

1 **Supporting information**

2 **“Megathrust reflectivity reveals**
3 **the updip limit of the 2014 Iquique earthquake rupture”**

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16 **1. Seismic Processing**

17 Seismic data processing was conducted with the Schlumberger OMEGA2 software.
18 Here we present the processing sequences and the corresponding results in terms of seismic
19 line MC04. The same processing strategy is used for all profiles to eliminate the possibility that
20 differences in reflectivity are due to differences in processing. An overview of the processing
21 scheme is shown below (Suppl. Fig. 1). The main seismic processing included four steps:

22

23 **A. Step1: Processing Geometry Implementation**

24 In this step, the seismic data are exposed to a series of pre-processing modules of the
25 Schlumberger OMEGA2 software, including sail extract module, geometry crooked module,
26 geometry update module and grid define module, aiming to implement the processing
27 geometry. In particular, the sail extract module is used to merge the geometry survey file
28 information with the seismic traces. The geometry crooked module analyses the location of
29 detectors, sources, and midpoints and projects them onto a smooth 2-D crooked common-mid-
30 point (CMP) profile. Subsequently, the geometry update module and the grid define module
31 were used to update the seismic trace header literals to the new processing grid. These pre-
32 processing steps prepare the seismic data set for further analysis. Seismic shooting was
33 conducted at an interval of 125 m to avoid interference from the previous shot on large offset
34 OBS data. Due to the large shot interval, the original CMP gathers are imaged with prominent
35 spatial aliasing. The aliasing has serious effects on the performance of multichannel data
36 analysis processes such as f-k filtering. Because of spatial aliasing, these processes can perceive
37 events with steep dips at high frequencies as different from what they are and, hence, do not
38 treat them properly.

39

40 **B. Step2: Increase Vertical Resolution and General Noise Cleaning**

41 In this step, four modules were used to increase the vertical resolution and clean the
42 general noise. In particular, the deconvolution module is applied to compress the basic wavelet,
43 thus increasing temporal resolution. The direct wave cleaning module was employed in shallow
44 seafloor areas to enhance near-surface wide-angle reflections. Attenuation of random noise
45 (RNA) was applied by a predictive deconvolution in the f_x -domain. The anomalous amplitude
46 attenuation (AAA) removed high amplitude events such as marine swell, rig and ship noise by
47 transforming the seismic data into frequency bands and applying a spatial median filter. These
48 pre-processing steps yield CMP gathers (Suppl. Fig. 2a) which will be interpolated in the next
49 step.

50

51 **C. Step3: Shot Interpolation**

52 To avoid prominent spatial aliasing between neighboring shots resulting from the large
53 shot interval, shot-ordered seismic data were interpolated three times from 125 m to 15.625 m
54 by f-k trace interpolation. As the result is not a multiple of the receiver interval, an irregular
55 interpolator module was used to interpolate it to 12.5 m, where aliased components have been
56 eliminated (Suppl. Fig. 2b).

57

58 **D. Step 4: Multiple attenuation**

59 Prior to multiple attenuation, the plate boundary reflection is severely obscured by
60 multiples (Suppl. Fig. 3a). The multiples were predicted by surface related multiple elimination
61 (SRME) using wavefield inversion based on the Kirchhoff integral and subtracted from the raw
62 data after shot interpolations. After the multiple attenuation, we applied the f-k filtering to
63 remove the high frequencies to enhance the image (Suppl. Fig. 3b). The plate boundary
64 reflection is clearly observed in the travel time range where the first multiple occurs and we can
65 now trace the plate boundary interface further landward on the time section of seismic line
66 MC04 (Suppl. Fig. 4). The extensive data volume generated by the shot interpolations led to
67 long calculation times in the pre-stack time domain and the pre-stack depth domain, so we
68 reconstructed the offsets of CMP gathers with an increment of 100 m. After the reconstruction
69 of offset distances, a preliminary v_p model based on an unpublished 3D v_p model from ocean
70 bottom seismometers (OBS) during cruise MGL1610 of the *R/V Marcus G. Langseth* in 2016
71 was applied for the pre-stack depth migration. Referring to the OBS velocity field, we
72 calculated the velocity gradient starting from the seafloor to yield an initial velocity model
73 (Suppl. Fig. 5) for application to the multichannel seismic profiles. Although this initial velocity
74 model is not sufficiently confining the shallow depth (< 8 km below seafloor) and may produce
75 an inaccuracy of 1-2 km in the depth of the plate boundary compared to an accurate velocity
76 model, it does not affect the spatial distribution of the reflectivity. In the meanwhile, this initial
77 velocity does not enable to carry out amplitude analysis at shallow depth. However, our main
78 focus is on the reflectivity pattern and the lateral coherence and continuity of the plate boundary

79 reflectivity. The pre-stack depth migrated seismic images are not significantly affected by
80 seafloor multiples anymore and resolve sub-seafloor structures in the upper plate and
81 subducting lower plate at high resolution (Suppl. Fig. 6).

