is attributed to the presence of geochemical boundaries and the occurrence of focused magma transport in separate mantle diapirs as has been also observed on other ultraslow spreading rifts. The

avolcanic–volcanic segmentation pattern along the Terceira Rift is comparable to the structures observed along the SWIR. This implies that these segmentation patterns are characteristic features of ultraslow spreading rifts with spreading rates <14 mm/a. The Sr, Nd, and Pb isotope composition of Graciosa lavas and their relation to the isotope trend in MORB form the adjacent MAR suggest a mantle flux from Graciosa toward the MAR causing the observed enriched MORB compositions along the MAR.

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