82

83 **E. Step 5: Amplitude calibration**

84 To quantify the amplitudes of the seismic sections, the reflection coefficient was
85 estimated based on the ratio of the seafloor reflection to the seafloor multiple reflection. Due to
86 interference of several reflector elements resulting in inverse and mix phased signals especially
87 in the crustal overburden the envelope was calculated representing the absolute reflection
88 strength (Suppl. Fig. 7). The plate boundary shows a unique reflection strength of 0.005 to
89 0.0075 in a depth range between 15 to 35 km, whereas the internal crustal reflector elements
90 show a higher variability ranging from 0.0075 to 0.025 at a depth range of 5 to 15 km.

91

92 **2. Thermal Model Setup**

93 It has long been recognized that the rupture zone of subduction zone megathrust
94 earthquakes is at least partially controlled by the thermal state of the fault zone^{1,2}. Analytical
95 models reveal that the geometry of the subduction zone, the thermal state of the incoming
96 subducting plate and the shear or frictional heating along the megathrust are critical parameters
97 controlling megathrust temperatures^{3,4}, which in turn, define the seaward and landward limit of
98 large subduction earthquakes rupture zones. The seaward or updip limit is generally associated
99 with temperatures of 100°-150°C, marking the smectite to chlorite transition² or a suite of
100 diagenetic reactions and release of water from underthrust sediments⁵ and/or basement⁶. The
101 landward or downdip limit is assumed to be associated with a critical temperature of 350°-
102 400°C, which marks the transition from stick-slip to stable sliding at the onset of quartz and
103 feldspar plasticity of continental crustal rocks⁷.

104

105 The geometry of the subduction zone is readily known from geophysical data or the
106 hypocentral depth of large megathrust earthquakes and the basal heat flow is defined by the age

107 of the incoming oceanic plate. Most thermal parameters of subduction zones show little
 108 variation along the Pacific Ring of Fire and are well established, especially for Chile^{8,9}. Based
 109 on these data, most subduction zones have been studied using two-dimensional thermal models.
 110 Here, we use two different approaches. First, we calculate the thermal state along the plate
 111 interface or megathrust fault using analytical solutions based on the formalism of ref.⁴ and
 112 second, we consider a 2-dimensional thermal model of ref.¹⁰ for northern Chile, incorporating
 113 corner flow in the mantle wedge.

114 Analytical expressions, which relate surface heat flux to temperature, geometrical
 115 constraints, and shear stress, provide an efficient approach to study the thermal state of the
 116 megathrust fault and are discussed in detail by refs.^{3,4,11}. We follow the approach of ref.⁴ and
 117 first calculate the temperature T_f as a function of the depth z on the interplate fault zone as:

$$118 \quad (1) \quad T(z) = K_m T_0 z / SK_s [\pi \kappa (t_0 + t_s)]^{1/2}$$

119 where $S(z) = 1 + bK_m [(V_n z \sin \delta) / \kappa]^{1/2} / K_s$. K_m ($3.3 \text{ Wm}^{-1}\text{K}^{-1}$) and K_s ($2.55 \text{ Wm}^{-1}\text{K}^{-1}$) are the
 120 mantle and forearc thermal conductivity, respectively. T_0 is the asthenospheric mantle
 121 temperature (1300°C) and κ is the thermal diffusivity ($10^{-6} \text{ m}^2 \text{ s}^{-1}$). t_0 is the average age of the
 122 subducting plate (50 Myr^{12}), whereas t_s is the time it takes the lithosphere to subduct to a depth
 123 z . V_n is the convergence rate normal to the subduction ($\sim 70 \text{ km/Myr}$), δ is the dip angle of
 124 subduction, and $b(\pi^{-1/2})$ is a factor that depends on the specific geometry¹¹.

125

126 The dip angle of subduction was taken directly from the seismic reflection images (Fig.
 127 2), and the time t_s is computed by dividing the integrated downdip length of the fault surface by
 128 V_n ¹³. We neglect the effect of the horizontal heat flow. To calculate the radiogenic heat
 129 production T_r in the forearc crust we used the following expression¹³:

$$130 \quad (2) \quad T_r(z) = A_r z^2 / (2K_s S(z))$$

131 where A_r is the radiogenic heat production rate (10^{-6} Wm^{-3})¹⁴. Radiogenic heat production adds
 132 $0\text{-}45^\circ\text{C}$ to the fault temperature from the trench axis up to the downdip limit.

133

134 Further, we include in our model frictional shear heating $T_{sh}(z)$ on the thermal field by
135 using:

$$136 \quad (3) T_{sh}(z) = \tau(z)V_t z/(K_s S(z))$$

137 where $\tau(z)$ is the shear stress on the fault and V_t is the total slip rate^{4,11}. $\tau(z)$ on a gently dipping
138 fault at shallow depth is approximately

$$139 \quad (4) \tau(z) = \mu(\sigma_n(z) - p(z))$$

140 where μ is the friction coefficient, σ_n is the normal stress applied on the fault plane
141 (approximately the overburden pressure), and p is the pore fluid pressure. Following ref.¹³, we
142 use $\mu = 0.85$, $\sigma_n(z) = \rho g z$ and $p(z) = 0.95 \sigma_n(z)$ with acceleration of gravity $g = 9.8 \text{ ms}^{-2}$ and the
143 average crustal density $\rho = 2500 \text{ kgm}^{-3}$.

144

145 The final predicted temperature on the fault plate boundary $T_f(z)$ is the sum of Eqs. (1)-(3)
146 (i.e., $T_f(z) = T(z) + T_r(z) + T_{sh}(z)$). Fig. 2 shows the estimated temperature values for $T_f(z)$ along
147 our seismic reflection lines.

148

149 We compare our model to the numerical model of ref.¹⁰. The geometry of this model is
150 based on a suite of geophysical data¹⁵ and thermal parameters were rated against a number of
151 observed features, including the maximum depth of subduction thrust earthquakes and observed
152 heat flow. Interestingly, the maximum depth of seismic faulting of megathrust earthquakes in
153 northern Chile occurs at 40-50 km^{16,17}, suggesting that temperatures of 350°-400°C are reached
154 at ~40-50 km, too. To mimic this feature, ref.¹⁰ had to introduce a considerable amount of shear
155 heating, in the order of $\tau = 33 \text{ MPa}$ to $\tau = 67 \text{ MPa}$, with the upper limit providing a better fit to
156 the data. The predictions from the $\tau = 67 \text{ MPa}$ model mimic the prediction of our preferred
157 analytic solution down to a depth of approx. 30 km. At greater depth, the models differ with
158 the ref.¹⁰ model showing somewhat lower temperatures. The observed differences may stem
159 from the effects of the asthenospheric corner flow incorporated into the numerical model and a

160 change in dip angle, which is not considered in the analytic model. We also compare our model
161 with other thermal models from the Northern Chilean margin. A comparison of all the models
162 for the temperature along the plate interface is shown in Suppl. Fig. 9, including the thermal
163 model with shear stress $\tau=67$ MPa¹⁰ as the black dashed line, the thermal model with frictional
164 heating we used in the main text as the red dashed line, the thermal model with shear stress
165 $\tau=33$ MPa¹⁰ as the orange dashed line, the thermal model of ref.¹⁸ as the magenta dashed line,
166 and a thermal model we built without the frictional heating as the green dashed line. Please note
167 that for the study of the up-limit limit at shallow depth (~ 15 km) and its correlation to the
168 reflectivity pattern of the seismic data, the model we established with frictional heating and the
169 model with shear stress $\tau=67$ MPa of ref.¹⁰ show consistent features.

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171 As previously mentioned in the main text, the established new thermal model uses a
172 friction coefficient $\mu=0.85$ and pore fluid pressure $\lambda=0.95$ ¹³. Since these values vary in each
173 tectonic setting, we applied different μ_b in the new analytical thermal model in Suppl. Fig. 10.
174 The effective coefficient of basal friction μ_b depends on both the friction coefficient μ and pore
175 fluid pressure λ along the fault zone: $\mu_b = \mu(1-\lambda)$ ¹⁹. Based on this formula, the μ_b of the model
176 in our main text is 0.0425, which is shown as the red dashed line. We applied a range of μ_b from
177 0.03-0.13, consistent with the global thermal measurement²⁰. In this range, the predictions depth
178 of the analytical model (< 20 km) is close to the downdip limit of the megathrust reflectivity
179 (~ 15 km) observed from MCS images at the upper threshold of the clay dehydration
180 temperature of 150°C (Suppl. Fig. 10). Moreover, the thermal model and reflectivity show a
181 better spatial matching with a higher effective coefficient of basal friction μ_b value.

182

183 An interesting feature is that the models for northern Chile show larger frictional heating
184 compared to those observed in south-central Chile^{8,9,14}. However, already ref.⁹ suggested that
185 frictional heating at the plate boundary increases northward, perhaps mimicking the increasing
186 age of the subducting plate. Furthermore, patterns are consistent with heat flow anomalies over

187 the marine forearc. Heat flow anomalies over the marine forearc are in the order of 50-60, 40-
188 50, and 24-31 mW/m² at 39°S, 36°S and 33°S, respectively¹⁴, decreasing northward and hence
189 reflecting increasing crustal age of the subducting plate and supporting a decrease of basal heat
190 flow. At 21°S, however, the age of the subducting plate has increased by roughly 20 Myr with
191 respect to 33°S, but the forearc heat flow is in the order of 30-40 mW/m²¹ and thus higher than
192 near 33°S, supporting higher values frictional heating than found further south. It might be
193 reasonable to hypothesize that sediment starved subduction erosion supports a higher degree of
194 friction than the accretionary margin of south-central Chile, but this interpretation is beyond
195 the scope of our work.

196

197 Both our analytical model and the numeric 2D model of ref.¹⁰ show higher temperatures
198 along the subduction megathrust fault with respect to other models for northern Chile. For
199 example, ref.¹⁸ (Suppl. Fig.8, magenta dashed line) did not consider any frictional heating with
200 the argument that in south-central Chile shear heating was low and therefore they obtained
201 lower temperatures.

202

203 **3. Supplementary Discussion**

204 Reflection energy absorption, seismic processing parameters, differences of gun energy
205 during seismic acquisition and shooting direction may all potentially cause a variation of
206 reflectivity along the plate boundary. To allow a spatial comparison of the reflectivity and avoid
207 issues caused by seismic processing, the same processing strategy is used for all profiles. In
208 addition, we discuss the following issues:

209

210 **A. Is the plate interface on seismic line MC04 visible to greater depth because of stronger** 211 **gun energy during data acquisition?**

212 The plate interface is traced to a depth of approximately 35 km on the northern dip line
213 MC04 but disappears at shallower depth on the dip lines that run through the rupture area of

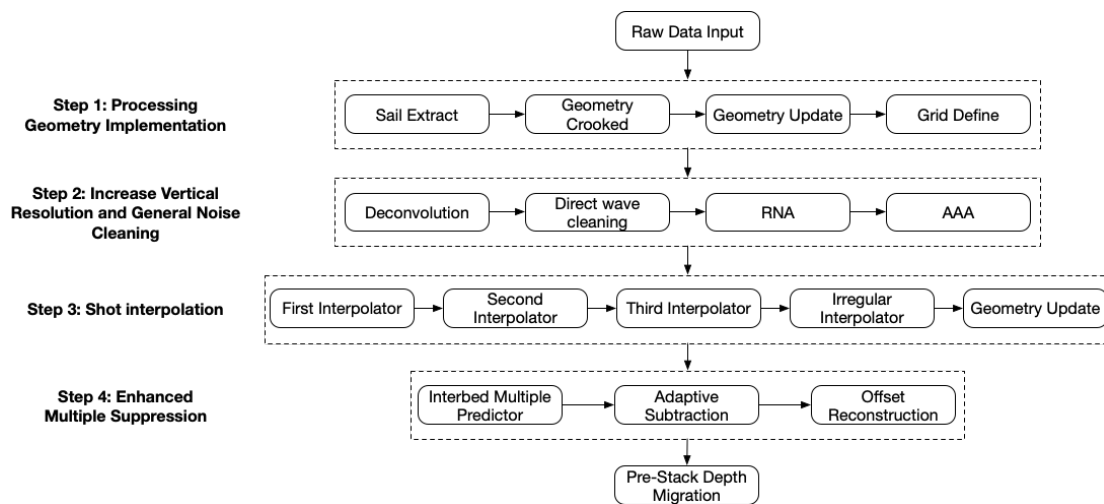
214 the 2014 Iquique earthquake. The acquisition geometry and gun array remained unchanged
 215 during the survey, which covered seismic lines parallel to the trench as well as profiles in the
 216 dip direction. Strike line MC30, which crosses the middle continental slope and is located
 217 furthest from the trench axis, documents that the plate interface along the northern part of the
 218 line can be traced to greater depth than along the southern part. This observation is augmented
 219 by all seismic dip lines and seismic strike lines of our survey.

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B. Is the difference in the plate interface reflection strength caused by shooting direction?

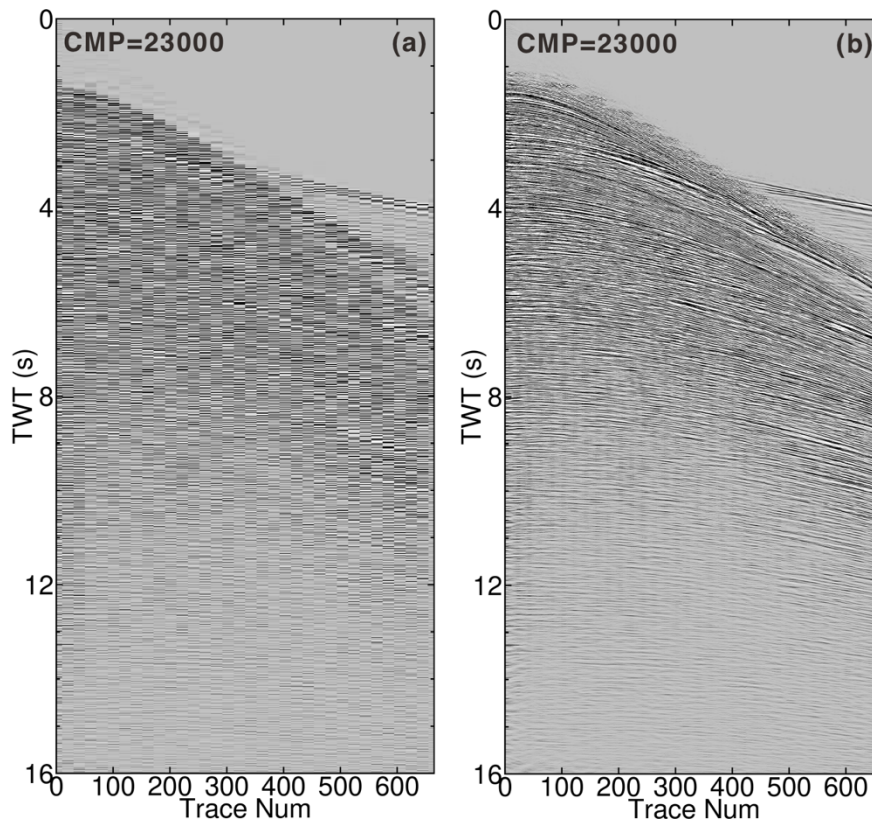
223 The shooting direction did not cause the observed differences of the plate interface
 224 reflection as seismic dip lines MC04, MC 06, and MC25 were shot from west to east, while
 225 seismic dip lines MC17 and MC23²² were shot from east to west. There are at least two seismic
 226 lines in the same direction, documenting that the shooting direction does not exert a major
 227 influence.

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Suppl. Fig. 1: Overview of the processing sequence.



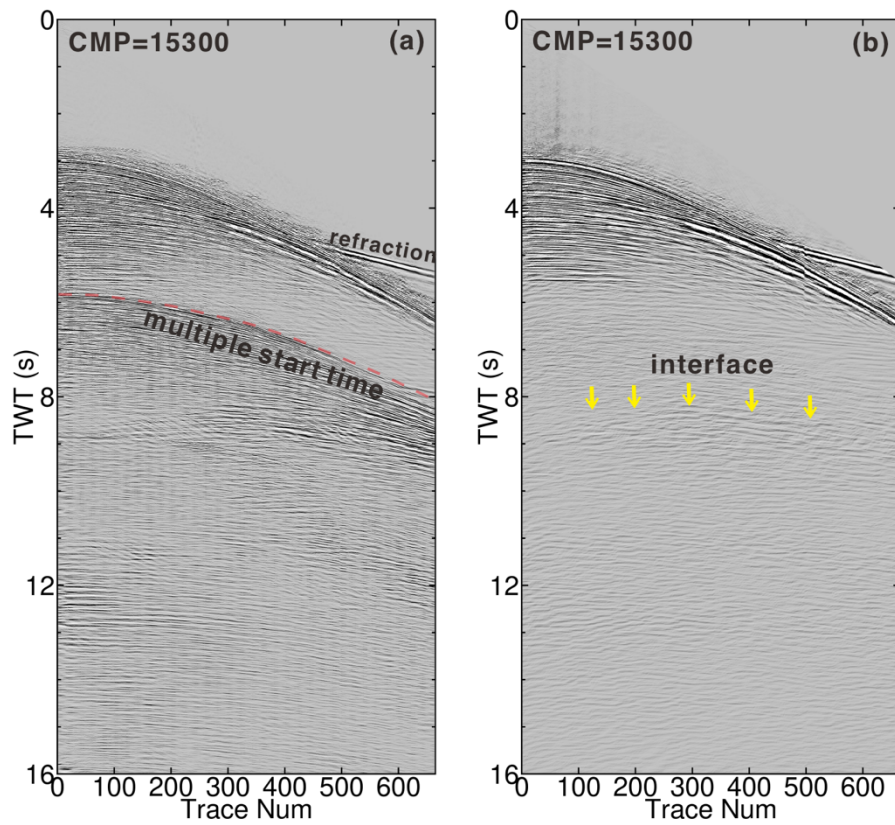
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237 **Suppl. Fig. 2:** The CMP gathers before and after interpolation of seismic line MC04. Before
238 the interpolation processing, the trace spacing is 250 m, which is shown in (a). After the
239 interpolation processing the trace spacing is 25 m in (b).

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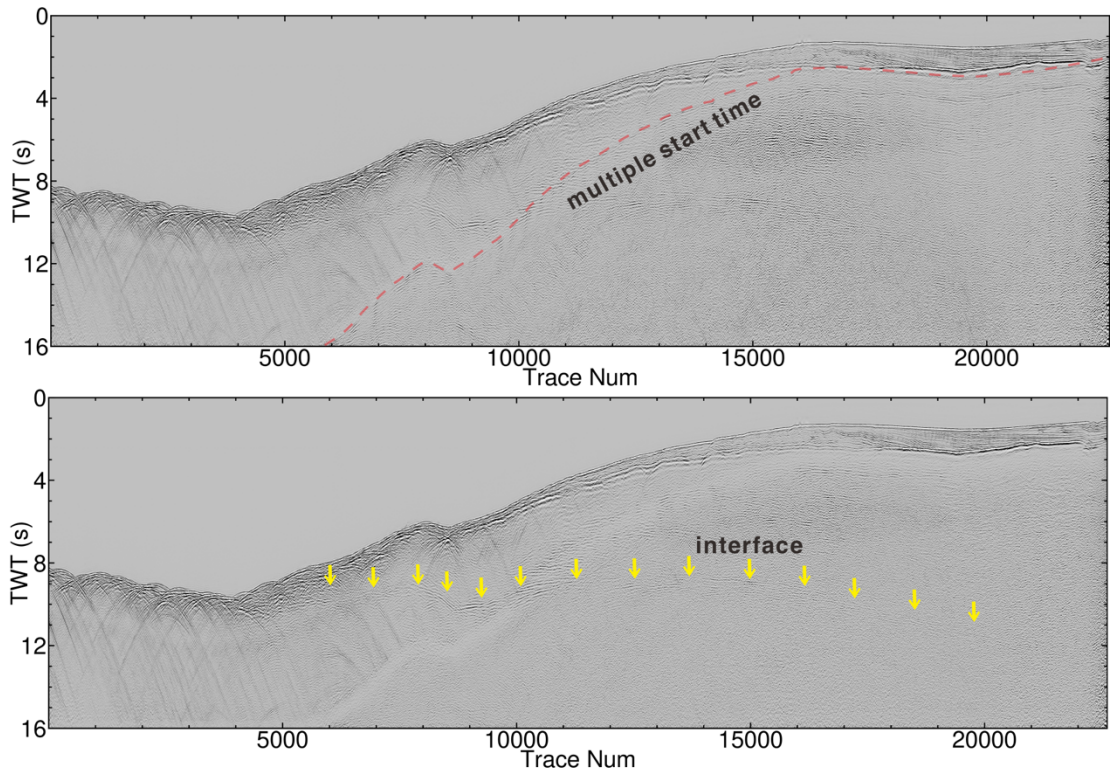
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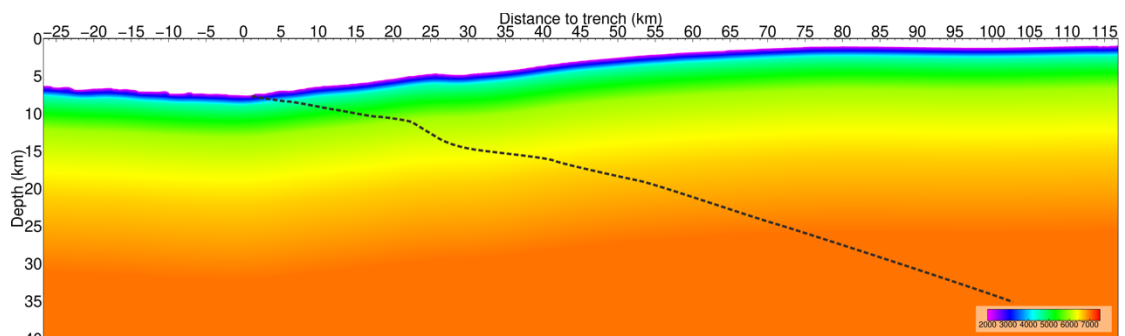
243 **Suppl. Fig. 3:** The CMP gathers before and after the de-multiple step of MC04. Before the
 244 demultiple processing, seismic signals are obscured by several multiple orders below the
 245 seafloor multiple start time indicated by the transparent red dashed line in (a). After the
 246 demultiple processing and high frequencies elimination, the multiples are mostly eliminated
 247 and we can see the interface traced by yellow arrows around 8-9 s TWT in (b).

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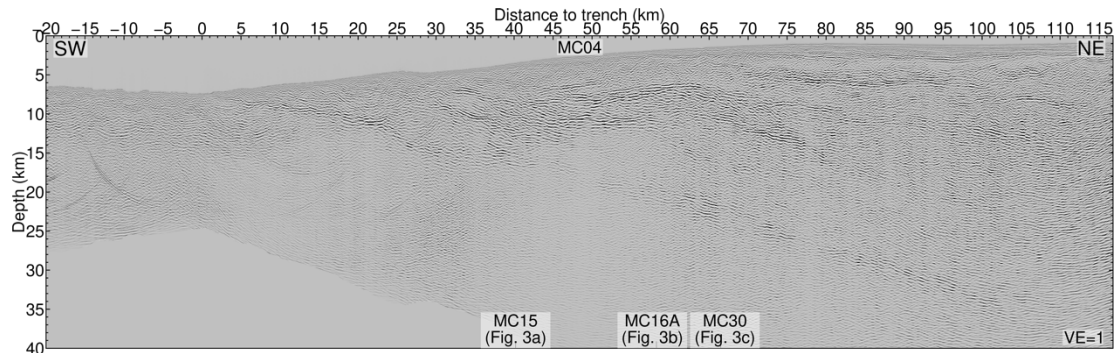
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 250 **Suppl. Fig. 4:** Stack section in the time domain of seismic line MC04 before and after the
 251 demultiple processing. Before the demultiple processing, the interface reflection is obscured by
 252 the multiples below the seafloor multiple start time indicated by the transparent red dashed line
 253 in (a). After the demultiple processing, the interface reflection can be traced around trace
 254 number 20000, as indicated by the yellow arrows in (b). Sections are shown after application
 255 of a depth customized gain of amplitudes.

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 261 **Suppl. Fig. 5:** An initial velocity model based on an unpublished 3D OBS velocity model (K.
 262 Davenport, personal communication) was applied for the pre-stack depth migration. The black
 263 dashed line indicates the plate boundary of profile MC04.

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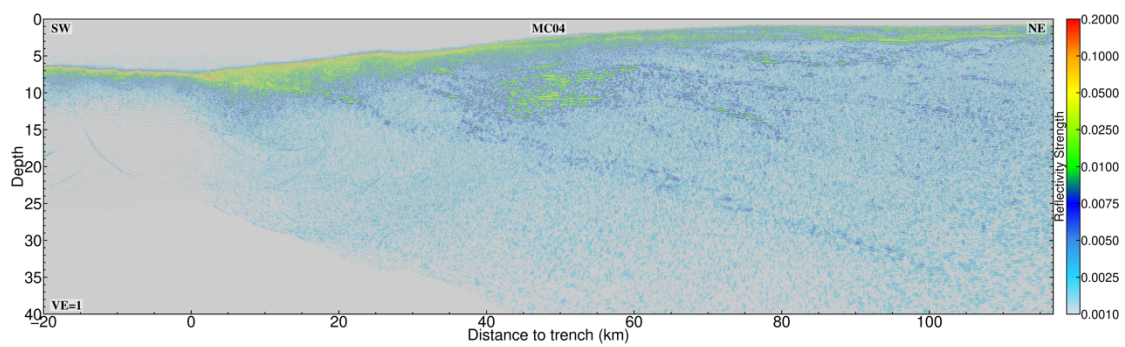
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Suppl. Fig. 6: The pre-stack depth migrated seismic image of seismic line MC04.

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Suppl. Fig. 7: Estimated reflection strength of pre-stack depth migrated seismic image of

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seismic line MC04. The plate boundary shows a unique reflection strength than the internal

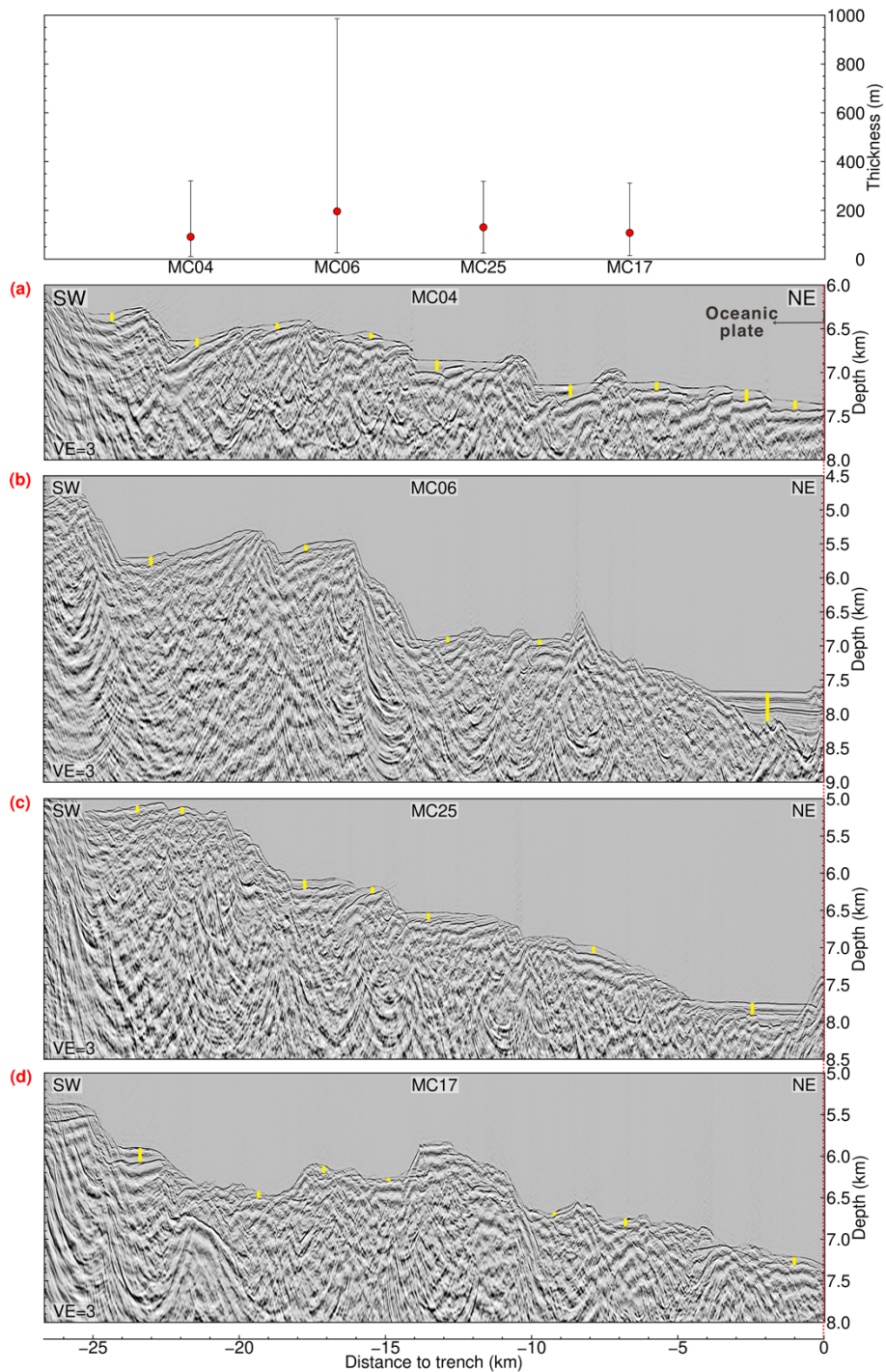
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crustal reflector elements. Vertical exaggeration is 1.

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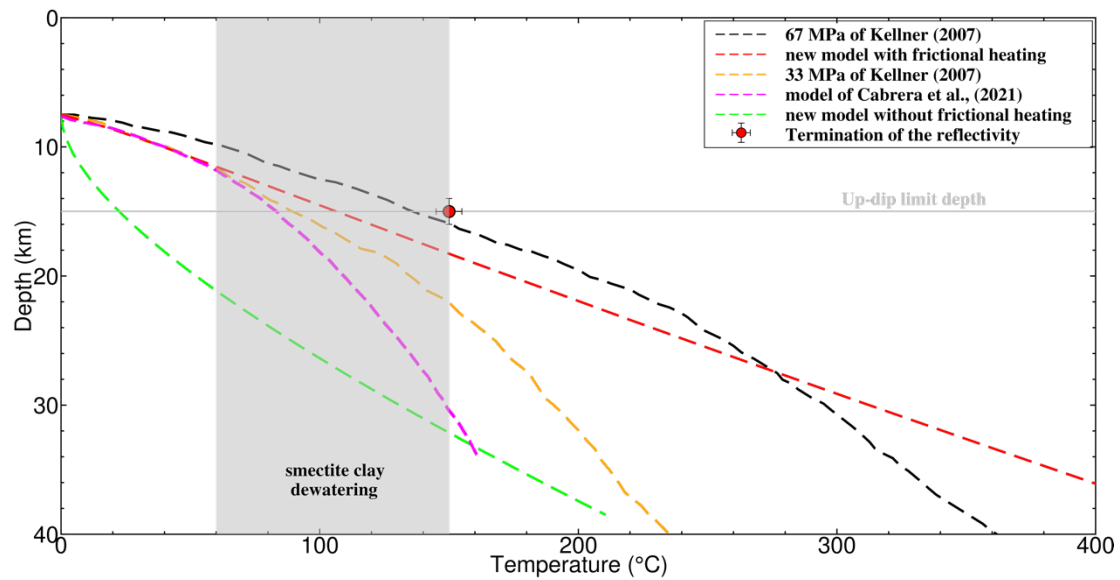


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278 **Suppl. Fig. 8:** Oceanic crust of pre-stack depth migrated section along seismic dip-lines. The
 279 error bar on top indicates the sediment upon the oceanic crust, in which the red dots show the
 280 average thickness of sediment on oceanic crust. The upper limit of the error bar represents the
 281 maximum thickness of sediments, while the lower limit indicates the minimum thickness of
 282 sediment. The maximum thickness anomaly along MC06 is due to more sediment in the trench
 283 than along the other seismic lines. In (a)-(d), yellow solid lines indicate the thickness of

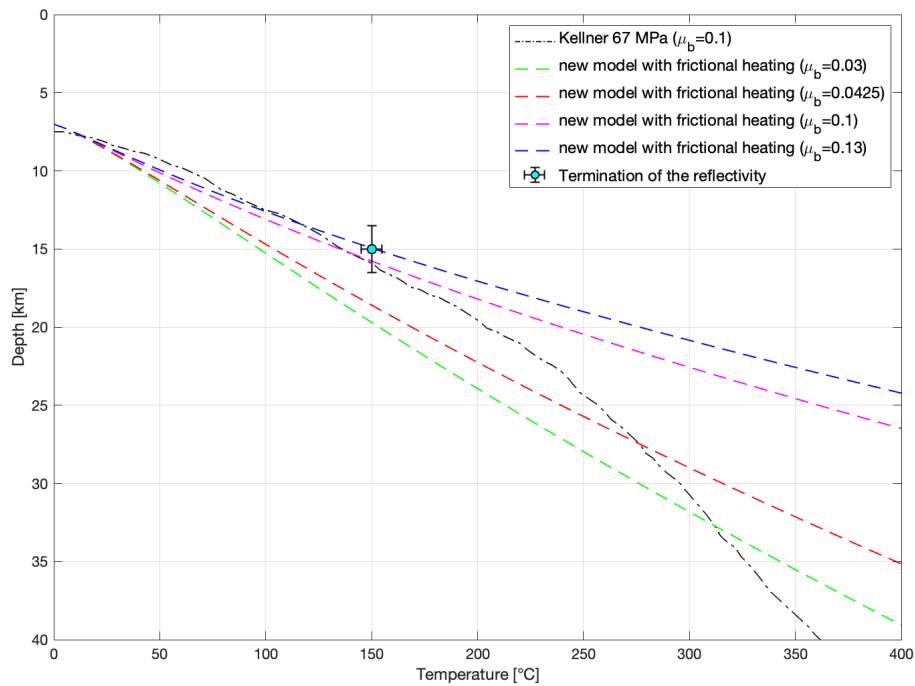
284 sediment on oceanic crust. Due to the bending of the oceanic crust, the sediments are
285 accumulated in half-graben structures. (a): seismic line MC04; (b): seismic line 06; (c): seismic
286 line 25; (d): seismic line 17. Vertical exaggeration is 3.

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289 **Suppl. Fig. 9:** Comparison of different thermal models. Black dashed line: thermal model with
290 shear stress $\tau=67$ MPa¹⁰; red dashed line: thermal model with frictional heating discussed in
291 main text; orange dashed line: thermal model with shear stress $\tau=33$ MPa¹⁰; magenta dashed
292 line: thermal model of ref.¹⁸; green dashed line: alternative thermal model computed without
293 frictional heating. The approximate depth range of smectite clay dehydration is based on refs.^{5,23-}
294 ²⁶.



295

296 **Suppl. Fig. 10:** New thermal model with varying effective coefficient of basal friction μ_b
 297 values. The range of μ_b is from 0.03-0.13, consistent with the global thermal measurement²⁰.
 298 The estimated average basal friction $\mu_b=0.1$ of Kellner 67 MPa is shown as a black dashed line.
 299 Using the same basal friction value, our new analytical model is shown as a magenta dashed
 300 line. In the main text, we used a friction coefficient of $\mu=0.85$ and pore fluid pressure $\lambda=0.95$ ¹³.
 301 The effective coefficient of basal friction μ_b depends on both the friction coefficient μ and pore
 302 fluid pressure λ along the fault zone: $\mu_b = \mu(1-\lambda)^{19}$. Based on this formula, the μ_b of the model
 303 in our main text is 0.0425, which is shown as red dashed line. $\mu_b=0.03$ and $\mu_b=0.13$ are shown
 304 as green and blue dashed lines, respectively.

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